

Jet Mass Spectra in Higgs + One Jet at NNLL

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The invariant mass of a jet is a benchmark variable describing the structure of jets at the LHC. We calculate the jet mass spectrum for Higgs plus one jet at the LHC at next-to-next-to-leading logarithmic (NNLL) order using a factorization formula. At this order, the cross section becomes sensitive to perturbation theory at the soft $m_{\text{jet}}^2/p_T^{\text{jet}}$ scale. Our calculation is exclusive and uses the 1-jettiness global event shape to implement a veto on additional jets. The dominant dependence on the jet veto is removed by normalizing the spectrum, leaving residual dependence from non-global logarithms depending on the ratio of the jet mass and jet veto variables. For our exclusive jet cross section these non-global logarithms are parametrically smaller than in the inclusive case, allowing us to obtain a complete NNLL result. Results for the dependence of the jet mass spectrum on the kinematics, jet algorithm, and jet size R are given. Using individual partonic channels we illustrate the difference between the jet mass spectra for quark and gluon jets. We also study the effect of hadronization and underlying event on the jet mass in PYTHIA. To highlight the similarity of inclusive and exclusive jet mass spectra, a comparison to LHC data is presented.

I. INTRODUCTION

There has been a rapidly expanding theoretical and experimental effort on techniques that exploit the substructure of jets (for a recent review see Ref. [1]). Jet substructure is of interest both for testing QCD and for identifying new physics. Much of the excitement in this field has been driven by the excellent performance of the ATLAS and CMS detectors, and the sophisticated jet measurements this has made possible at the LHC. Jet substructure measurements can for example be used to tag boosted heavy particles, whose decay products get collimated into a fat jet, or to test and tune Monte Carlo programs. Most theoretical work has focused on designing these techniques and observables with the help of Monte Carlo programs. At the same time, one would also like to know that these methods are under theoretical control and build confidence that higher-order effects are not significant. (For some recent progress in this direction see e.g. Refs. [2–7].)

As our underlying hard process we consider $pp \rightarrow H+1$ jet with gluon fusion $gg \rightarrow H$ as the underlying Higgs production mechanism. This process is convenient as it provides a clean setup with a single quark or gluon jet in the final state via the three basic partonic channels $gg \rightarrow Hg$, $gq \rightarrow Hq$, and $q\bar{q} \rightarrow Hg$. Of course, it is also important in its own right for Higgs measurements at the LHC, which rely on exclusive jet channels.

Here we focus on one of the simplest jet substructures: the invariant mass of a jet. A successful calculation of this benchmark observable will instill confidence in our ability to carry out analogous calculations for other more complicated jet substructure observables. Such analyses require incorporating both a resummation of large logarithms $\alpha_s^i \ln^j(m_J^2/p_T^{J,2})$ where m_J is the jet mass and p_T^J is the transverse momentum of the jet, as well as fixed-order perturbative corrections. This is made intricate by

the dependence on multiple variables. There has been a lot of recent work on the calculation (resummation) of the jet invariant mass spectrum for jets with a realistic angular size [2, 8–15] which we will review in more detail below. Some of the key theoretical issues that must be addressed for the LHC case include:

- Impact of summing large logarithms, $\ln(m_J^2/p_T^{J,2})$
- Soft radiation effects at the scale m_J^2/p_T^J
- Impact of initial-state radiation
- Color flow and hard process dependence
- Dependence on kinematics including rapidity cuts
- Jet algorithm and dependence on jet size R
- Inclusive ($\geq N$ jets) vs. exclusive ($= N$ jets)
- Impact of non-global logarithms (NGLs)
- Effect of hadronization on the spectrum
- Effect of underlying event on the spectrum
- Effect of pile-up on the spectrum
- Utility of using groomed jets with trimming [16], filtering [17], or pruning [18]

We now elaborate on several of these items. For a jet with $p_T^J \sim 300$ GeV, the jet mass peaks at $m_J \sim 50$ GeV, leading to large logarithms of $m_J^2/p_T^{J,2} \sim 36$. Therefore, a description of the peak region of the jet mass spectrum requires the all-order resummation of these logarithms. Soft radiation with momentum $k^\mu \sim m_J^2/p_T^J$ is generated by both initial and final-state particles and contributes at leading order in the power expansion to the jet mass. Since fixed-order corrections start to become relevant for resummation at next-to-next-to-leading logarithmic (NNLL) order, a proper treatment of the soft scale $\sim m_J^2/p_T^J$ is crucial at this order [2, 19–21]. Numerically, the importance of these fixed-order soft corrections is also well known from recent work up to N³LL for event shapes

in $e^+e^- \rightarrow \text{jets}$ [15, 22–24]. For processes with ≥ 2 jets at hadron colliders there are multiple color structures, and the corresponding color flow must be taken into account starting at next-to-leading-logarithmic (NLL) order [25].

The available freedom in defining a jet introduces a dependence of the jet mass spectrum on the choice of algorithm/clustering method and the jet size parameter R . There is also a choice of whether to use an inclusive or exclusive jet cross section, where the latter involves a veto on additional jets. The inclusive case has been studied at the LHC [26], and inclusive calculations tend to focus on the anti- k_T algorithm [27]. (Use of the anti- k_T jet algorithm avoids issues associated to clustering effects [28–31].) As we will emphasize further below, a key difference between the inclusive and exclusive cases are the form of the non-global logarithms [32, 33] that arise at $\mathcal{O}(\alpha_s^2)$ beyond the Born cross section due to multiple restrictions on phase space.

Let us summarize how the above issues have been studied so far in the literature on jet mass calculations. The first calculations were carried out for event shapes in $e^+e^- \rightarrow \text{jets}$ using hemisphere jet masses. Here factorization theorems are well established and calculations exist up to N³LL [15, 19, 21, 34–37]. In Refs. [2, 8] a factorization formula for exclusive N -jet cross sections at e^+e^- colliders was derived, where the angularity of a jet (which includes the jet mass as a special case) is measured. This result only depends on the class of the jet algorithm (such as cone or k_T -type), but suffers from non-global logarithms involving the jet veto and jet size R . The resummation of the jet mass in $e^+e^- \rightarrow 2$ jets with a jet veto was carried out at NLL in Ref. [10], which includes a resummation of NGLs in the large- N_c approximation. This same process was considered in Ref. [11], where the dominant R dependence of asymmetric thrust (which is related to jet mass) was obtained using a refactorization of the soft function. In Ref. [12], this refactorization was verified at $\mathcal{O}(\alpha_s^2)$ and the leading NGLs were obtained at this order.

For jet mass calculations in pp collisions one considers jets with large transverse momentum, p_T^J , and with rapidities η^J away from the beam axis. Recently, several inclusive jet mass calculations have been carried out [13–15]. In Ref. [13], the jet mass was calculated using only a jet function. This ignores important contributions from wide-angle soft radiation, which couples together multiple hard partons, depends on the choice of jet algorithm, and contains NGLs. In Ref. [14], the jet mass in $pp \rightarrow 2$ jets and $Z + 1$ jet were calculated at NLL, including a resummation of NGLs in the large- N_c approximation. Although this is an inclusive calculation (no jet veto), one should also note that hard emissions giving rise to additional jets are beyond the NLL order considered. In this case the dominant effect of the NGLs is on the peak of the jet mass distribution. Another inclusive calculation of the jet mass was carried out to obtain partial NNLL results in Ref. [15], by expanding around the threshold limit. Here dynamical threshold enhancement [38–40]

was used to argue that additional hard emissions are suppressed. Although NGLs were not resummed, their size was estimated, and found to mainly effect the peak region of the jet mass, as in Ref. [14].

