

Dirac Particle on  $S^2$

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SUMMARY

The problem of a Dirac particle in stationary motion on  $S^2$  - a two dimensional sphere embedded in Euclidean space  $E^3$  - is discussed. It provides a particularly simple case of an exactly solvable constrained Dirac particle whose properties are here studied, with emphasis on its magnetic moment.

\* With financial support of FINEP and CNPq, Brazil.

## 1. INTRODUCTION

The quantum mechanical problem of a nonrelativistic particle constrained to move on a  $N$ -sphere  $S^N$  is well-known to admit exact solutions<sup>(1)</sup>. For the corresponding problem of a constrained Dirac particle on  $S^N$ , solutions, to our knowledge are not known for a general value of  $N$ , although the particular case of  $N=2$  (the two-dimensional sphere  $S^2$ , embedded in Euclidean space  $E^3$ ), as we shall see, admits exact solutions. This is a direct consequence of the embedding property of  $S^2$  in  $E^3$ , which allow us to derive stationary solutions on  $S^2$  as a restriction ( $r=R$ ) imposed on the free motion parametrized by spherical coordinates  $(r, \theta, \varphi)$  in  $E^3$ ,  $R$  being the radius of the 2-sphere.

After discussing the energy spectrum of a Dirac particle on  $S^2$  in section 2, we carry out, in section 3, perturbative calculations of the magnetic moment of the system. Section 4 is devoted to some final remarks on the problem.

## 2. THE ENERGY SPECTRUM OF THE FREE DIRAC PARTICLE ON $S^2$ .

Our starting point are the stationary solutions of a free Dirac particle in  $E^3$ , in spherical coordinates. Since this is a rather well-known subject we go directly to the main results to be used in the sequel.

The Dirac equation for a free particle of mass  $M$  ( $\hbar=c=1$ ) is

$$\left( \underline{\alpha} \cdot \underline{p} + \beta M - E \right) \psi = 0 . \quad (2.1)$$

Spherical solutions of (2.1) may be written in the form<sup>(2)</sup>

$$\psi(r, \theta, \varphi) = \begin{pmatrix} g(r) y_{j_1 l_1}^{j_1}(\theta, \varphi) \\ i f(r) y_{j_2 l_2}^{j_2}(\theta, \varphi) \end{pmatrix} \quad (2.2)$$

where  $g(r)$  and  $f(r)$  are real functions of the Euclidean distance  $r$  in  $E^3$  and the  $y_{j l}^{j_1}$ 's are spinor harmonics, responsible for the angular dependence of  $\psi$ .

The two radial functions obey the equations

$$\begin{cases} \frac{d}{dr} f(r) + (1-k) \frac{f(r)}{r} + (E-M)g(r) = 0 \\ \frac{d}{dr} g(r) + (1-k) \frac{g(r)}{r} - (E+M)f(r) = 0 \end{cases} \quad (2.3)$$

in which  $k$  is the eigenvalue of the Dirac  $K$  operator,  $K = \beta(\tilde{\Sigma} \cdot \tilde{L} + 1)$ .

Constraining the particle to move on  $S^2$ , of radius  $R$ , we

put

$$\begin{aligned} g(r) &= g(R) = a, & \frac{dg}{dr} &= 0 \\ f(r) &= f(R) = b, & \frac{df}{dr} &= 0. \end{aligned} \quad (2.4)$$

Consequently, the two real constants  $a$  and  $b$  will obey the linear system

$$\begin{cases} (E-M)Ra - (k-1)b = 0 \\ (k+1)a - (E+M)Rb = 0. \end{cases} \quad (2.5)$$

The normalization of the wave-function on  $S^2$  is defined by

$$\int \psi^+ \psi dS = 1, \quad (2.6)$$

with  $dS = R^2 d\Omega$ .

