

CP 7 10 1985

APPLICATION OF NORMAL FORM METHODS TO THE ANALYSIS  
OF RESONANCES IN PARTICLE ACCELERATORS

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ABSTRACT

The transformation to normal form in a Lie-algebraic framework provides a very powerful method for identifying and analysing non-linear behaviour and resonances in particle accelerators. The basic ideas are presented and illustrated by a simple yet important example.

**1. Introduction**

In this paper, we will discuss the advantages of using state-of-the-art mathematical tools, in particular normal form methods and maps in a Lie-algebraic framework<sup>(1-3)</sup>, to help us solve very complicated non-linear behaviour in particle accelerators. Although these methods were invented about 100 years ago, they would still have limited utility without the availability of modern computers and considerable mathematical development<sup>(3)</sup>. Two important developments are the advent of algebraic codes and differential algebra<sup>(4)</sup>.

In general, this is a study of systems of differential equations of the form

$$\dot{x} = f(x, \lambda, t); \quad x \in \mathbf{R}^n, \lambda \in \mathbf{R}^m, t \in \mathbf{R}. \quad (1.1)$$

In accelerator physics, these equations become Hamilton's equations

$$\zeta' = -[H, \zeta] \quad (1.2)$$

where  $f(x, \lambda, t) \Rightarrow H$ ,  $' = d/d\theta$ ,  $t \Rightarrow \theta$ , the new "time";  $x \in \mathbf{R}^n \Rightarrow \zeta = (x, p_x, z, p_z, \tau, p_\tau) \in \mathbf{R}^6$ , the "canonical variables" and the  $[\cdot, \cdot]$  are the familiar Poisson brackets. The relativistic Hamiltonian,  $H$ , in cylindrical coordinates will be used for explicitness.

$$H = -e\rho A_\theta - \rho \left\{ \left( \frac{p_0}{c} \right)^2 - m^2 c^2 - [p_\rho - eA_\rho]^2 - [p_z - eA_z]^2 \right\}^{1/2} \quad (1.3)$$

where  $\rho = x + \rho_0$ ;  $\rho_0$  is the radius of curvature of the central or "design" trajectory,  $p_0$  is the total energy of the system and  $e$  is the electronic charge. The magnetic vector potential  $A$  has the form

$$A = A(x, z, \tau; \theta); \quad A \in \mathbf{R}^3. \quad (1.4)$$

Equation (1.2) could be solved by brute-force numerical integration, which works up to a point for small systems like cyclotrons, but not for large systems. For example, the main ring of the Superconducting Supercollider has about 10 000 magnets; the Hamiltonian has 20 000 piece-wise continuous sections! Each particle trajectory must be integrated separately. Consequently, exploring phase space is extremely expensive and is accompanied by loss of numerical accuracy. In particular, the *symplectic condition* (conservation of phase space) for Hamiltonian systems is generally not preserved. Alternatively, we can use "state-of-the-art" methods to simplify the problem. This is essential for large systems. A truly enormous simplification of the problem is achieved by employing: 1) maps  $\zeta_j = \mathcal{M}(\theta)\zeta_i$ ; 2) normal forms (reduces the problem to its "simplest" form — isolates resonances); 3) a Lie-algebraic formulation (ensures symplecticity, minimizes the number of terms, enhances understanding).

## 2. Lie-Algebraic Maps

In the development, we use a Lie-algebraic formulation. The Poisson bracket  $[H, \zeta]$  in Hamilton's equations (1.2) satisfies all of the requirements of a Lie algebra, namely, the commutation relation

$$A * B = (A, B) = AB - BA \quad (2.1)$$

and the Jacobi identity

$$(A, (B, C)) + (B, (C, A)) + (C, (A, B)) = 0. \quad (2.2)$$

Note: This Lie algebra is *neither commutative nor associative*.

Traditionally, Taylor-series maps of the form  $\zeta_f = \mathcal{M}(\theta)\zeta_i$  have been employed, which when expanded become

$$\zeta_i^{FIN} = R_{ij}\zeta_j^{IN} + T_{ijk}\zeta_j^{IN}\zeta_k^{IN} + U_{ijkm}\zeta_j^{IN}\zeta_k^{IN}\zeta_m^{IN} + \dots \quad (2.3)$$

where, in general,  $R$ ,  $T$  and  $U$  have 36, 126, and 336 terms, respectively. In a Lie-algebraic representation of  $\mathcal{M}(\theta)$ , we write

$$\mathcal{M} = \mathcal{M}_f; \quad \mathcal{M}_f = e^{:f:}; \quad :f:g = [f, g] \quad (2.4)$$

where

$$e^{:f:} \stackrel{\text{def}}{=} \sum_{n=0}^{\infty} \frac{:f:^n}{n!}; \quad :f:^0 = I = \{\text{identity map}\}. \quad (2.5)$$

