

MASTER

Current Schemes for National Synchrotron Light Source UV Beamlines*

Gwyn P. Williams, Malcolm R. Howells and Wayne R. McKinney

National Synchrotron Light Source
Brookhaven National Laboratory, Upton, New York 11973

BNL--26287

DE83 013292

ABSTRACT

We describe in some detail four beamlines proposed for the National Synchrotron Light Source UV ring at Brookhaven National Laboratory. Three grazing-incidence instruments, one of the plane grating Mijake type and two with toroidal gratings at grazing angles of $2\frac{1}{2}^\circ$ and 15° are described. Two normal incidence instruments, one using the source as entrance slit and accepting 75 milliradians horizontally are also discussed. In each case we have estimated the output fluxes expected from such beamlines.

DISCLAIMER

This report was prepared as an account of work sponsored by an agency of the United States Government. Neither the United States Government nor any agency thereof, nor any of their employees, makes any warranty, express or implied, or assumes any legal liability or responsibility for the accuracy, completeness, or usefulness of any information, apparatus, product, or process disclosed, or represents that its use would not infringe privately owned rights. Reference herein to any specific commercial product, process, or service by trade name, trademark, manufacturer, or otherwise does not necessarily constitute or imply its endorsement, recommendation, or favoring by the United States Government or any agency thereof. The views and opinions of authors expressed herein do not necessarily state or reflect those of the United States Government or any agency thereof.

NOTICE

PORTIONS OF THIS REPORT ARE ILLEGIBLE.

It has been reproduced from the best available copy to permit the broadest possible availability.

* Research supported by the U.S. Department of Energy.

1. Introduction

This paper describes four beamlines devised for installation on the National Synchrotron Light Source (NSLS) UV ring. Three of the four lines (U2, U3, and U4) use normal incidence monochromators, U4 having access to the synchrotron radiation through a LiF window. Beamline U1 incorporates three grazing incidence instruments, one using a plane grating and two using toroidal gratings. The NSLS UV ring which is designed to run at 0.7 GeV with 1 amp. stored current has a bending radius of 1.91 meters giving it a critical wavelength of about 32\AA . Up to 90 milliradians (horizontal) x 10 milliradians (vertical) of radiation opening angle are available from each of the 16 beamlines. Since the light source has such a small vertical beam size (FWHM $\sim .2\text{mm}$) and high brightness, it is often attractive to use the source itself as a monochromator entrance slit. The savings in beamline hardware and reflection losses are often very considerable. Three of the four beamlines described employ this approach. The high brightness remains an advantage even if a collection mirror is used, however, helping to minimize the losses at entrance slits for instruments of high resolution.

2. Grazing Incidence Beamline U1

This line is divided into three, one of which (20 milliradians) is spare. It is shown schematically in Fig. 1. U1A intercepts 5 mrad horizontally with an 86° angle of incidence pre-mirror at 10 meters from the source. This mirror is made up of four segments of spherical figure and focuses mainly in the horizontal plane. Detailed ray tracing indicates that such a mirror will give adequate image quality and large cost saving compared to an aspheric mirror some .75 meters long. The horizontal focus is arranged to occur at the sample position i.e. 150mm after the monochromator exit slit and is $< 1\text{mm}$ wide. 2.5 meters from the pre-mirror the radiation is dispersed downwards in the vertical plane by a plane grating^(1,2). The dispersed beam is then collected and focussed on the (horizontal) exit slit by a parabolic mirror.

The gratings have a rectangular profile^(2,3). These so-called "laminar" gratings have a "wavelength of maximum efficiency" (λ_m) very like a blaze. λ_m can be chosen at will by correct choice of the groove depth and angle of incidence for a given groove density. Once the groove depth has been chosen λ_m becomes a function of α . It turns out by good fortune that in negative order with a gold coated 600 lines/mm grating, λ_m is roughly the same function of α as λ_j , where λ_j is the typical wavelength covered by range j over which reflection filtering of second and higher orders occur, and which extends from $\lambda_j/2$ to $2\lambda_j$. The instrument is designed to work in several "ranges" and the above discovery means that one grating can cover several such ranges and never be far off blaze. Within a given range the post-mirror remains stationary and the grating is rotated with a sine bar arrangement such that the linear position of a drive rod is directly proportional to wavelength.

Six ranges are envisaged for the instrument and their properties are listed in table 1. The best resolution with the source as entrance slit is .013Å. Estimated fluxes for the various ranges are presented in Fig. 4.

Beamline UIB (Fig. 1) houses two toroidal grating monochromators (TGM's) one vertically above the other and sharing a common target point. The upper TGM has an 87.5° angle of incidence and collects 3 milliradians horizontally. The source is used as entrance slit and the grating is the only beamline optical component. The resolution is intended to be 2eV. The lower TGM has unit magnification pre and post mirrors and collects 15 milliradians horizontally. Two gratings are used to give a wavelength range from 80-320Å with .25Å resolution and from 320 - 1680Å with ~ 1Å resolution. Estimates of the output fluxes from these instruments are shown in Figs. 5 and 6.

Depending on the requirements of a particular experiment it is possible that different TGM configurations will be required. For example for some experiments such as photoemission it may be desirable to design the instrument for constant

energy, rather than wavelength resolution. Fig. 6 compares fluxes from two TGM's, one designed for constant energy and one designed for constant wavelength resolution. It turns out that if the grating from the $\Delta\lambda$ instrument were masked down to achieve $\Delta E = 0.1\text{eV}$, the output flux would be much lower than the ΔE instrument⁽⁴⁾. Clearly any choice of instrument must involve considerations of flux and resolution.