Our calculation at NNLL is for the exclusive jet mass spectrum, so it is useful to highlight differences with the inclusive case. At NLL, for a given partonic channel and fixed momenta of the hard partons, the two cases simply differ by a multiplicative factor, except for their respective NGLs. In both cases the leading NGLs involve terms of the form

$$\alpha_s^2 \ln^2 \left(\frac{m_J^{\text{cut}2}}{p_{\text{cut}}^2} \right). \quad (1)$$

for the cumulant jet mass spectrum integrated up to m_J^{cut} . For the inclusive jet mass spectrum, p_{cut} is a hard scale $\simeq p_T^J$ and the NGLs are therefore large logarithms that are parametrically of the same size as other $\alpha_s^i \ln^j(m_J^2/p_T^{J2})$ terms, and are thus part of the NLL result. Hence, in this case a complete resummation at NLL (or beyond) requires the NGLs to be resummed to all orders, which practically is currently only possible in the large- N_c approximation. In contrast, in the exclusive case p_{cut} is an adjustable parameter and is related to the jet veto. Here we have both $m_J^2 \ll p_T^{J2}$ and $p_{\text{cut}}^2 \ll p_T^{J2}$, so the logarithms in Eq. (1) are smaller than in the inclusive case. In particular, for fixed p_{cut} there is a point in the m_J spectrum where the NGLs vanish, and there is a region about this point where the NGLs are not large logarithms. An estimate for the size of this region can be obtained from the series of three NGL terms (log-squared, log, and non-log) that are known for the hemisphere jet masses [41, 42]. When all the terms in this series are of similar magnitude the logarithmic enhancement is not dominant, and the NGLs do not need to be resummed. This occurs for $1/8 \leq m_J^{\text{cut}2}/p_{\text{cut}}^2 \leq 8$. We will numerically explore the size of this region in our exclusive jet mass calculation, and demonstrate that the region is large enough that we may consider the non-global logarithms to not be large. This can be contrasted with Fig. 3 of Ref. [10], which shows that the presence of an unmeasured region of phase space makes large NGLs unavoidable in the inclusive case [14, 15].

It should also be noted that although exclusive jet cross sections are not necessary for jet mass spectra, they are important in their own right because many Higgs and new physics searches categorize the data by the number of jets to improve their sensitivity. For example, the importance of the Higgs + 1 jet channel in $H \rightarrow \tau\tau$ and $H \rightarrow WW^*$ was pointed out in Refs. [43, 44]. Recently a NLL resummation of jet veto logarithms was carried out in the context of Higgs plus jets in Ref. [45].

Our calculation of the jet mass is centered on using the N -jettiness global event shape [46] to define jets, instead of a more traditional jet algorithm. For an event with N jets, N -jettiness assigns all particles to $N + 2$ regions, corresponding to the N jets and two beams. We calculate the cross section for $pp \rightarrow H + 1$ jet at NNLL,

fully differential in the contributions of each region to 1-jettiness. For the jet region, this contribution yields the jet invariant mass. The contribution from the remaining two beam regions are used to implement the jet veto. In each of these variables there is a series of large double logarithms that must be summed.

An advantage of using N -jettiness is that the jet veto is made through a jet mass-type variable, rather than a p_T variable. Therefore, the structure of the perturbation theory, which is simultaneously differential in these two kinematic variables, is simpler. In particular, there is a QCD factorization formula for this cross section [46, 47], obtained by making use of Soft-Collinear Effective Theory (SCET) [20, 48–50]. For the experimentally more realistic case of measuring m_J with a p_T veto variable one must simultaneously deal with a thrust-like resummation and a p_T -type resummation.

Returning to our list of theoretical issues from the beginning, the use of N -jettiness allows us to carry out the summation of large logarithms at NNLL while properly accounting for soft radiation effects and initial-state radiation. We also use it to calculate the dependence of the jet mass spectrum on the jet kinematics, the jet size, and the definition of the jet region. Results are shown for individual partonic channels, $gg \rightarrow Hg$ and $gq \rightarrow Hq$, illustrating the differences between quark and gluon jets, as well as the full $pp \rightarrow H + 1$ jet process from the Higgs coupling through a top quark loop. To investigate the differences between exclusive and inclusive jet mass measurements we compare our results with PYTHIA and also to ATLAS jet mass data [26]. We also analytically explore the effect of NGLs on the jet mass spectrum, and the effect of hadronization and underlying event with PYTHIA [51, 52].

Thus, we address all items in the list of issues except for the last two, for which some brief comments are in order. Methods for removing pile-up contributions to jet observables have been discussed in e.g. Refs. [53, 54], and direct pile-up calculations are beyond the scope of our work. Finally, it is known that grooming jets has a large impact on their soft radiation and causes significant changes to the jet mass spectrum. We do not attempt to analytically control the effects of jet-grooming methods here.

In calculating the jet mass we consider both absolute and normalized spectra. Normalizing the jet mass spectrum reduces the perturbative uncertainty, and turns out to remove the dominant dependence on the jet veto variable. In particular, the jet veto dependence cancels up to NLL if we consider a particular partonic channel and fixed jet kinematics. We will show that this cancellation remains effective when summing over partonic channels and integrating over a range of jet momenta.

In Sec. II, we discuss the kinematics and several jet definitions based on N -jettiness, exploring their features. The technical details of our calculation are presented in Sec. III. Here we discuss the factorization formula for the cross section, the refactorization of the soft func-

tion, non-global logarithms, and the choice of running scales. Sec. IV and Sec. V contain our numerical results for the individual partonic channels and for $pp \rightarrow H + 1$ jet, showing the dependence of the jet mass spectrum on the jet veto cut, the order in perturbation theory, the jet kinematics, the jet definition, the jet area, on gluon versus quark jets, and on NGLs. Using PYTHIA8, in Sec. VI we analyze the hard process dependence for gluon jets, compare inclusive versus exclusive jet mass spectra, study the dependence on classic jet algorithms, and look at the impact of hadronization and underlying event. We also compare our NNLL exclusive jet results with PYTHIA for the same jet definition and kinematics, and compare them with inclusive jets from the LHC data. We conclude in Sec. VII. Detailed ingredients for the NNLL cross section are summarized in appendices.

II. KINEMATICS AND JET DEFINITIONS

We describe the process $pp \rightarrow H + 1$ jet using the transverse momentum p_T^J of the jet, the pseudorapidity η_J of the jet, and the rapidity Y of the hard collision relative to the CM frame of the colliding protons. The 1-jettiness event shape is defined as [46]

$$\mathcal{T}_1 = \sum_k \min \left\{ \frac{2q_J \cdot p_k}{Q_J}, \frac{2q_a \cdot p_k}{Q_a}, \frac{2q_b \cdot p_k}{Q_b} \right\}, \quad (2)$$

where a, b denote the two beams and J the jet, the q_i are massless reference momenta and the Q_i are normalization factors. For the reference momenta we take

$$q_J^\mu = E_J (1, \vec{n}_J), \quad q_{a,b}^\mu = x_{a,b} \frac{E_{\text{cm}}}{2} (1, \pm \hat{z}). \quad (3)$$

The jet energy E_J and jet direction \vec{n}_J can be predetermined with a suitable jet algorithm. The jet algorithm dependence this induces on \mathcal{T}_1 is power suppressed [46], and we will use anti- k_T .¹ The unit vector \hat{z} points along the beam axis, and the momentum fractions x_a and x_b are fixed in terms of the total invariant mass Q and rapidity Y ,

$$x_a x_b E_{\text{cm}}^2 = Q^2 = (q_J + q_H)^2, \quad \ln \frac{x_a}{x_b} = 2Y = \ln \frac{(1, -\hat{z}) \cdot (q_J + q_H)}{(1, \hat{z}) \cdot (q_J + q_H)}, \quad (4)$$

where q_H^μ denotes the momentum of the Higgs. For later convenience we also introduce the notation

$$s_{ij} = 2q_i \cdot q_j. \quad (5)$$

¹ If $Q_J = 2E_J$ then an equally good choice would be to minimize \mathcal{T}_1 with respect to the axis \vec{n}_J . A fast algorithm to carry out this minimization has been devised in Ref. [55], using a slightly different N -jettiness measure than the ones we use here.