The map  $\mathcal{M}_f$  can be factored <sup>(1)</sup> into

$$\mathcal{M}_f = \mathcal{M}_2\mathcal{M}_3\mathcal{M}_4\dots = e^{:f_2:}e^{:f_3:}e^{:f_4:}\dots \quad (2.6)$$

where  $\mathcal{M}$  and  $\mathcal{M}_n$  are *symplectic maps* and  $f_n$  are homogeneous polynomials of degree  $n$ ;  $f_2, f_3$ , and  $f_4$  contain 21, 56 and 126, terms respectively (compare with (2.3)). The maps  $\mathcal{M}_n$  are elements of the symplectic group  $sp(2n, \mathbb{R})$ .

## 3. Normal Form

The normal form is the transformation  $\mathcal{N} = \mathcal{A}\mathcal{M}\mathcal{A}^{-1}$  such that  $\mathcal{N}$  is in its "simplest" form;  $\mathcal{M}$  must be expanded about a "centre manifold" or "closed orbit" (i.e. no  $H_1$  term in the Hamiltonian,  $H$ ), which implies no  $\mathcal{M}_1$  term in the map  $\mathcal{M}$ . The *symplectic map*  $\mathcal{A}$  is an  $n^{\text{th}}$  order canonical transformation that isolates the *tune shifts* and *resonances* to  $n^{\text{th}}$  order. Using the factored map, we can normalize  $\mathcal{M}$  order-by-order to obtain

$$\mathcal{N} = \mathcal{A}_n \dots \mathcal{A}_4 \{ \mathcal{A}_3 \{ \mathcal{A}_2 \mathcal{M} \mathcal{A}_2^{-1} \} \mathcal{A}_3^{-1} \} \mathcal{A}_4^{-1} \dots \mathcal{A}_n^{-1}. \quad (3.1)$$

We proceed as follows. From the eigenvectors of the matrix representation  $R$ , of  $\mathcal{M}_2$ , we get a canonical transformation that *rescales and block diagonalizes*  $R$ .

$$\mathcal{R} = \mathcal{A}_2 R \mathcal{A}_2^{-1} \Rightarrow \mathcal{R} = \begin{pmatrix} \mathcal{R}_1 & & 0 \\ & \mathcal{R}_2 & \\ 0 & & \mathcal{R}_3 \end{pmatrix}; \mathcal{R}_i = \begin{pmatrix} \cos(\mu_i) & \sin(\mu_i) \\ -\sin(\mu_i) & \cos(\mu_i) \end{pmatrix} \text{ or } \begin{pmatrix} 1 & L_i \\ 0 & 1 \end{pmatrix}. \quad (3.2)$$

Next we normalize  $\mathcal{M}_3$

$$\mathcal{N} = A_3 A_2 R e^{i g_3} \dots A_2^{-1} A_3^{-1} = e^{i G_3} \mathcal{R} e^{i g_3} \dots e^{-i G_3} = \mathcal{R} \exp \{ -(I - \mathcal{R}^{-1}) G_3 + g_3 \}. \quad (3.3)$$

A further simplification results if we expand  $(I - \mathcal{R}^{-1})$  and  $g_3$  in a resonance or complex basis, as follows:

$$h_{\pm} = x \pm i p_x; \quad v_{\pm} = z \pm i p_z. \quad (3.4)$$

Hence,

$$(I - \mathcal{R}^{-1}) |n, m\rangle = \left[ 1 - \sum_{n+m=3} e^{i(n-m)\mu} \right] |n, m\rangle, \quad (3.5)$$

and

$$g_3 = \sum_{n+m=3} A_{nm} |n, m\rangle \quad \text{where} \quad |n, m\rangle = h_+^{n_1} h_-^{m_1} v_+^{n_2} v_-^{m_2} \dots \quad (3.6)$$

and  $n, m, \mu$  are vectors with components  $(n_1, n_2, \dots)$  etc. Substituting (3.5) and (3.6) into (3.3) and solving for  $G_3$ , we obtain

$$G_3 = \sum_{n+m=3} \frac{A_{nm}}{\{1 - \exp[i(n-m)\mu]\}}. \quad (3.7)$$

All terms in  $g_3$  can be incorporated into  $G_3$  except when  $n - m = 0$  or  $(n - m) \cdot \mu = 2\pi k$ ;  $k = \text{integer}$ ; i.e.,  $(n_1 - m_1)\mu_x + (n_2 - m_2)\mu_z \neq 2\pi k$ ;  $k = 0, \pm 1, \pm 2, \dots$ . In a similar fashion, we could proceed to find  $G_4$ , etc. A further substitution of  $I_x = h_+ h_-$ ,  $I_z = v_+ v_-$ , ... — the *action invariants*— leads to