The general philosophy regarding TGM's is to design these instruments in such a way that slit to grating distances and angles of incidence may fairly readily be changed to accommodate different optimization schemes. The designs presented here should therefore be regarded as illustrative examples.

3. Normal Incidence Beamlines U2 and U3

The two beamlines U2 and U3 are identical. Each has a 3.5 meter normal incidence monochromator which uses the source as an entrance slit⁽⁵⁾. The beam is collected by three plane pre-mirrors which superpose their respective sources to a common source point. Since the total number of milliradians collected is 75, the tri-mirror arrangement is necessary to reduce the effective source size and maintain resolution. As shown in Fig. 2, the pre-mirrors also serve to direct the emerging beam away from the storage ring. With a .2mm source and a 3600 $\text{\AA}/\text{mm}$ grating, the resolution is $\sim .2\text{\AA}$. Some flux estimates are shown in Fig. 7.

4. LiF Windowed Beamline U4

This line is divided into two, each of which has a commercial Czerny-Turner 0.5 meter monochromator installed (Fig. 3). In each case the radiation is intercepted by a plane mirror which collects some 40 milliradians. This light is focussed through a LiF window onto the entrance slit of the monochromator with a magnification of 0.5. Beamline U4A has the monochromator installed with vertical dispersion, while U4B has horizontal dispersion⁽⁶⁾, the choices being made solely on the basis of convenience for the experimental chambers. Estimated fluxes are presented in Fig. 8.

References

- (1) K. P. Mijake, R. Kato and H. Yamashita, *Sci. Light* 18, 1, 39 (1969).
- (2) M. R. Howells, D. Norman, G. P. Williams and J. B. West, *J. Phys. E* 11 199 (1978).
- (3) A. Franks, K. Lindsey, J. M. Bennett, R. J. Speer, D. Turner and D. J. Hunt, *Philos. Trans.* A277 503 (1978).
- (4) W. R. McKinney and M. R. Howells, *This Conference Proceedings*.
- (5) E. L. Ederer, B. Cole and J. B. West, *This Conference Proceedings*.
- (6) J. C. Sutherland, E. J. Desmond and P. Z. Takacs, *This Conference Proceedings*.

Table 1. NSIS UV beamline Schemes and Parameters.

	W1(A) ^a	W1(B)	W1(C)	W2	W3	W4(A)	W4(B)
1. # of horizontal mirrors	HORIZ 5	30	20	75	75	40	40
	VERT 51.4	10	10	10	10	10	10
2. Type of experiment	High res. solid state	AMES, SEXAFS ^b	SPARE mirrors	None at present	Not beam chemistry	Time resolved fluor	Biology ¹⁴
3. Type of monochromator(s)	Mijaka	Double TGM ^c	AT PRESENT	Normal incidence	Normal incidence	0.5m Czerny-Turner ¹⁵	0.5m Czerny-Turner
4. Wavelength range(s) Å	R1:30-66 R2:60-110 R3:90-240 R3:150-475 R5:400-1200 R1(A):25-50	M1:0-100 M2:80-1600 ^d G1:80-320 G2:320-1600		300-3000	300-3000	1050-12000	1050-12000
5. Limitation on resolution	Source size	M1 mostly source M2 Aberr		Source size ^e	Source size ^e	Aberrations	Aberrations
6. Best resolution (Å)	R1:0.016 R2:0.022 R3:0.039 R4:0.062 R5:0.1 R(A):0.013	M1:2eV M2:61 ~ .25 G2 ~ 1.0		0.25 ¹⁶	0.25 ¹⁶	0.3	0.3
7. Special features		2 Mon'rs feed 1. sample					
8. Plane of dispersion	Vertical	Vertical		Vertical ^g	Vertical ^g	Window, no fast valve	Window, no fast valve
9. Dispersion (Å/mm)	^h R1:1.20 R2:1.90 R3:3.4 R4:5.7 R5:9.1 R(A):0.91	G1 ~ .55 G2 ~ 2.2		1.19	1.19	Vertical	Horizontal
10. Angular acceptance(mr)(H×V)	Noncollinearly parallel light	M1:3×.25 M2:15×10		(75×10)	(75×10)	(115×115)	(115×115)
11. Entrance slit	None	M1:No M2:Yes		None	None	Yes	Yes
12. Blaze wavelength (Å)	R1:45 R2:60 R3:135 R4:225 R5:600 R(A):35	M1:12 M2:G1:120 G2:480		500 (say)	1000 (say)	1500 (say)	1500 (say)
13. Grating type	Plane, lamina ⁱ	Toroidal, lamina ⁱ		Spherical ¹¹	Spherical ¹¹	Plane ¹³	Plane
14. Groove density gr/mm	600 and/or 1200	M1:600 G1:1800 G2:450		2400	2400	1200	1200
15. Ruled area (mm×mm)	50 × 19	M1:27 × 14 M2:75 × 25		264 × 72	264 × 72	52 × 52	52 × 52
16. Grating/slit distance(mm) ENT		M1:4490 M2:1151					
17. Grating/slit distance(mm) EXII		M1:4336 M2:1930		3500	3500		
18. Scanning method	Simple rotation	Simple rotation		Rot & translation	Rot & translation	Simple rotation	Simple rotation
19. Admittance at best resn. (mradian.m)(Horiz×vert)	(.25) × (.50-.28)						
COLLECTION MIRROR DETAILS							
20. Size (mm×mm)	800 × 35	M1:None M2:66 × 106		248 × 32 ¹²	248 × 32 ¹²	Two A & B ¹⁴	Two A & B ¹⁴
21. Shape [radius(m)]	Spherical (80)	Ellipsoidal		Plane	Plane	A Plane B Sphere	A Plane B Sphere
22. Angle of incidence	86°	M1:87.5° M2:75°		7.5°	7.5°	A 5° B 5°	A 5° B 5°
23. Magnification	0.25	1.0		0.5	0.5	A - B 0.5	A - B 0.5
24. Focal length(m)	2.8					A - B 0.77	A - B 0.77

Footnotes

¹Using 1200 gr/mm grating.