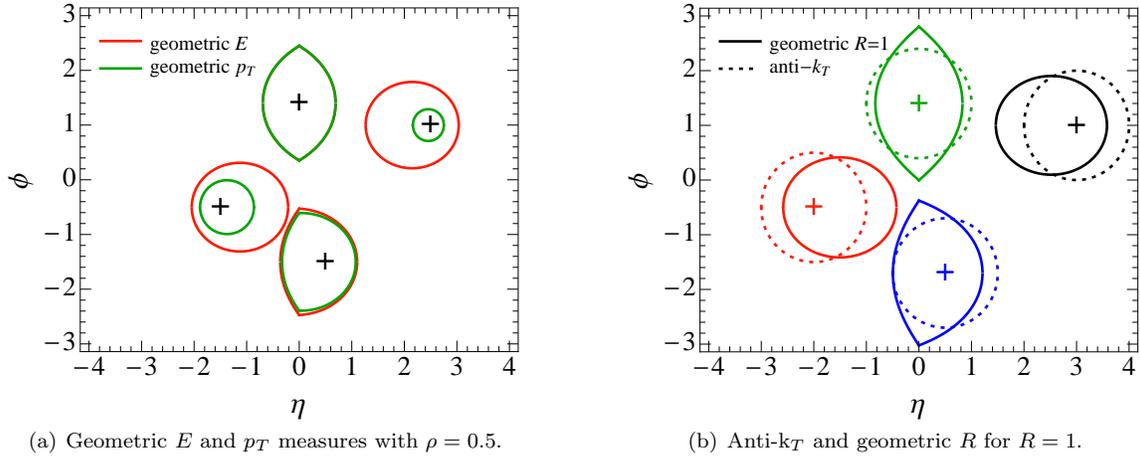


FIG. 1: Comparison of the jet regions for different jet measures at different η and ϕ . The “+” marks the jet direction \vec{n}_J .

The minimum in Eq. (2) divides the total phase space into 3 regions, one for each beam and one for the jet. We denote their contributions to \mathcal{T}_1 as \mathcal{T}_a and \mathcal{T}_b for the two beam regions, and \mathcal{T}_J for the jet region, so

$$\mathcal{T}_1 = \mathcal{T}_J + \mathcal{T}_a + \mathcal{T}_b. \quad (6)$$

The contribution of the jet, \mathcal{T}_J , is directly related to the jet’s invariant mass m_J

$$\begin{aligned} m_J^2 &= p_J^2 = (\vec{n}_J \cdot p_J)(n_J \cdot p_J) - \vec{p}_{J\perp}^2 \\ &= 2q_J \cdot p_J [1 + \mathcal{O}(\lambda^2)] \\ &= Q_J \mathcal{T}_J [1 + \mathcal{O}(\lambda^2)], \end{aligned} \quad (7)$$

where p_J^μ is the full jet momentum defined by summing all particles in the \mathcal{T}_J -region, $n_J^\mu = (1, \vec{n}_J)$ and $\vec{n}_J^\mu = (1, -\vec{n}_J)$ are defined by the predetermined jet direction \vec{n}_J , and the power counting parameter λ scales as $\lambda^2 \sim \mathcal{T}_J/E_J \sim m_J^2/E_J^2$. In the second line of Eq. (7) we used the fact that \vec{n}_J and the exact direction of the N -jettiness jet, \vec{p}_J , differ by very little, such that $p_{J\perp}/(\vec{n}_J \cdot p_J) \sim \lambda^2$. The difference between these two jet directions affects the jet boundary, which changes the contribution of soft radiation to the jet p_T , but only by a small amount $\sim \lambda^2$. We also used that the large jet momentum $\vec{n}_J \cdot p_J = \vec{n}_J \cdot q_J [1 + \mathcal{O}(\lambda^2)]$. For a jet with $p_T^J \sim 300$ GeV these $\mathcal{O}(\lambda^2)$ power corrections are $1/36 \sim 3\%$ in the peak region, and hence negligible relative to the perturbative uncertainties at NNLL. Investigating the jet mass spectra for the exact $m_J^2 = p_J^2$ vs. using $m_J^2 = Q_J \mathcal{T}_J$ in PYTHIA, we also find that they are indistinguishable.

The details of the beam and jet regions selected by the minimum condition in Eq. (2) depend on the normalization factors Q_i . Since their values affect which particles are grouped into the beam and jet regions, they constitute a jet measure. They also impact the geometric shape of the jet area. Differences between measures are therefore similar to the different choices for jet-algorithms (anti- k_T , Cambridge-Aachen, cone, etc.). We will consider a variety of choices:

- invariant-mass measure:

$$Q_J = Q_a = Q_b = Q \quad (8)$$

- geometric p_T measure:

$$\begin{aligned} Q_J &= 2\rho |\vec{q}_{iT}| = 2\rho E_J / \cosh \eta_J \\ Q_{a,b} &= x_{a,b} E_{\text{cm}} = e^{\pm Y} Q \end{aligned} \quad (9)$$

- geometric E measure:

$$\begin{aligned} Q_J &= 2\rho E_J \\ Q_{a,b} &= x_{a,b} E_{\text{cm}} = e^{\pm Y} Q \end{aligned} \quad (10)$$

- geometric R measure:

$$\begin{aligned} Q_J &= 2\rho(R, \eta_J) E_J \\ Q_{a,b} &= x_{a,b} E_{\text{cm}} = e^{\pm Y} Q \end{aligned} \quad (11)$$

where $\rho(R, \eta_J)$ fixes the area of the jet in (η, ϕ) -space to be πR^2 .

In all cases ρ is a dimensionless parameter that allows one to change the size of the jet region. In the geometric R case ρ is fixed in terms of the jet radius parameter R .² The choice of $Q_{a,b}$ in the geometric measures removes the dependence in q_a^μ/Q_a and q_b^μ/Q_b on the total rapidity Y . This is useful in the presence of missing energy, which prohibits the measurement of the boost Y of the partonic center-of-mass frame. Since for the geometric measures $Q_J \sim E_J$, they are all insensitive to the total jet energy. For the geometric p_T case we have explicitly

$$\frac{2q_i \cdot p_k}{q_{iT}} = p_{kT} \left(2 \frac{m_{kT}}{p_{kT}} \cosh \Delta y_{ik} - 2 \cos \Delta \phi_{ik} \right) \quad (12)$$

² For the multijet case we would use the same $\rho(R, \eta_J)$ for each jet that is determined when they do not overlap.

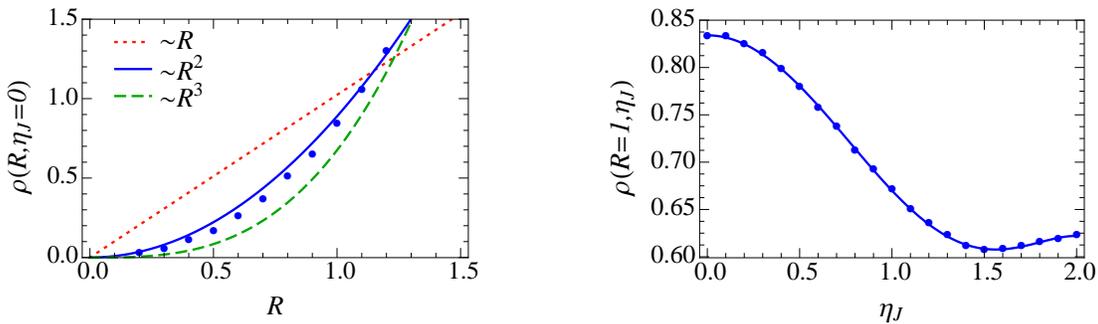


FIG. 2: Numerical results for $\rho(R, \eta_J)$ in the geometric R measure. Left: Dependence on R for $\eta_J = 0$, which is $\sim R^2$ as expected. Right: Dependence on η_J for $R = 1$. To solve for ρ we use a fit (solid line) to the true η_J dependence (dots).

where $\Delta y_{ik} = y_i - y_k$, $\Delta \phi_{ik} = \phi_i - \phi_k$ are the differences in rapidity and azimuthal angle between the direction of jet i and particle k , and $m_{kT}^2 = p_{kT}^2 + m^2$ for a particle of mass m . For massless particles we thus get

$$\begin{aligned} \frac{2q_i \cdot p_k}{q_{iT}} &= p_{kT} (2 \cosh \Delta y_{ik} - 2 \cos \Delta \phi_{ik}) \\ &\approx p_{kT} [(\Delta y)^2 + (\Delta \phi_{ik})^2]. \end{aligned} \quad (13)$$

The jet regions for geometric p_T and E are roughly circular, as shown in Fig. 1(a). They become smaller at large rapidities for geometric p_T , while they stay of comparable size for geometric E .