Condition (2.6) gives immediately

$$R^2(a^2 + b^2) = 1. \quad (2.7)$$

The solutions of the system (2.5), with condition (2.7) are:

(i) if  $k = +1$  :

$$a = \frac{M}{(1+M^2R^2)^{1/2}}, \quad b = \frac{1}{R(1+M^2R^2)^{1/2}} \quad \text{for } E=M, (2.8_1)$$

$$a = 0, \quad b = \frac{1}{R} \quad \text{for } E=-M;$$

(ii) if  $k = -1$  : (2.8\_2)

$$a = \frac{1}{R}, \quad b = 0 \quad \text{for } E=M,$$

$$a = \frac{1}{R(1+M^2R^2)^{1/2}}, \quad b = \frac{M}{(1+M^2R^2)^{1/2}} \quad \text{for } E=-M;$$

(iii) if  $k \neq \pm 1$ ,

$$a = \frac{1}{R} \left[ 1 + \frac{(E-M)^2 R^2}{(k-1)^2} \right]^{-\frac{1}{2}}, \quad b = \frac{1}{R} \left[ 1 + \frac{(E+M)^2 R^2}{(k+1)^2} \right]^{-\frac{1}{2}}$$

with 
$$E = \pm \frac{1}{R} \left[ M^2 R^2 + (k^2 - 1) \right]^{\frac{1}{2}} \quad (2.8_3)$$

$$(k = \pm 1, \pm 2, \dots).$$

The energy levels are seen to have a double degeneracy with respect to  $k$ . Since there is still a  $(2j + 1)$  - fold degeneracy due to  $j_z$ , for the total degeneracy  $d$  we, of course have

$$d = 2(2j+1) = 4|k|. \quad (2.9)$$

The parity  $\pi$  of a state of given  $k$  is

$$\pi = \begin{cases} (-1)^{|k|} & \text{if } k > 0 \\ (-1)^{|k|-1} & \text{if } k < 0. \end{cases}$$

Therefore, the parities of the states  $k$  and  $-k$  are opposite. It follows that the double degeneracy corresponds to a parity doubling of the spectrum.

### 3. THE MAGNETIC MOMENT

We will now discuss the magnetic moment of a Dirac particle, of charge  $Q$ , constrained to move on  $S^2$ .

If we assume that a weak, constant and uniform magnetic field, taken in the  $Oz$  direction, is acting on our system, the magnetic moment  $\mu$  will be given by<sup>(3)</sup>

$$\begin{aligned} \mu &= \frac{Q}{2} \int \psi^\dagger (\underline{r} \times \underline{\alpha})_3 \psi dS \\ &= \frac{QR^3}{2} \int \psi^\dagger (\hat{r} \times \underline{\alpha})_3 \psi d\Omega, \end{aligned} \quad (3.1)$$

where  $\hat{r} = \frac{\underline{r}}{R}$  and  $\psi$  is taken for  $j_3 = j$ :

$$\psi = \begin{pmatrix} a y_{j l_A}^j \\ i b y_{j l_B}^j \end{pmatrix} = \begin{pmatrix} a y_{j l_A}^j \\ -i b \underline{\sigma} \cdot \hat{r} y_{j l_A}^j \end{pmatrix}. \quad (3.2)$$

The last step in eq. (3.2) is a consequence of the equality

$$\underline{\sigma} \cdot \hat{r} y_{j l_A}^j = - y_{j l_B}^j, \quad (3.3)$$

where  $\underline{\sigma}$  are  $2 \times 2$  Pauli spin matrices.

By a simple manipulation, we have

$$\mu = \frac{QR^2}{2} \int \psi^+ \psi d\Omega \quad (3.4)$$

where

$$\psi^+ \psi = 2Rab \begin{pmatrix} Y_{j l_A}^{j+} \\ Y_{j l_A}^{j-} \end{pmatrix} \begin{pmatrix} -\sin^2 \theta & \sin \theta \cos \theta e^{-i\varphi} \\ \sin \theta \cos \theta e^{+i\varphi} & \sin^2 \theta \end{pmatrix} \begin{pmatrix} Y_{j l_A}^j \\ Y_{j l_A}^j \end{pmatrix} \quad (3.5)$$

We shall consider separately two cases, according to whether  $k = \pm(j + \frac{1}{2})$ .