²Using 600 gr/mm grating which gives best order sorting

³AMES = Angle resolved photoelectron spectroscopy

⁴SEXAFS = Surface extended x-ray absorption fine structure

⁵Called M1: 175° included angle and M2: 150° included angle

M1 has one grating M2 has two called G1 and G2

Note that NSIS TGM's will be configured so that the slit distances and included angle can be chosen at will. This allows a choice of the wavelength range of best resolution and of the low wavelength threshold.

To change from one configuration to another would involve a resurvey.

⁶This is achieved with two gratings G1 and G2

⁷Without aberration correction

⁸See Brookhaven Informal report #26027

⁹A slit can be placed with some difficulty near the source so that better resolution can be achieved.

¹⁰Has horizontal beam coming away from the storage ring.

¹¹With 2400 gr/mm grating

¹²Special shape to 'match' shape of SR beam. Rollings are parallel to longest direction.

¹³In three segments each at the same distance from their own sources.

¹⁴Al/MgF₂ coating. Should be 80% reflection efficient at Lyman α

¹⁵Acton Research instrument is suggested.

¹⁶A double reflection is used, first a plane mirror then a spherical mirror both at near normal incidence. The latter images the source on to the monochromator entrance slit.

¹⁷Existing monochromator and experiment.

Figure Captions

1. Schematic layout of NSLS beamline U1.
2. Schematic layout of NSLS beamline U2.
3. Schematic layout of NSLS beamline U4.
4. Estimated output flux from the Mijake type plane grating monochromator in beamline U1. The numbers refer to the ranges defined in table I.
5. Output flux estimated from the 2.5° TGM in beamline U1.
6. Output flux estimated from 15° TGM's in beamline U1. Solid line - instrument designed for constant energy resolution. Dashed line - instrument designed for constant wavelength resolution.
7. Estimated output flux from beamlines U2 and U3 for two gratings.
8. Estimated output flux for three gratings for beamline U4.

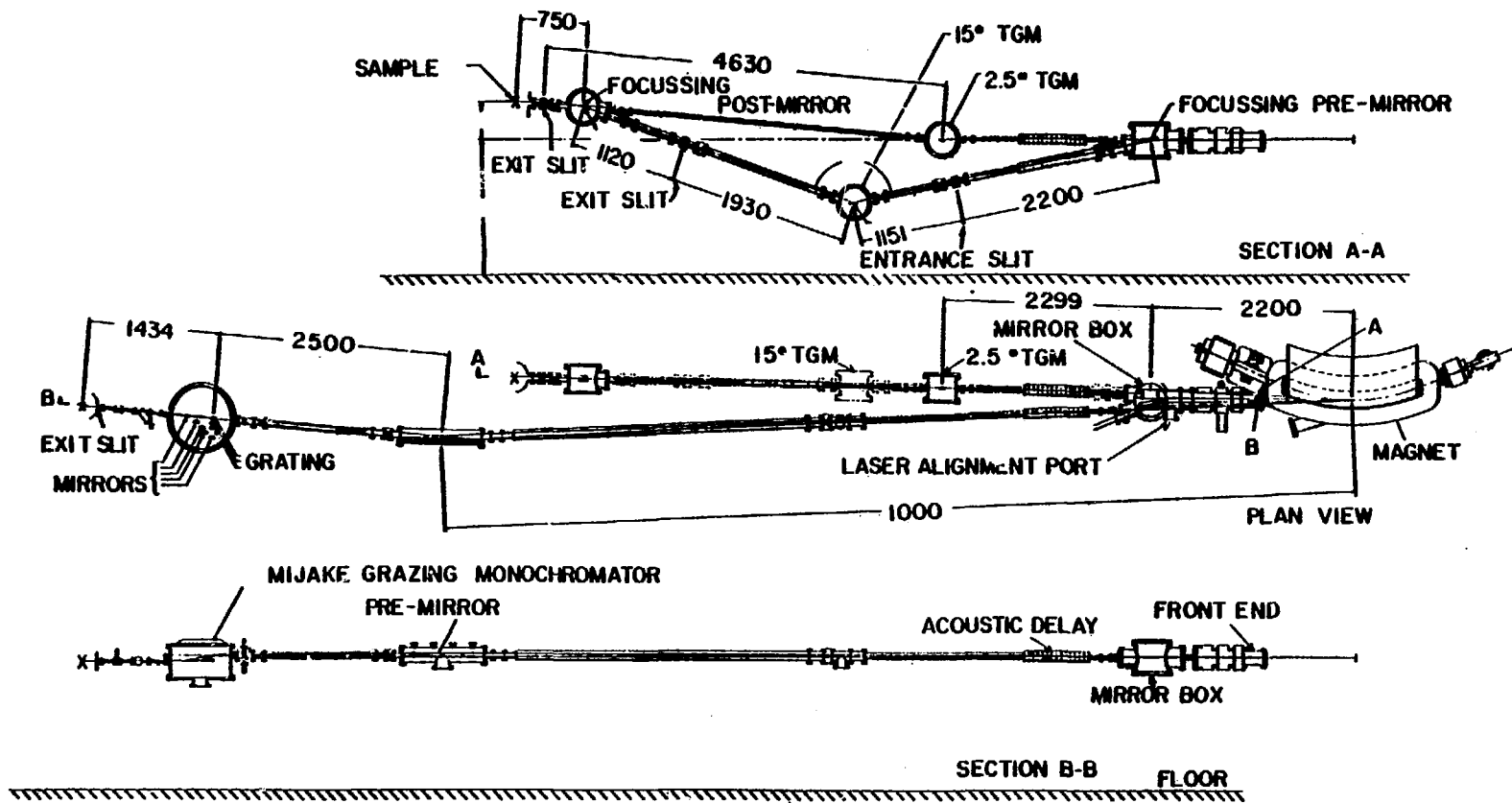


Figure 1.