For geometric R , numerical results for the parameter $\rho(R, \eta_J)$ as function of R and η_J are shown in Fig. 2. The left panel shows that the dependence on the jet radius R is approximately $\rho \sim R^2$, as expected. The right panel illustrates the dependence on η_J for fixed $R = 1$, showing that ρ approaches a constant for large η_J , i.e. when the jet becomes close to the beam. When using geometric R in our results below, we use for convenience a fit of the η_J dependence for fixed value of R . For example, for $R = 0.5, 1, 1.2$ we have for $|\eta_J| \leq 2$

$$\begin{aligned} \rho(R = 0.5, \eta_J) &= 0.164 + 0.037\eta_J^2 - 0.009\eta_J^4 + 0.0008\eta_J^6, \\ \rho(R = 1, \eta_J) &= 0.834 - 0.233\eta_J^2 + 0.077\eta_J^4 - 0.008\eta_J^6, \\ \rho(R = 1.2, \eta_J) &= 1.272 - 0.376\eta_J^2 + 0.101\eta_J^4 - 0.010\eta_J^6. \end{aligned} \quad (14)$$

Note that for $R = 0.5$ the parameter ρ increases rather than decreases with η_J . A comparison of the jet regions for geometric R with anti- k_T jets is shown in Fig. 1(b). Although their areas are chosen to be the same, the geometric R jets are not perfectly circular and have an “offset” between the jet direction and the center of the jet

region. The former (latter) effect decreases (increases) with $|\eta_J|$. For a smaller jet radius of $R = 0.5$ the geometric R jets become more circular also at central rapidities and very close to anti- k_T jets. In Ref. [55] a modification of N -jettiness was introduced that matches anti- k_T closely for any R . However, this definition reintroduces a region of phase space that belongs neither to the jet nor the beams, making it more complicated for calculations.

III. CALCULATION

A. Factorization Formula

We start by rewriting the phase space integrals for the hard kinematics in terms of the rapidity η_J and transverse momentum p_T^J of the jet and the total rapidity Y ,

$$\begin{aligned} \int \frac{dx_a}{x_a} \int \frac{dx_b}{x_b} \int \frac{d^3 \vec{q}_H}{E_H} \int \frac{d^3 \vec{q}_J}{E_J} (2\pi)^4 \delta^4(q_a + q_b - q_J - q) \\ = \int d\eta_J dp_T^J dY \frac{1}{2\pi} \frac{p_T^J}{Q^2 + m_H^2}. \end{aligned} \quad (15)$$

The variables were defined in Sec. II, and we used azimuthal symmetry and the relations

$$\begin{aligned} p_T^J &= \frac{Q^2 - m_H^2}{2Q \cosh(\eta_J - Y)}, \\ Q &= p_T^J \cosh(\eta_J - Y) + \sqrt{p_T^{J2} \cosh^2(\eta_J - Y) + m_H^2}. \end{aligned} \quad (16)$$

Most of our plots will be normalized and for fixed values of η_J , p_T^J , and Y , in which case the phase space factor in Eq. (15) drops out.

Our calculation relies on the N -jettiness factorization formula in Ref. [47], which we here specialize to the case of 1-jettiness:

$$\begin{aligned} \frac{d^3 \sigma^{H+1j}}{d\eta_J dp_T^J dY d\mathcal{T}_a d\mathcal{T}_b d\mathcal{T}_J} &= \frac{p_T^J}{4\pi E_{\text{cm}}^2 (Q^2 + m_H^2)} \sum_{\kappa} H_{\kappa}(\{q_i^{\mu}\}, \mu) \int dt_a B_{\kappa_a}(t_a, x_a, \mu) \int dt_b B_{\kappa_b}(t_b, x_b, \mu) \\ &\times \int ds_J J_{\kappa_J}(s_J, \mu) S_{\kappa} \left(\mathcal{T}_a - \frac{t_a}{Q_a}, \mathcal{T}_b - \frac{t_b}{Q_b}, \mathcal{T}_J - \frac{s_J}{Q_J}, \left\{ \frac{q_i^{\mu}}{Q_i} \right\}, \mu \right). \end{aligned} \quad (17)$$

channel	κ_a	κ_b	κ_J
$gg \rightarrow Hg$	g	g	g
$gq \rightarrow Hq$	g	q	q
$qg \rightarrow Hq$	q	g	q
$g\bar{q} \rightarrow \bar{H}q$	g	\bar{q}	\bar{q}
$\bar{q}g \rightarrow \bar{H}q$	\bar{q}	g	\bar{q}
$q\bar{q} \rightarrow Hg$	q	\bar{q}	g
$\bar{q}q \rightarrow Hg$	\bar{q}	q	g

TABLE I: Values of κ for the different partonic channels.

The N -jettiness variables \mathcal{T}_a , \mathcal{T}_b , and \mathcal{T}_J were defined in Sec. II. The hard function H_κ contains the short-distance matrix element for producing a Higgs plus a jet, the beam functions B_{κ_a} and B_{κ_b} describe the collinear initial-state radiation and contain the PDFs, the jet function J_{κ_J} characterizes the collinear final-state radiation, and the soft function S_κ describes soft radiation effects.³ The sum over $\kappa = \{\kappa_a, \kappa_b, \kappa_J\}$ runs over the possible flavors $\kappa_i \in \{g, u, \bar{u}, d, \dots\}$ of the two incoming and one outgoing parton. The possible combinations, corresponding to the various partonic channels, are listed in Table I.

The power of factorization is that it allows one to evaluate the various fixed-order pieces at their natural scales, where they contain no large logarithms. We then use the RG evolution of each of these functions to evolve them to a common scale μ , resumming the logarithms of m_J^2/p_T^2 and $Q_i \mathcal{T}_i/p_T^2$. This evolution is implicit in Eq. (17), by writing all functions as evaluated at the common scale μ . The factorization formula with all evolution factors written out explicitly is given in Eq. (28) below. Our choice of scales is discussed in Sec. III C. Power corrections to Eq. (17) arise from so-called nonsingular corrections, which are suppressed by a relative $\mathcal{O}(m_J^2/Q^2)$ in this differential cross, and are not considered here.