(i)  $k = -(j + \frac{1}{2})$ , hence  $l_A = j - \frac{1}{2}$ .

In this case, we get

$$\psi^+ \psi = -2Rab \left| Y_{j-\frac{1}{2}}^{j-\frac{1}{2}}(\theta, \varphi) \right|^2 \sin^2 \theta. \quad (3.6)$$

With the help of the integral

$$\int \left| Y_{j-\frac{1}{2}}^{j-\frac{1}{2}} \right|^2 \sin^2 \theta d\Omega = \frac{l_A + 1}{l_A + \frac{3}{2}}, \quad (3.7)$$

we obtain

$$\begin{aligned} \mu &= -QR^3ab \frac{l_A + 1}{l_A + \frac{3}{2}} \\ &= -QR^3ab \frac{k}{k - \frac{1}{2}}. \end{aligned} \quad (3.8)$$

(ii)  $k = j + \frac{1}{2}$ , hence  $l_A = j + \frac{1}{2}$ .

We get

$$\begin{aligned} \psi + \theta \psi &= -\frac{2Rab \sin \theta}{2j+2} \left[ Y_{j+\frac{1}{2}}^{j-\frac{1}{2}*} Y_{j+\frac{1}{2}}^{j-\frac{1}{2}} \sin \theta + \right. \\ &+ (2j+1)^{1/2} Y_{j+\frac{1}{2}}^{j-\frac{1}{2}*} Y_{j+\frac{1}{2}}^{j+\frac{1}{2}} \cos \theta e^{-i\varphi} + \\ &+ (2j+1)^{1/2} Y_{j+\frac{1}{2}}^{j+\frac{1}{2}*} Y_{j+\frac{1}{2}}^{j-\frac{1}{2}} \cos \theta e^{i\varphi} + \\ &\left. - (2j+1) Y_{j+\frac{1}{2}}^{j+\frac{1}{2}*} Y_{j+\frac{1}{2}}^{j+\frac{1}{2}} \sin \theta \right]. \end{aligned} \quad (3.9)$$

Now, making use of the following expressions for the spherical harmonics involved in (3.9), namely

$$Y_l^l = \left[ \frac{(2l+1)!! 2^l}{(2l)!! 4\pi} \right]^{\frac{1}{2}} \hat{x}_+^l,$$

$$Y_l^{l-1} = \left[ \frac{(2l+1)!! 2^{l-1}}{(2l-2)!! 4\pi} \right]^{\frac{1}{2}} \hat{x}_0 \hat{x}_+^{l-1}$$

$$\left( \hat{x}_+ = \frac{-1}{\sqrt{2}} (\hat{x} + i\hat{y}), \quad \hat{x}_0 = \frac{z}{R} \right),$$

we calculate the four terms appearing in that equation and obtain for  $\mu$  the result

$$\mu = \frac{1}{2} \frac{QR^3 ab}{2j+2} \frac{(2l_A+1)!!}{(2l_A-1)!!} \int_0^\pi \sin^{2l_A+1} \theta d\theta$$

$$= + QR^3 ab \frac{l_A}{l_A + \frac{1}{2}}$$

$$= + QR^3 ab \frac{k}{k + \frac{1}{2}}.$$

(3.10)

Therefore, the two cases (i) and (ii) can be written together as

$$\begin{aligned} \mu_{\pm} &= \pm QR^3 ab \frac{k}{k \pm \frac{1}{2}} \\ &= \pm QR^3 ab \frac{j + \frac{1}{2}}{j + 1} \end{aligned} \quad (3.11)$$

corresponding to  $k = \pm (j + \frac{1}{2})$  respectively.