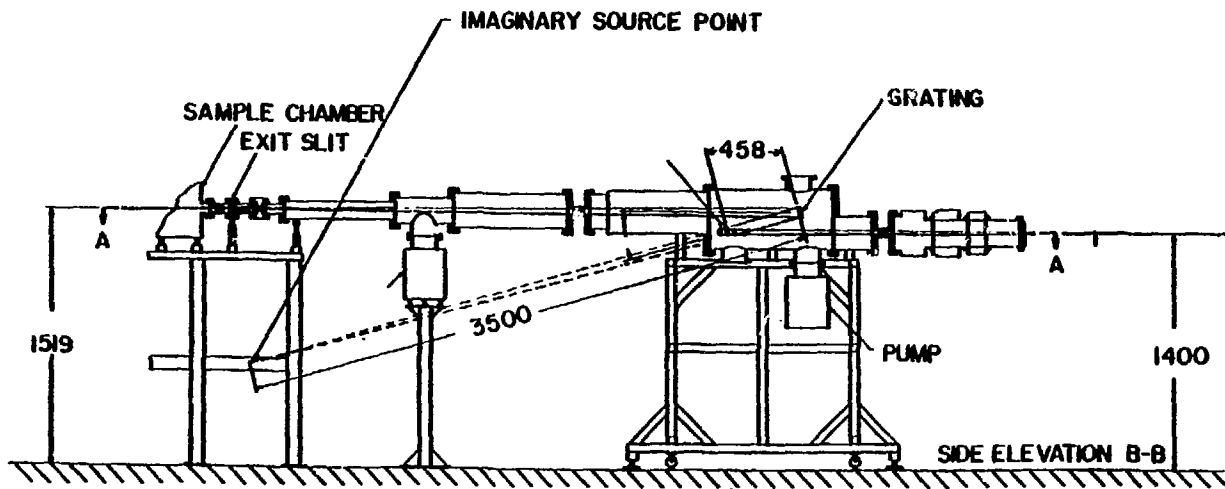
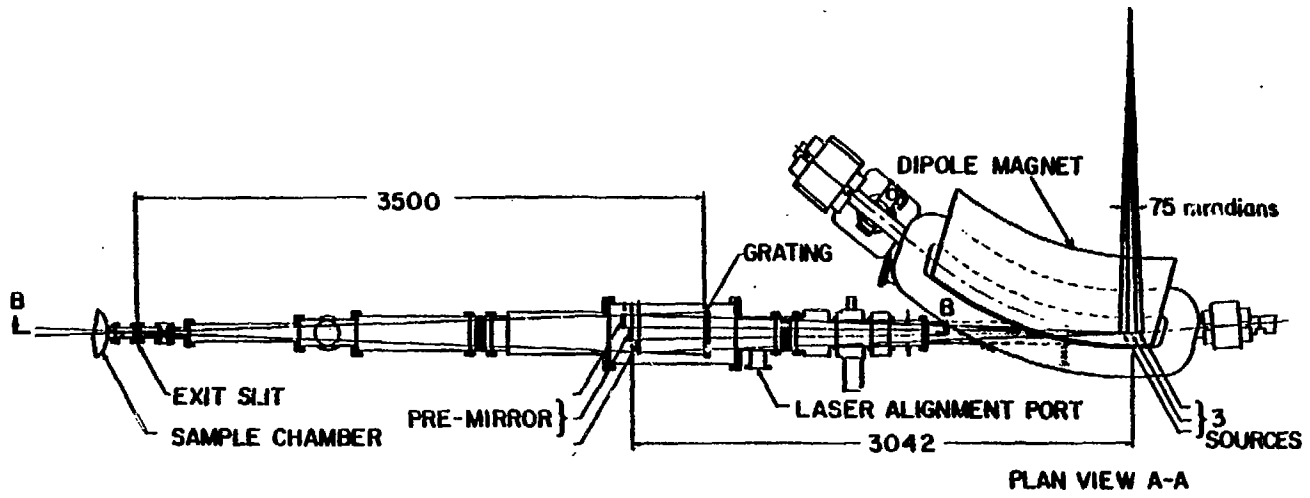


Figure 2.