The cross section in Eq. (17) is differential in the 1-jettiness contributions from the jet and the beams \mathcal{T}_J , \mathcal{T}_a , and \mathcal{T}_b . As we will see, the shape of the jet mass spectrum is independent of the jet veto for a reasonable range of $\mathcal{T}_{a,b}$ values. For simplicity we impose a common cut $\mathcal{T}_{a,b} \leq \mathcal{T}^{\text{cut}}$. We also convert \mathcal{T}_J to the jet mass m_J using Eq. (7), and so consider

$$\sigma(m_J^{\text{cut}}, \mathcal{T}^{\text{cut}}) = \int_0^{\mathcal{T}^{\text{cut}}} d\mathcal{T}_a \int_0^{\mathcal{T}^{\text{cut}}} d\mathcal{T}_b \int_0^{m_J^{\text{cut}2}/Q_J} d\mathcal{T}_J \frac{d^3\sigma}{d\mathcal{T}_a d\mathcal{T}_b d\mathcal{T}_J}. \quad (18)$$

The differential jet mass cross section, $d\sigma/dm_J$, is obtained by taking the numerical derivative of this cumulant cross section. We define the normalized jet mass

	matching	γ_x	Γ_{cusp}	β	PDF
LL	0-loop	-	1-loop	1-loop	NLO
NLL	0-loop	1-loop	2-loop	2-loop	NLO
NNLL	1-loop	2-loop	3-loop	3-loop	NLO

TABLE II: Perturbative ingredients at different orders in resummed perturbation theory.

spectrum over the range $[0, m_J^{\text{cut}}]$ as $d\hat{\sigma}/dm_J$, so

$$\frac{d\hat{\sigma}}{dm_J}(m_J^{\text{cut}}, \mathcal{T}^{\text{cut}}) \equiv \frac{1}{\sigma(m_J^{\text{cut}}, \mathcal{T}^{\text{cut}})} \frac{d\sigma(\mathcal{T}^{\text{cut}})}{dm_J}. \quad (19)$$

The ingredients in the resummed cross section are needed at different orders in perturbation theory, as summarized in Table II, where the columns correspond to the fixed-order matching, non-cusp anomalous dimension γ_x , cusp anomalous dimension Γ_{cusp} , the β function, and the PDFs. All ingredients necessary for a NNLL resummation of the global logarithms are known and are collected in App. A: The one-loop hard function for the three basic processes $gg \rightarrow Hg$, $gq \rightarrow Hq$, and $q\bar{q} \rightarrow Hg$ via gluon fusion (in the large m_t limit) are obtained from the one-loop helicity amplitudes calculated in Ref. [56] following the procedure in Ref. [57]. The one-loop quark and gluon jet function were calculated in Refs. [58–60], the one-loop quark and gluon beam functions in Refs. [61–64], and the one-loop soft function in Ref. [47]. We also require the cusp anomalous dimension to three loops [65, 66], and the non-cusp anomalous dimensions to two loops, which are known from Refs. [60, 62, 67–70].

There is some freedom in how to treat products of the fixed-order corrections in Eq. (17), specifically the higher-order cross terms that are generated, such as the one-loop correction to H times the one-loop correction to J , which we denote $H^{(1)}J^{(1)}$. The series for the individual objects are fairly convergent, except for the hard function whose one-loop correction is known to be rather large.⁴ We therefore expand the convolutions of the fixed-order B , J , and S functions order by order in α_s to the order needed, but keep the hard function $H^{(0)} + H^{(1)}$ as an overall multiplicative factor. We anticipate that the cross terms involving $H^{(1)}$ will account for a large portion of the NNLO cross section. Schematically, this means that the fixed-order components of our cross section take the form

$$\begin{aligned} & (H^{(0)} + H^{(1)}) \left[(B^{(0)} B^{(0)} J^{(0)}) \otimes S^{(0)} \right. \\ & + (B^{(1)} B^{(0)} J^{(0)}) \otimes S^{(0)} + (B^{(0)} B^{(1)} J^{(0)}) \otimes S^{(0)} \\ & \left. + (B^{(0)} B^{(0)} J^{(1)}) \otimes S^{(0)} + (B^{(0)} B^{(0)} J^{(0)}) \otimes S^{(1)} \right]. \end{aligned} \quad (20)$$

³ Note that we do not call Eq. (17) a factorization theorem since the decoupling of Glauber gluons for hadron collider processes with a specific number of jets has not been proven.

⁴ We find that the improvement of using a complex scale μ_H for the hard function is only marginal, because some logarithms are minimized by a negative μ_H^2 and other by a positive μ_H^2 .

The difference between expanding the hard function or treating it as multiplicative is within our perturbative uncertainty, being a $\lesssim 20\%$ effect for the unnormalized m_J spectrum, and only a $\lesssim 2\%$ effect for the normalized m_J spectrum.

B. Refactorization of the Soft Function

For a process with one or more jets there are multiple directions for collinear radiation and various kinematic variables so a few additional hierarchies become possible. The factorization formula assumes that there are no additional strong hierarchies beyond the collinearity of the jet $m_J^2 \ll p_T^2$, and the absence of additional central jets away from the beam directions, namely $Q_a \mathcal{T}_a \ll p_T^2$ and $Q_b \mathcal{T}_b \ll p_T^2$. Physically, this corresponds to the following four assumptions

- 1) $Q_i \mathcal{T}_i \sim Q_j \mathcal{T}_j$ commensurate m_J and jet veto
- 2) $\frac{q_i \cdot q_j}{E_i E_j} \sim \frac{q_i \cdot q_k}{E_i E_k}$ well separated jet and beams
- 3) $E_i \sim E_j$ jet and beam-jets of similar energy
- 4) $Q_i \sim Q_j$ jet and beam regions of similar size

Assumption 1) ensures that we are in the region where NGLs are not large logarithms. Assumption 2) implies that the jet is not too close to the beam direction, and avoids having large angular logarithms, which would require an additional “ninja summation” [3].

Three combinations of these four assumptions are necessary to avoid introducing additional large logarithms that are not summed by the renormalization group evolution of terms in the factorization formula, namely

$$\frac{s_{ij}}{s_{ik}} \sim 1, \quad \frac{\mathcal{T}_i}{\mathcal{T}_j} \sim 1, \quad \frac{Q_i}{Q_j} \sim 1. \quad (21)$$

The first implies that the logarithms in the hard function can be minimized with a common scale μ , and all three combine to imply that a common scale also minimizes all logarithms in the soft function. One combination of assumptions, $E_i/Q_i \sim E_j/Q_j$, does not appear explicitly in arguments of functions in the factorized cross section, and hence does not show up in logarithms for the leading power result. However, it is in general necessary as part of the derivation of Eq. (17) to ensure that certain neglected terms are power suppressed.

An important consideration in carrying out the summation of large logarithms is the order in α_s and logarithms at which violations of Eq. (21) first become apparent. For the soft function the first terms that appear for the various logarithms are

$$\alpha_s \ln^2 \left(\frac{Q_i \mathcal{T}_i^c Q_j \mathcal{T}_j^c}{\mu^2 s_{ij}} \right), \quad \alpha_s \ln \left(\frac{s_{ij} Q_k}{s_{ik} Q_j} \right), \quad \alpha_s^2 \ln^2 \left(\frac{\mathcal{T}_i^c}{\mathcal{T}_j^c} \right), \quad (22)$$

where we integrate the soft function over \mathcal{T}_i up to the cumulant variable \mathcal{T}_i^c . The first of these is part of the LL series. The second is an angular logarithm. It is part of the NLL series if it counts as a large logarithm. Otherwise, it is part of the $\sim \alpha_s$ fixed-order terms that start to contribute at NNLL. The third is a NGL. It is part of the NLL series if it is a large logarithm. Otherwise it is part of the $\sim \alpha_s^2$ fixed-order terms that start to contribute at N³LL. Therefore, there is a nontrivial constraint on the choice of scales μ in the soft function. The scales must be chosen to minimize the first type of logarithm in Eq. (22) *without* inducing terms of the form of the second and third types already at LL order. In particular, this implies that a poor scale choice could introduce unphysical angular logarithms or NGLs into the LL series. For our choice of kinematics and Q_i the second type of angular logarithm in Eq. (22) is never large. However, since we are exploring a spectrum in $m_J^2 = Q_J \mathcal{T}_J$ the third term in Eq. (22) will grow as the parameters are varied. To surmount this problem requires a refactorization of the soft function which we will consider below.

For the hard function the series of leading double logarithms involves terms of the form

$$\alpha_s \ln^2 \left(\frac{\mu^2}{s_{ij}} \right), \quad \alpha_s \ln^2 \left(\frac{s_{ij}}{s_{ik}} \right). \quad (23)$$

For the choice of jet kinematics explored in this paper we will always satisfy the assumption $s_{ij} \sim s_{ik}$, so there is no additional constraint on the scale associated with the hard function.