$$\mathcal{N} = \exp : \{ (\mu_x + \mu'_x \rho_r + \mu''_x p_r^2) I_x + (\mu_z + \dots) I_z + a_x I_x^2 + a_z I_z^2 + \dots \} : \quad (3.8)$$

where we have assumed no acceleration and no explicit resonances;  $\mu$  are the *phase advances*,  $\mu'$  and  $\mu''$  are the first- and second-order *chromaticities* and all terms proportional to  $I^n$  are non-linear *tune shifts*. If explicit resonances occur with  $(n - m) \cdot \mu = 2\pi k$ ,  $k = \text{integer}$ , these must be added to (3.8). Finally, the *normal form* has the remarkable property that

$$\mathcal{M}^n = A^{-1} \mathcal{N}^n A \quad (3.9)$$

which also results in an enormous saving of work when applicable.

As an example, consider the Hamiltonian for the superposition of a uniform magnetic field with quadrupole and sextupole components. The vector potential for this Hamiltonian up to 3<sup>rd</sup> order is

$$A_\theta = -\frac{x + \rho_0}{2} B_0 - \frac{1}{\rho_0} \{ Q_x x^2 + Q_z z^2 \} - \frac{S_x}{\rho_0} \left\{ \frac{x^3}{3} - x z^2 \right\}, \quad (3.10)$$

with  $A_\rho = A_z = 0$ ;  $Q_x, Q_z, S_x$  are the quadrupole and sextupole strength, respectively. (With a proper choice of parameters, this Hamiltonian models all of the non-linearities of the TASC superconducting cyclotron magnetic field up to third order.) Because (3.10) is independent of  $\theta$ , the Hamiltonian (1.3) has the trivial solution

$$\mathcal{M}(\theta) = \exp \left\{ - \int_0^\theta : H : d\theta' \right\} \implies \mathcal{M}(\theta) = e^{-\theta : H :} \quad (3.11)$$

which can be factored (with some work) into the form of (2.6). Because we are neglecting acceleration, the energy of each particle is conserved and we define  $p_r = (E_0 - E)/E_0$ , where  $E$  is the particle energy and  $E_0$  the energy of the central trajectory. If we factor (3.11), we find that  $f_3$  contains 19 complicated terms.

If we carry out the normalization procedure described above and there are no explicit resonances, the normal form contains  $g_2 = \mu_x h_+ h_- + \mu_z v_+ v_-$ , the “rotation matrix  $\mathcal{R}$ ” and

$$\begin{aligned}
g_3 = & \frac{(1 - \nu_x)}{4\nu_x\beta} h_+ h_- p_t + \frac{(1 - \nu_z)}{4\nu_z\beta} v_+ v_- p_t \\
& + \pi S_x \frac{k_x^2}{\nu_x\beta} h_+ h_- p_t - \pi S_x \frac{k_x k_z}{\nu_x\beta} v_+ v_- p_t \\
& + \pi k_x \frac{1 - \beta^2}{\nu_x\beta} p_t^3 + 2\pi S_x \frac{k_x^3}{3\nu_x^3\beta^3} p_t^3
\end{aligned} \tag{3.12}$$

which contains 4 *chromaticity* terms proportional to  $p_r$  and two time-of-flight aberration terms proportional to  $p_r^3$ , a substantial simplification. Furthermore, these are the terms that could potentially cause us trouble.

We see from (3.12) that if  $\nu_x = \nu_z = 1 \Rightarrow \mu_x = \mu_z = 2\pi$  (that is, we make the system resonant!), then 2 terms disappear and we are left with the terms proportional to  $S_x$ . If the remaining *chromaticity* terms can be removed (which they can), then  $\mathcal{R}_1 = \mathcal{R}_2 = \mathcal{I}$  or  $\mathcal{R}_\perp = \mathcal{I}$ . Hence

$$\mathcal{M} = \{A_3^{-1} A_2^{-1} \mathcal{N} A_2 A_3\} = \mathcal{I} e^{-i\alpha p_r^3} !! \tag{3.13}$$

Eq. (3.13) tells us that under these circumstances, all of the non-linearities, except the one proportional to  $p_r^3$ , annihilate each other up to 3<sup>rd</sup> order in the map! Although one would never want to make the whole accelerator resonant in this way, parts of it can be made so to great advantage.

#### 4. Conclusions

In the design and analysis of particle accelerators, we can achieve truly enormous savings in time and effort by:

1. using *maps*,
2. transforming to *normal form* to study *tune-shifts* and *resonances*, and
3. using the *Lie-algebraic* framework which enhances *insight and understanding*.

#### Acknowledgements

It is a pleasure to acknowledge many useful discussions with Stephen Douglas and Gordon Pusch.

#### References

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