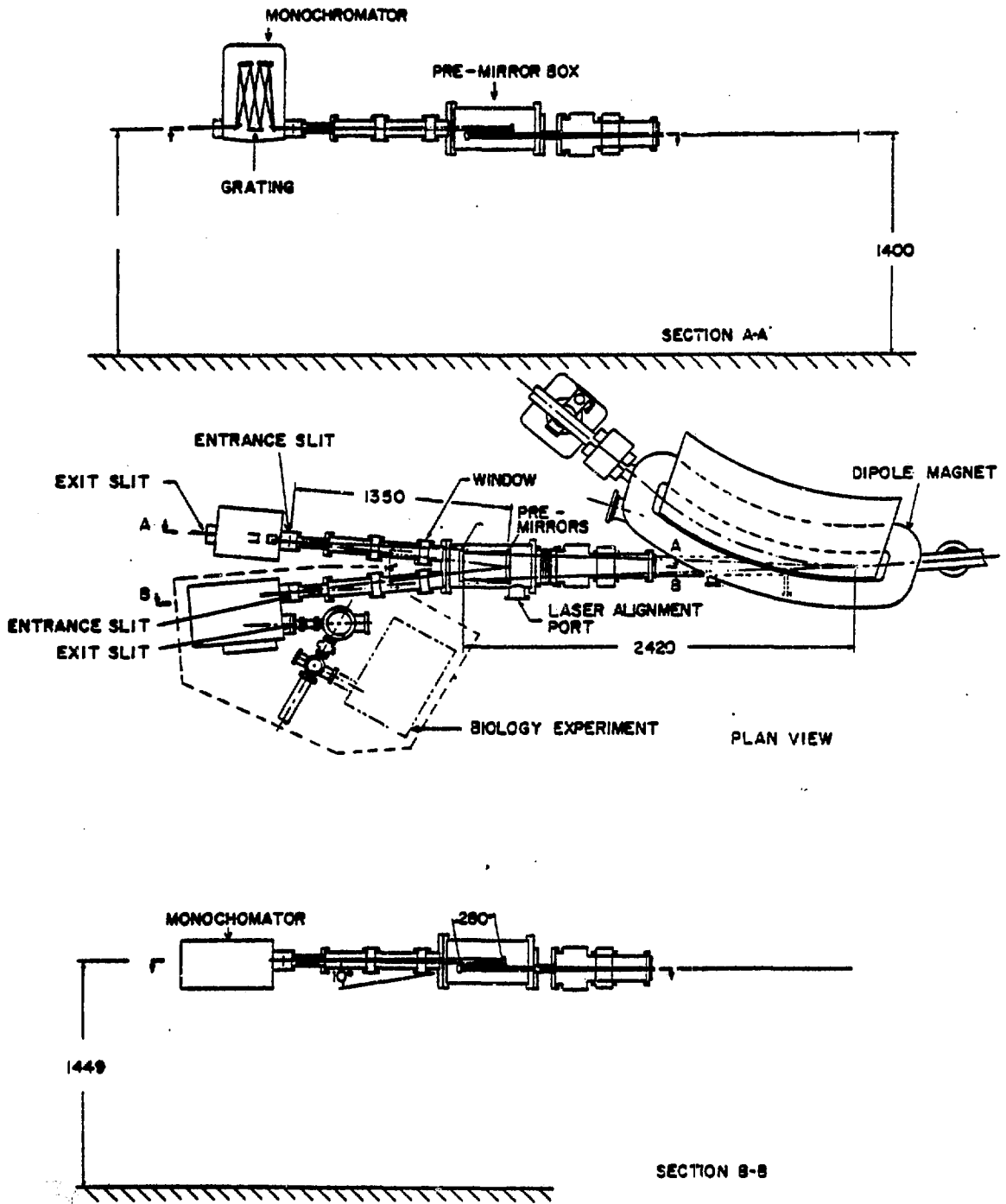


Figure 3.