The hierarchy between \mathcal{T}_J and \mathcal{T}^{cut} leads to unphysical large logarithms if a single scale μ_S is used for the initial conditions for the soft function evolution. Here we address how these can be removed by a refactorization of the soft function, with corrections from the true higher order non-global logarithms (see Refs. [8, 12, 15, 37] for earlier refactorization discussions).

In general, the all-order soft function has the form

$$S(\{k_i\}, \{\hat{q}_i^\mu\}, \mu) = \prod_i S_i(k_i, \{\hat{q}_i^\mu\}, \mu) + S_{\text{NGL}}(\{k_i\}, \{\hat{q}_i^\mu\}, \mu), \quad (24)$$

where $\hat{q}_i^\mu = q_i^\mu/Q_i$. Here S_{NGL} contains all non-global terms, and hence has an intrinsic dependence on the ratios k_i/k_j . At NLO there is only one soft gluon emitted, which can contribute to only one of the \mathcal{T}_i at a time. Thus the NLO soft function factorizes, and $S_{\text{NGL}} \sim \mathcal{O}(\alpha_s^2)$. Truncating to $\mathcal{O}(\alpha_s)$ there is still some freedom in the definition of the S_i . Whereas the terms with explicit k_i dependence in $S(\{k_i\}, \mu)$ clearly belong to $S_i(k_i, \mu)$, the pure delta function terms $\delta(k_J)\delta(k_a)\delta(k_b)$ can in principle be split in multiple ways between the various $S_i(k_i, \mu)$. We choose to split these terms evenly, as detailed in App. A 4, and we introduce an additional parameter r in the scale variation to estimate uncertainty from this freedom as discussed further below and in detail in Sec. III C.

Due to the consistency of the factorization formula, the evolution of the soft function factorizes exactly to all orders in perturbation theory,

$$\begin{aligned} U_S(\{k_i\}, \mu, \mu_0) &= U_H(\mu_0, \mu) \prod_i Q_i U_{J_i}(Q_i k_i, \mu_0, \mu) \\ &= \prod_i U_{S_i}(k_i, \mu, \mu_0), \end{aligned} \quad (25)$$

(Here we used the fact that the beam and jet functions have the same evolution [62].) This involves the factorization of the evolution of the hard function $H = CC^\dagger$, which follows from the form of the anomalous dimension for C [71, 72],

$$\begin{aligned} \widehat{\gamma}_C(\mu) &= \Gamma_{\text{cusp}}[\alpha_s(\mu)] \left[\sum_i \mathbf{T}_i^2 \ln \frac{\mu}{\mu_0} \right. \\ &\quad \left. + \sum_{i < j} \mathbf{T}_i \cdot \mathbf{T}_j \ln \left(-\frac{s_{ij}}{\mu_0^2} - i0 \right) \right] + \widehat{\gamma}_C[\alpha_s(\mu)]. \end{aligned} \quad (26)$$

The sum on i and j runs over the colored partons participating in the short-distance interaction and \mathbf{T}_i denotes the corresponding color charge matrix. (For $pp \rightarrow H + 1$ jet the color space is still trivial, so these color matrices are just numbers.) To associate the $\ln \mu$ terms to individual partons we introduced a dummy variable μ_0 and used color conservation. It is not a priori clear how to associate the remaining terms within the $\sum_{i < j}$ to each U_{S_i} , and we choose to split each term evenly between i and j . The explicit expression for the factorized hard function evolution that we employ is given in App. A 5. Other potential

choices of splitting up these terms are again probed by the scale parameter r , which is discussed in more detail around Eq. (31), and the corresponding uncertainty is found to be small. The two-loop non-cusp anomalous dimension has the structure $\widehat{\gamma}_C(\alpha_s) = n_q \gamma_q + n_g \gamma_g$, where n_g and n_q are the number of gluon and (anti)quark legs, so it naturally factors.

The factorization of the evolution and fixed-order soft function in Eqs. (24) and (25) suggests that we can evaluate the piece of the soft function corresponding to \mathcal{T}_i at a scale μ_{S_i} ,

$$S(\{k_i\}, \mu) = \prod_i \int dk'_i U_{S_i}(k_i - k'_i, \mu, \mu_{S_i}) S_i(\{k'_i\}, \mu_{S_i}). \quad (27)$$

This factorization does not hold for all the terms at order α_s^2 , since there are diagrams that contribute to multiple \mathcal{T}_i , leading to non-global logarithms of the form $\alpha_s^2 \ln^2(k_i^c/k_j^c)$ appearing in S_{NGL} in Eq. (24). We discuss in Sec. III D how we estimate the size of these NGL contributions in the jet mass spectrum.

In our implementation we find it simplest to run the hard, jet, and beam functions, rather than the soft function, as summarized in Fig. 3, though the final results are completely independent of this choice. Since the cut on both beams is the same, they have a common μ_B and μ_{S_B} . We summarize the work in this section by presenting the factorization formula valid at NNLL which includes the evolution factors and refactorization of S ,

$$\begin{aligned} \frac{d^3 \sigma^{H+1j}(\mathcal{T}^{\text{cut}})}{d\eta_J dp_T^J dY dm_J} &= \frac{(2p_T^J m_J / Q_J)}{4\pi E_{\text{cm}}^2 (Q^2 + m_H^2)} \sum_{\kappa} H_{\kappa}(\{q_i^\mu\}, \mu_H) U_{H_{\kappa_a}}(\{q_i^\mu\}, \mu_{S_B}, \mu_H) U_{H_{\kappa_b}}(\{q_i^\mu\}, \mu_{S_B}, \mu_H) U_{H_{\kappa_J}}(\{q_i^\mu\}, \mu_{S_J}, \mu_H) \\ &\quad \times \int dt_a dt'_a U_{J_{\kappa_a}}(t_a - t'_a, \mu_{S_B}, \mu_B) B_{\kappa_a}(t'_a, x_a, \mu_B) \int dt_b dt'_b U_{J_{\kappa_b}}(t_b - t'_b, \mu_{S_B}, \mu_B) B_{\kappa_b}(t'_b, x_b, \mu_B) \\ &\quad \times \int ds_J ds'_J U_{J_{\kappa_J}}(s_J - s'_J, \mu_{S_J}, \mu_J) J_{\kappa_J}(s'_J, \mu_J) \int_0^{\mathcal{T}^{\text{cut}}} d\mathcal{T}_a S_a \left(\mathcal{T}_a - \frac{t_a}{Q_a}, \left\{ \frac{q_i^\mu}{Q_i} \right\}, \mu_{S_B} \right) \\ &\quad \times \int_0^{\mathcal{T}^{\text{cut}}} d\mathcal{T}_b S_b \left(\mathcal{T}_b - \frac{t_b}{Q_b}, \left\{ \frac{q_i^\mu}{Q_i} \right\}, \mu_{S_B} \right) S_J \left(\frac{m_J^2 - s_J}{Q_J}, \left\{ \frac{q_i^\mu}{Q_i} \right\}, \mu_{S_J} \right). \end{aligned} \quad (28)$$

All necessary perturbative results for H_{κ} , J_{κ_J} , S_i , and the U_i are collected in App. A.

C. Choice of Running Scales

The factorization formula in Eq. (28) sums the large logarithms of $Q_i \mathcal{T}_i / p_T^{J^2}$ from the cuts on the beams and of $Q_J \mathcal{T}_J / p_T^{J^2} = m_J^2 / p_T^{J^2}$ from the jet mass measurement

by carrying out perturbation theory for the hard, beam, jet, and soft functions at their natural scales

$$\mu_H \simeq p_T^J, \quad \mu_i \simeq \sqrt{Q_i \mathcal{T}_i}, \quad \mu_{S_i} \simeq Q_i \mathcal{T}_i / p_T^J, \quad (29)$$

and then running them to an arbitrary common scale. Here $i = a, b, J$, since the situation for the beams and the jet are fully analogous. As discussed above we must use the combination $Q_i \mathcal{T}_i$ to ensure we have the correct leading logarithms. The remaining dependence of the cross

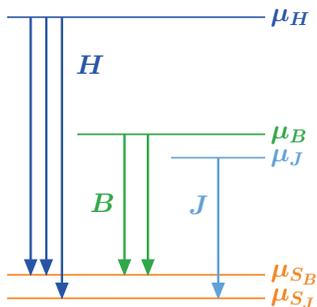


FIG. 3: Illustration of the different fixed-order scales appearing in the factorized cross section and our evolution strategy.

section on Q_i occurs due to its impact on the boundaries between the jet and beam regions. This is encoded in the fixed-order terms in the soft function that do not involve large logarithms.