In eq. (3.11), the factor  $R^3 ab$  may be more conveniently calculated as follows. By multiplying the first of eq. (2.5) by  $b$ , the second by  $a$  and adding, we have get

$$\begin{aligned} R^3 ab &= \frac{1}{2M} (1 + k - 2kR^2 b^2) \\ &= \frac{1}{2M} \left[ 1 + k - \frac{2k}{1 + \frac{(E+M)^2 R^2}{(k+1)^2}} \right], \end{aligned} \quad (3.12)$$

where in the last step we made use of eq. (2.8). By applying (3.12) to (3.11), we obtain

$$\mu_{\pm} = \pm \frac{Q}{2M} \left[ 1 + k - \frac{2k}{1 + \frac{(E+M)^2 R^2}{(k+1)^2}} \right] \frac{k}{k \pm \frac{1}{2}}. \quad (3.13)$$

Finally, it may be of interest to point out that for the magnetic moment calculated via the iterated Dirac equation, as the expectation value of

$$\mu = \frac{Q}{2E} (L_3 + \Sigma_3), \quad (3.14)$$

we obtained

$$\mu_{\pm} = \pm \frac{Q}{2E} k \begin{cases} 1 - \frac{R^2 a^2}{k + \frac{1}{2}} & \text{if } k > 0, \\ 1 + \frac{R^2 b^2}{k - \frac{1}{2}} & \text{if } k < 0. \end{cases} \quad (3.15)$$

where the expressions for  $a$  and  $b$  are given by eq. (2.8<sub>3</sub>).

#### 4. FINAL REMARKS

The solutions obtained are remarkably simple and are given in exact form for any value of  $k$  (or  $j$ ). The energy spectrum as given by formula (2.8<sub>3</sub>) is a discrete one and corresponds to parabolic Regge trajectories ( $E^2 \propto j^2$ ). It also presents a double degeneracy in  $k$ , which may be interpreted as a parity doubling.

We remark that this degeneracy can be easily broken up - and exact solutions still obtained - if we introduce in the Dirac Hamiltonian an additional term of the form  $F(K, J^2)$ ,  $f$  denoting an arbitrary function of the commuting angular operators  $K$  and  $J^2$ .

Finally, a few words about the magnetic moment calculations are in order. It should be noted that magnetic moments calculated via (3.1) or via (3.14) give, in general, different results<sup>(4)</sup>, the differences being of the order of 25% .

An exceptional case is that of the ground state ( $E=0$ ) with  $k=-1$ . In this case, eq.(3.12) gives  $\mu_z=0$  whereas, from eq. (3.15) the result  $\frac{Q}{2M}$  is obtained. This may be easily understood having in mind that in this state  $b=0$  and remarking that in the first case the magnetic moment is given by the product  $a b$  whereas in the second, it is given by a sum of two terms, one proportional to  $a^2$ , the other to  $b^2$ .

Let us add a few words about the future outlook. This work can proceed in at least two directions: in the first, a generalization to  $N$ -spheres is envisaged. In the second, instead of spheres, the Dirac particle is constrained to move on a surface of different symmetry as, for instance, a circular cylinder of given length, a case also amenable to an exact treatment. We hope to return to these interesting topics at a later date.

#### ACKNOWLEDGMENT

One of us (P.L.F.) is grateful to Prof. E. Ley Kroo, of UNAM, México for interesting discussions.

REFERENCES AND NOTES

- (1) For a recent discussion, see P.Leal Ferreira, Rev.Brasil. Fís., Special volume, p.183, July (1984).
- (2) For the Dirac equation and related quantities we follow the notation and conventions of J.J.Sakurai, "Advanced Quantum Mechanics", Addison -Wesley Publ. Co., Reading, MA (1967).
- (3) This definitions is consistent with the secular equation resulting from the double degeneracy since the off-diagonal elements of the perturbation vanish in first order.
- (4) This fact was recognized long ago by H.Margenau, Phys. Rev. 57, 383 (1940).