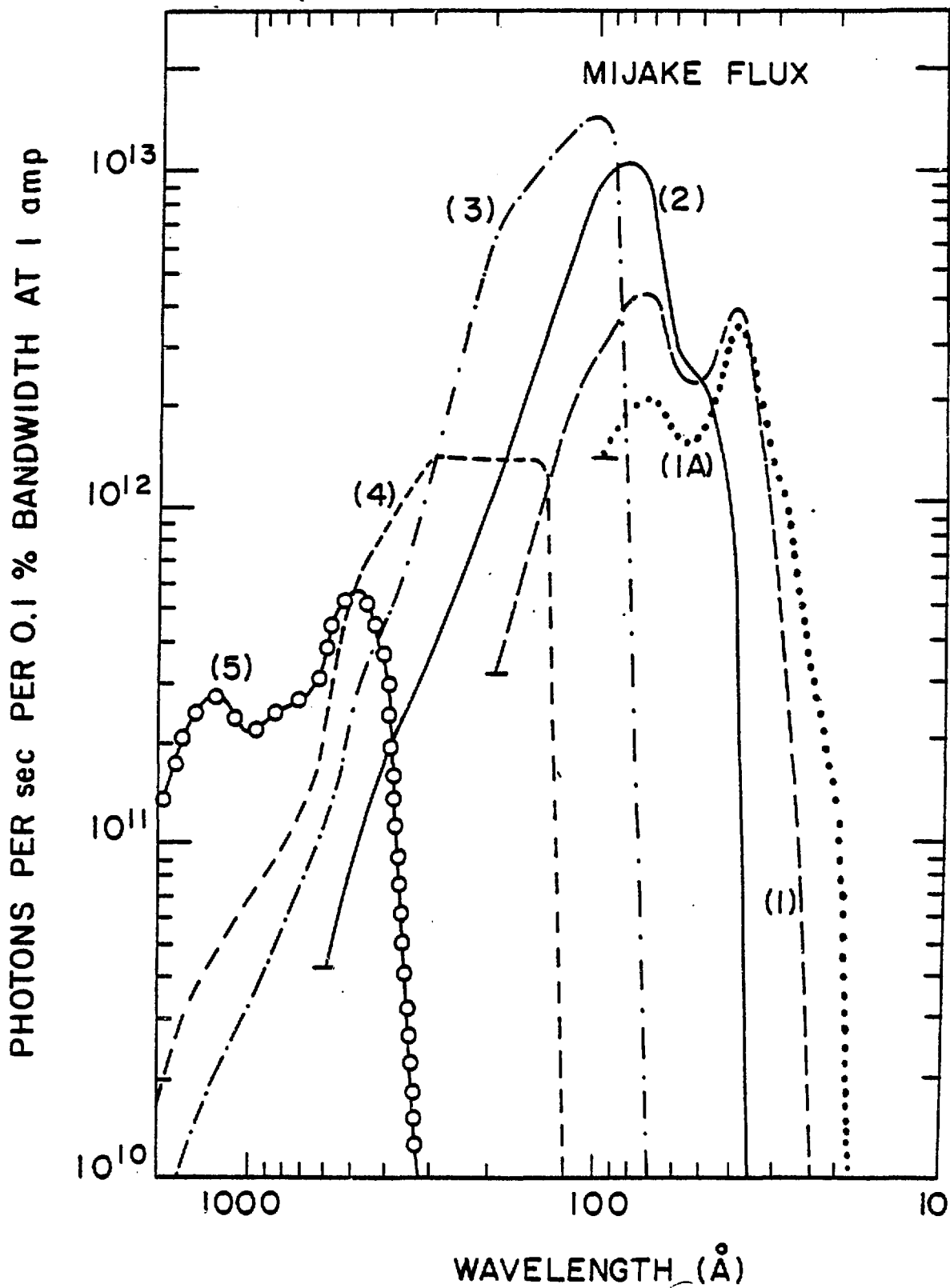


Figure 4.

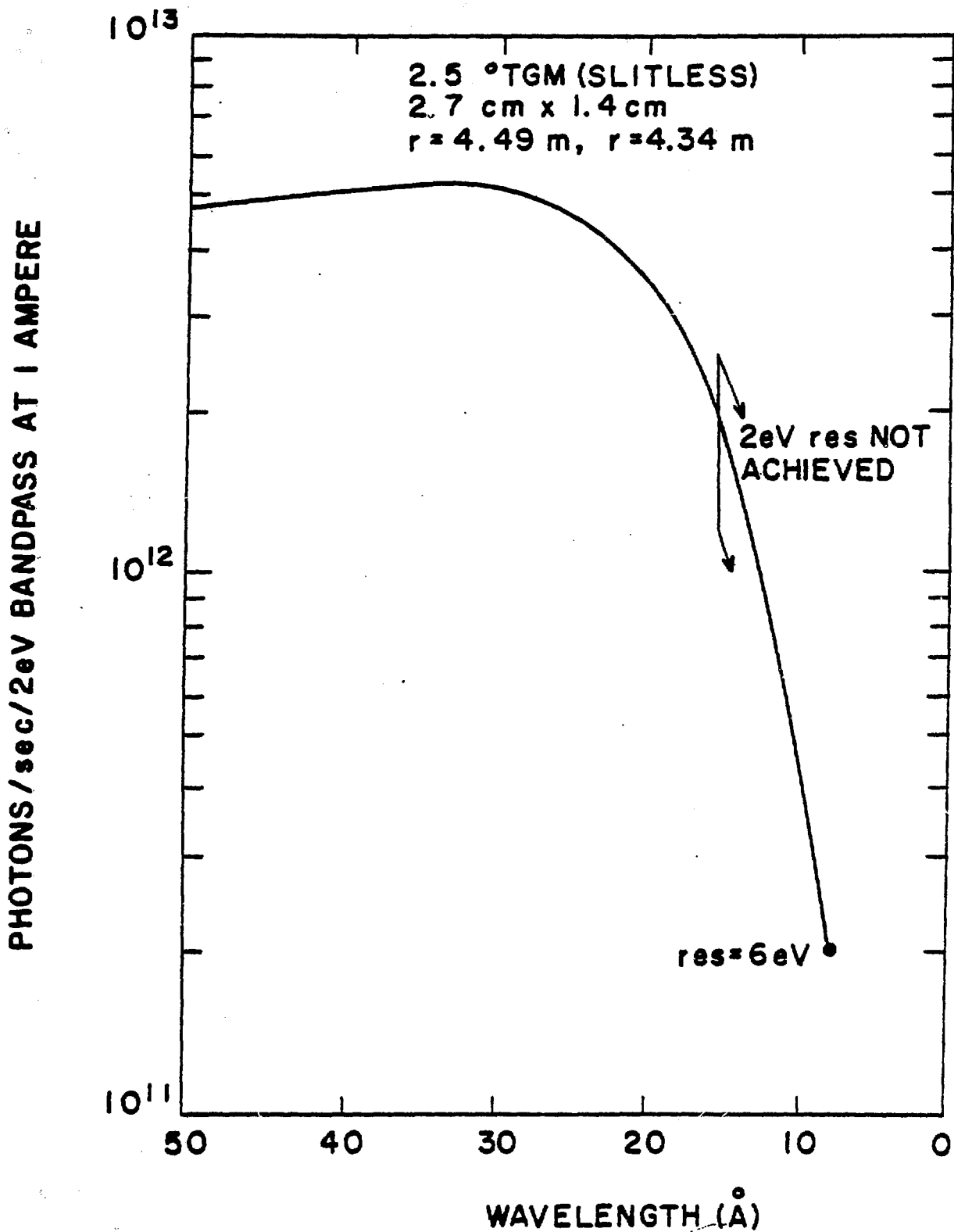


Figure 5.