If any \mathcal{T}_i becomes very small, $\sim \Lambda_{\text{QCD}}$, the nonperturbative corrections to the soft function become important and the scales are $\mu_H \simeq p_T^J$, $\mu_i \sim \sqrt{\Lambda_{\text{QCD}} p_T^J}$, and $\mu_{S_i} \gtrsim \Lambda_{\text{QCD}}$. On the other hand, for large $\mathcal{T}_i \sim p_T^J$ the resummation of $Q_i \mathcal{T}_i / p_T^{J2}$ is not important and is turned off by having the scales merge, $\mu_H = \mu_i = \mu_{S_i} \simeq p_T^J$.

To connect the various regions where the resummation must be handled differently, we use \mathcal{T}_i -dependent scales, which are known as profile functions [23, 73]. A transition between these three regions is given by the following running scales, adopted from Ref. [64],

$$\begin{aligned} \mu_H &= \mu, \\ \mu_i(\tau) &= \left[1 + e_i \theta(\tau_3 - \tau) \left(1 - \frac{\tau}{\tau_3} \right)^2 \right] \sqrt{\mu \mu_{\text{run}}(\tau, \mu)}, \\ \mu_{S_i}(\tau) &= \left[1 + e_{S_i} \theta(\tau_3 - \tau) \left(1 - \frac{\tau}{\tau_3} \right)^2 \right] \mu_{\text{run}}(\tau, \mu), \end{aligned} \quad (30)$$

where $\tau = Q_i \mathcal{T}_i / Q^2$. The expression for μ_{run} can be found in App. B, along with the choice of parameters for the central values, and details on the variations of μ , e_i , and e_{S_i} used to estimate the perturbative uncertainties in our predictions.

To estimate the additional perturbative uncertainty associated with the refactorization of the soft function in Sec. III B, we reintroduce correlations between the soft scales using a parameter r satisfying $0 \leq r \leq 1$,

$$\begin{aligned} \mu_{S_J}^{(r)} &= (\bar{\mu}_S)^r (\mu_{S_J})^{1-r}, \quad \mu_{S_B}^{(r)} = (\bar{\mu}_S)^r (\mu_{S_B})^{1-r}, \\ \ln \bar{\mu}_S &\equiv \frac{(T_a^2 + T_b^2) \ln \mu_{S_B} + T_J^2 \ln \mu_{S_J}}{T_a^2 + T_b^2 + T_J^2}. \end{aligned} \quad (31)$$

Here $T_i^2 = C_F$ for $i = q$ and $i = \bar{q}$, and $T_i^2 = C_A$ for $i = g$. For $r = 0$ we have the original uncorrelated soft scales. By increasing r the scales move towards the ‘‘color average’’ value $\bar{\mu}_S$. At $r = 1$ they are all equal to this average soft scale, so the refactorization is turned off (which as explained earlier causes unphysical NGLs in the LL series). To estimate the size of the freedom in the

refactorization we take $r = 0.2$ as our default choice and include $r = 0$ and $r = 0.4$ as separate scale variations in our uncertainty estimate.

D. Non-Global Logarithms

If the NGLs are not large logarithms then they enter beyond NNLL order, and should be of comparable size to other higher-order perturbative terms. This is obviously only possible for some range of $m_J^2 / (p_T^J \mathcal{T}^{\text{cut}})$, which determines where our result is valid at NNLL order. To determine this range we include the leading $\mathcal{O}(\alpha_s^2)$ NGL into our resummed calculation and compare the results with and without this term for various parameter choices. In the factorized cross section all NGLs enter through the soft function S . For simplicity we restrict this analysis of the size of non-global logarithms to the $gg \rightarrow Hg$ channel, as the results for other channels are similar.

The leading NGL in the cumulant soft function is

$$\begin{aligned} S_{\text{NGL}}(\{k_i^c\}, \mu_S) &= \prod_i \left(\int_0^{k_i^c} dk_i \right) S_{\text{NGL}}(\{k_i\}, \mu_S) \\ &= -\frac{\alpha_s^2(\mu_S) C_A^2}{(2\pi)^2} \sum_{i < j} G_{ij} \ln^2 \left(\frac{k_i^c}{k_j^c} \right). \end{aligned} \quad (32)$$

Here G_{ij} is a geometric factor that depends on the boundaries of the jet and beam regions. Note the absence of explicit μ -dependence in the NGLs. These expressions for S_{NGL} follow from the known result for $e^+e^- \rightarrow 2$ jets [32, 33, 41, 42], by replacing the color factor $C_F C_A \rightarrow C_A^2$. Unlike the global logarithms this contribution does not factor, so we assign it a common soft scale which we take to be $\bar{\mu}_S$ given in Eq. (31).

For the purpose of our numerical analysis we take $G_{ij} = \pi^2/3$, which is the result for a hemisphere. This may be thought of as reasonable estimate and in reality the values may differ by about 15% to 30% [14]. Converting the cumulant space result in Eq. (32) into a full distribution yields

$$\begin{aligned} S_{\text{NGL}}(\{k_i\}, \bar{\mu}_S) &\simeq -\frac{\alpha_s^2(\bar{\mu}_S) C_A^2 \pi^2}{(2\pi)^2} \frac{1}{3} \left[\sum_i \frac{4}{\mu'} \mathcal{L}_1 \left(\frac{k_i}{\mu'} \right) \right. \\ &\quad \left. - 2 \sum_{i < j} \frac{1}{\mu'} \mathcal{L}_0 \left(\frac{k_i}{\mu'} \right) \frac{1}{\mu'} \mathcal{L}_0 \left(\frac{k_j}{\mu'} \right) \right], \end{aligned} \quad (33)$$

where the \mathcal{L}_n denote standard plus distributions as defined in Eq. (A5). Note that the μ' dependence cancels out explicitly between the terms, so the choice of this scale is arbitrary and irrelevant. It is introduced for coding purposes, since it is convenient to have the same type of \mathcal{L}_n distributions as in the non-NGL part of the soft function. When the NGLs are included in this manner, via the soft function in the factorization, one automatically resums an infinite series of global logarithms that multiply the NGL. In particular, this includes terms that are schematically $[\alpha_s^2 \ln^2] [\sum_k (\alpha_s \ln^2)^k]$ where the first \ln^2

is non-global and the second \ln^2 is a large global logarithm. The all-order structure of this series of terms is correctly predicted by the factorization.

For our analysis we will mostly be interested in the normalized spectrum in Eq. (19). Here in the numerator the two jet veto variables are in cumulant space and m_J is in distribution space, while in the denominator all the variables are in cumulant space. This result has two types of NGLs

$$\begin{aligned} \text{i)} \quad & \alpha_s^2(\mu_S) \ln^2 \left(\frac{m_J^{\text{cut}2}}{p_T^J \mathcal{T}^{\text{cut}}} \right), \\ \text{ii)} \quad & \alpha_s^2(\mu_S) \frac{2}{\mathcal{T}^{\text{cut}}} \mathcal{L}_1 \left(\frac{m_J^2}{p_T^J \mathcal{T}^{\text{cut}}} \right). \end{aligned} \quad (34)$$

For the denominator the relevant form of the NGL logarithms is as in Eq. (32), yielding the terms i). For the numerator the form of the NGL is as in ii). The presence of two types of NGLs in the normalized spectrum implies a somewhat different dependence than for the unnormalized cross section. The effect of NGLs in these two cases are analyzed in detail in Sec. IV C. There we will show that there is indeed a fairly large range of m_J^{cut} values where the NGL terms in the exclusive jet cross section are not large logarithms.