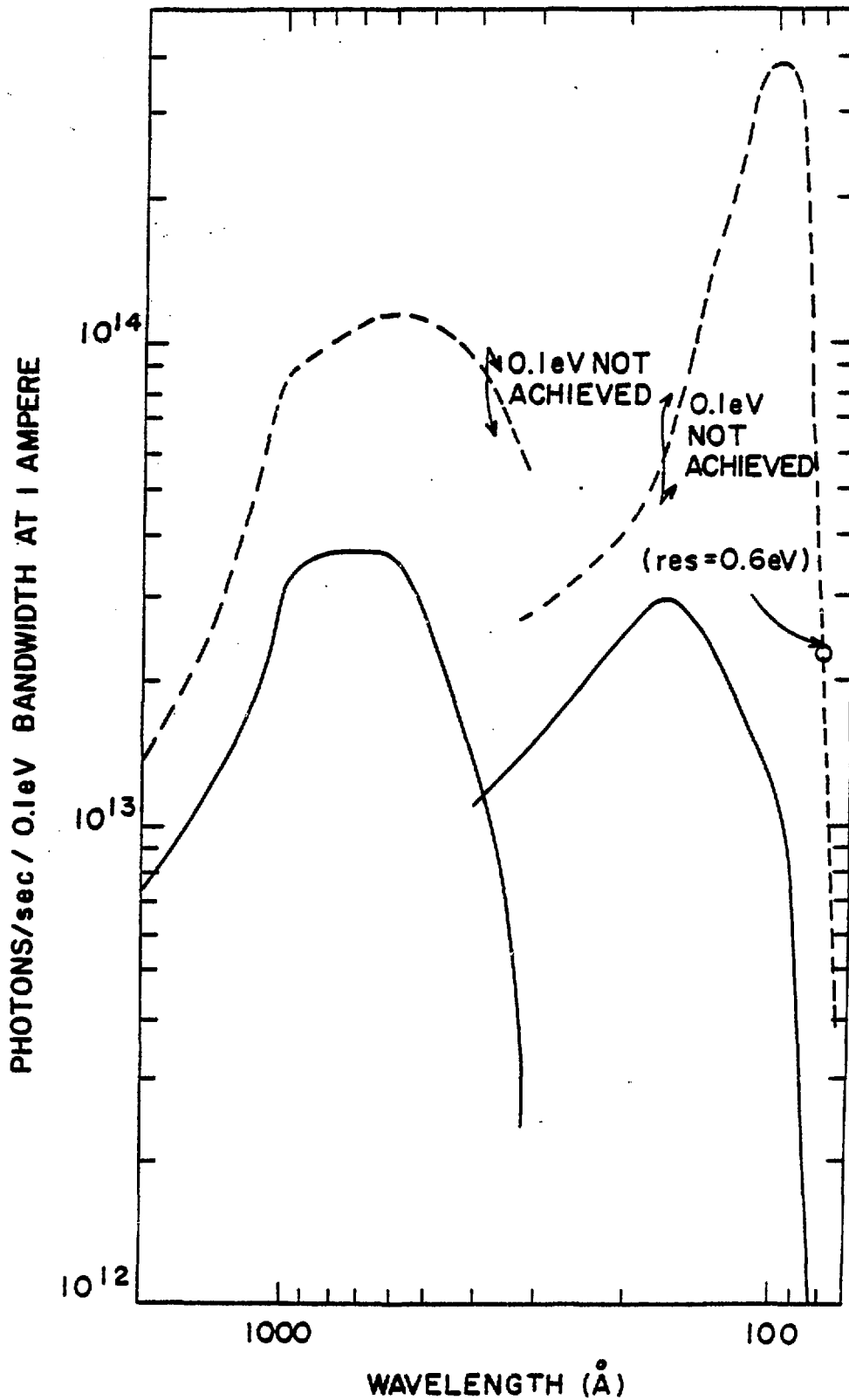
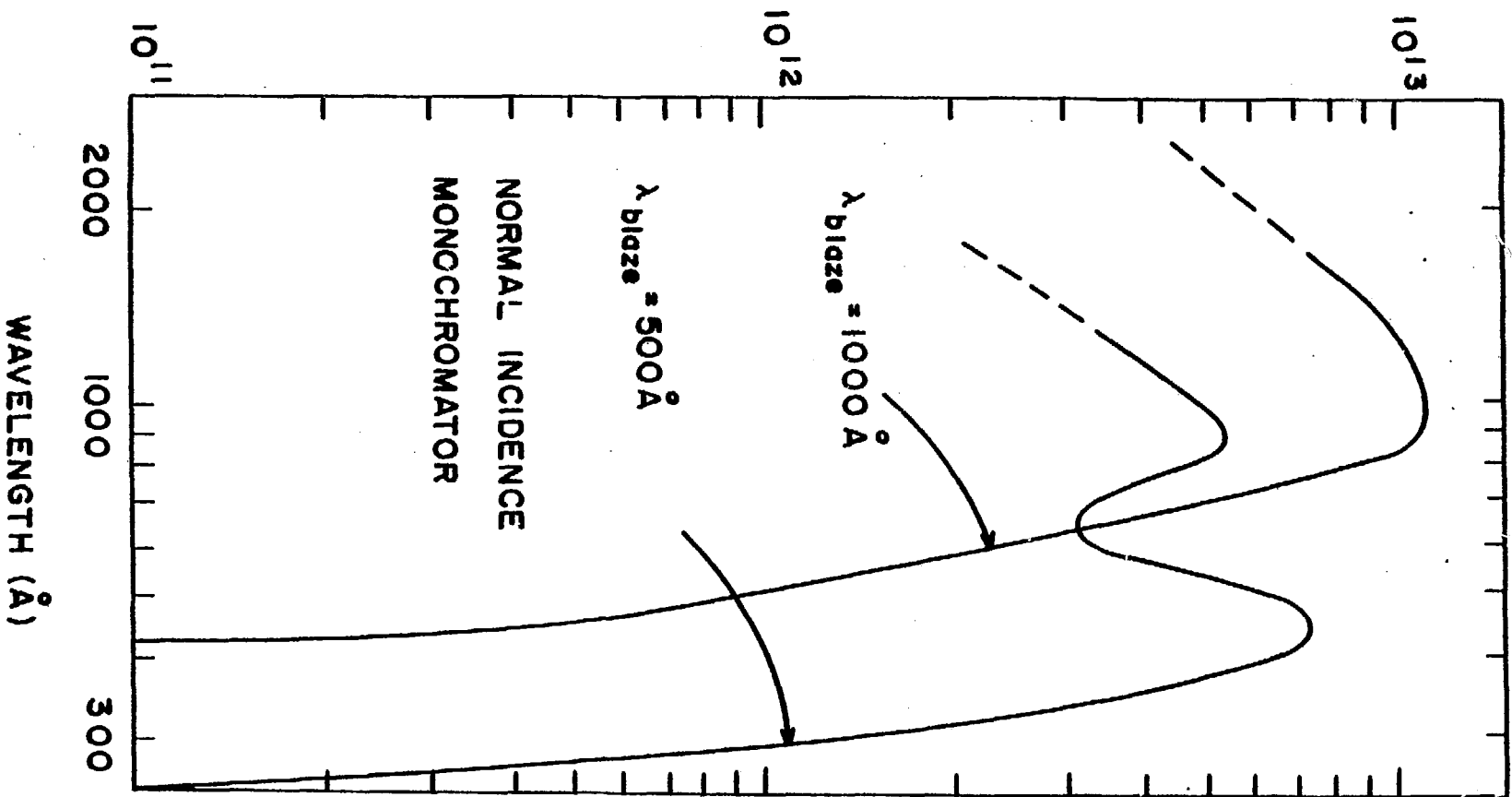


Figure 6.

PHOTONS/sec / 0.1 % BANDPASS AT 1 AMPERE



WAVELENGTH (\AA)

Figure 7.

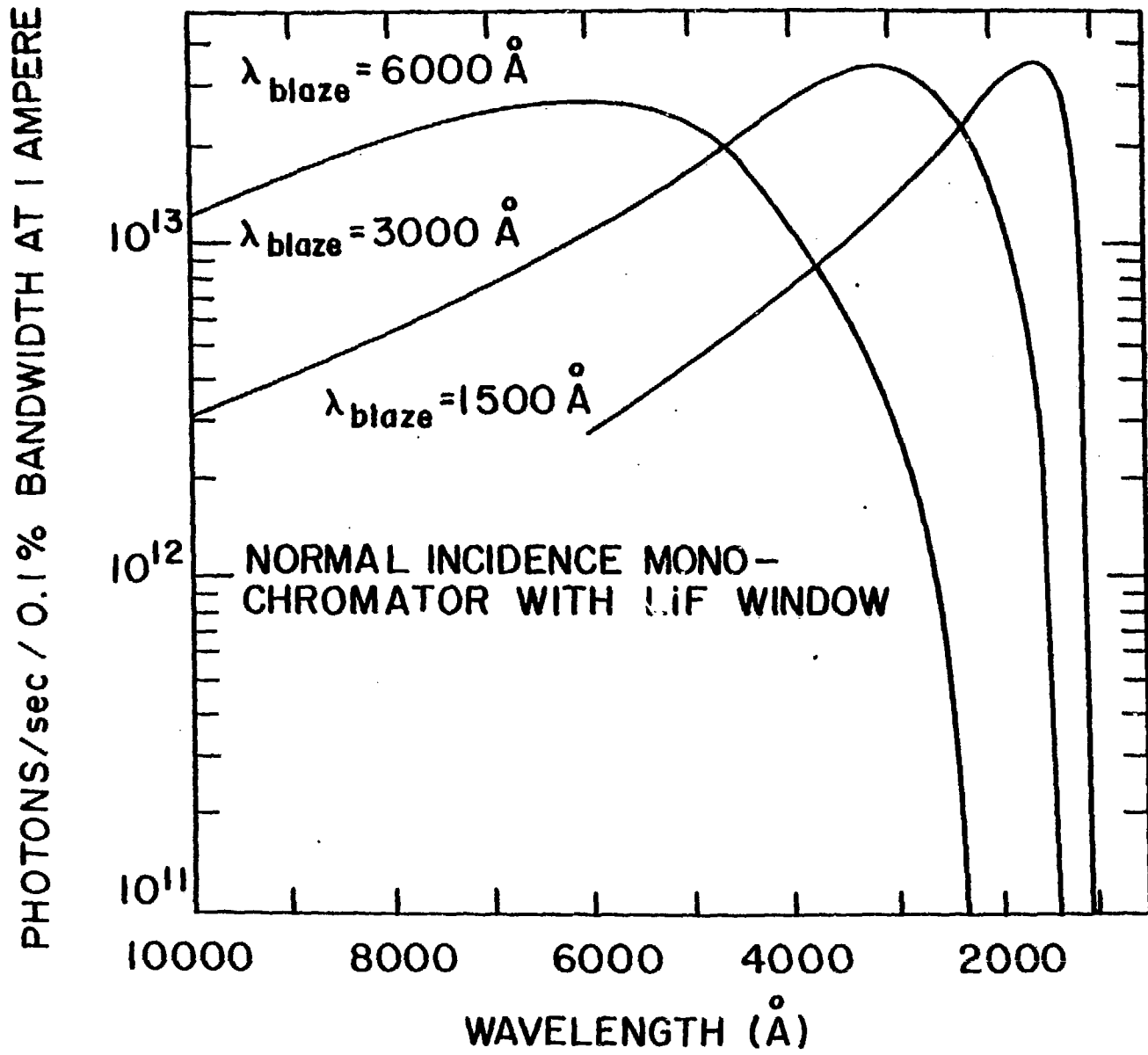


Figure 8.