IV. RESULTS FOR GLUON AND QUARK JETS

In this section we focus on the individual quark and gluon channels, leaving results for $pp \rightarrow H + 1$ jet to be discussed in Sec. V below. We first study the theoretical predictions for the m_J spectrum with and without normalization, and show that normalizing substantially reduces the perturbative uncertainty. We also study the order-by-order convergence of this differential cross section, and the size of various contributions to the perturbative uncertainty bands. Next, the dependence on the jet veto \mathcal{T}^{cut} is studied. Finally, we investigate the size of non-global logarithms as a function of m_J and \mathcal{T}^{cut} .

A. Default Parameter Choices

Unless indicated otherwise we use the following default parameter choices for all plots in Secs. IV, V, and VI. For the Higgs mass we take $m_H = 125$ GeV [74, 75], and for the LHC center-of-mass energy we take $E_{\text{cm}} = 7$ TeV. We always use the MSTW NLO PDFs [76] with the corresponding value of $\alpha_s(m_Z) = 0.1202$ for the strong coupling constant. As our default we use the geometric $R = 1$ measure for defining the jets, $\mathcal{T}^{\text{cut}} = 25$ GeV for the jet veto, and $m_J^{\text{cut}} = 200$ GeV for the normalization range. Our default hard kinematics are $p_T^J = 300$ GeV, $\eta_J = 0$, and $Y = 0$. Finally, for the scale functions μ_H , $\mu_B(\tau)$, $\mu_J(\tau)$, and $\mu_{S_i}(\tau)$ defined in Sec. III C, the central parameter values are given in App. B. There we

also discuss the combination of scale variations used for estimating the perturbative uncertainties.

B. Normalization and Convergence

The unnormalized jet mass spectrum at NNLL with our default inputs for the quark and gluon channels are shown in Fig. 4(a). As one expects, the gluon jets peak at a much higher jet mass than the quark jets. We also see that the perturbative uncertainties are quite sizable, even at NNLL.

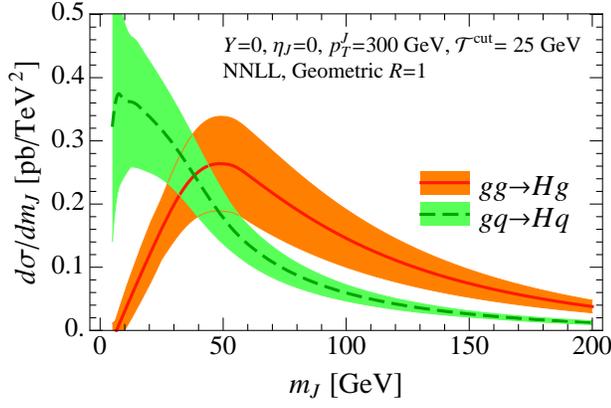
Normalizing the jet mass spectrum allows one to study its shape without contamination from the slow convergence of the integrated 1-jet cross section, and also reduces the experimental uncertainties significantly. We denote the normalized cross section as $d\hat{\sigma}/dm_J$ and calculate it using Eq. (19) where we normalize over the range $0 \leq m_J \leq m_J^{\text{cut}}$.

We first study the impact of normalization on the perturbative uncertainty. To preserve the normalization, we simultaneously vary the scales in the numerator and denominator of Eq. (19). Comparing the unnormalized cross section at NNLL for the gluon and quark channels shown in Fig. 4(a) to the normalized ones in Fig. 4(b), we observe that a substantial portion of the uncertainty is related to the integrated cross section rather than the shape. In the peak region of the m_J spectrum, $30 \text{ GeV} \leq m_J \leq 150 \text{ GeV}$ the normalized cross sections have a quite reasonable remaining perturbative uncertainty of $\simeq 5$ –10%.

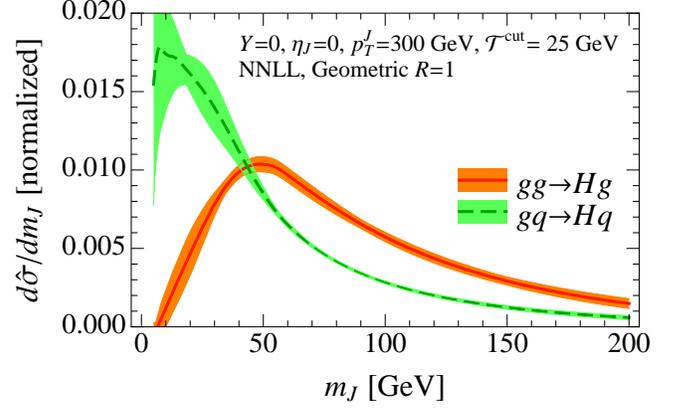
A big part of the sizable uncertainty in the unnormalized 1-jet cross section is due to the poor convergence of the hard function for $pp \rightarrow H + 1$ jet, and thus specific to the Higgs process. By keeping the hard function as an overall multiplicative factor as in Eq. (20), it cancels exactly in the normalized cross section for a given partonic channel and fixed phase space point (which we use for most of our plots). This cancellation still takes place approximately for the integrated cross section summed over partonic channels as we show below in Sec. V B. Our results with fixed kinematics are therefore representative of results integrated over the jet phase space.

The order-by-order convergence of our resummed jet mass calculation is displayed in Figs. 4(c) and 4(d) for the gluon and quark jet channels, where results at LL, NLL, and NNLL are shown. The various bands overlap with those of lower orders, providing direct evidence that our scale variations yield a reasonable estimate of the higher-order perturbative uncertainties.

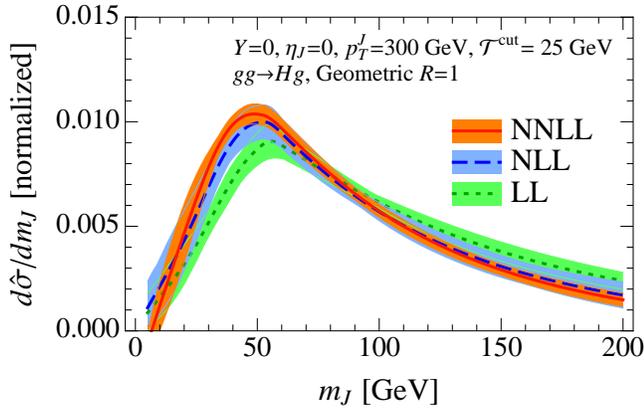
There are several classes of perturbative scale uncertainties, the ‘‘Fixed Order’’ scale variation that is correlated with the total cross section, the ‘‘Beam’’ scale variation from varying μ_B and μ_{S_B} that is related to the presence of the jet veto, the ‘‘Jet’’ scale variation from varying μ_J and μ_{S_J} that is related to the jet mass measurement, and the uncertainty from ‘‘r’’ that is related to the perturbative freedom in the refactorized formula



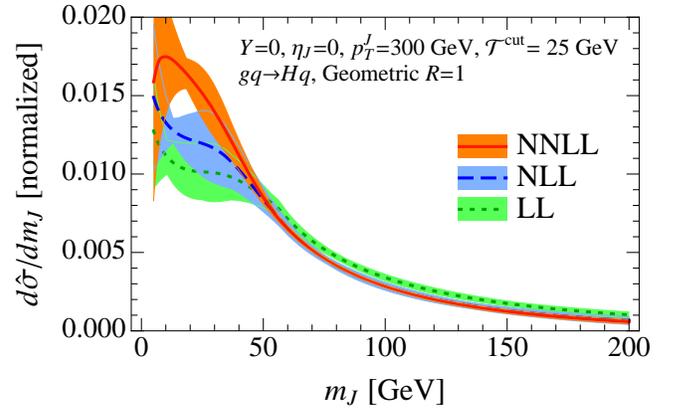
(a) Unnormalized jet mass spectrum for quark and gluon jets at NNLL. The uncertainties are sizable even at NNLL.



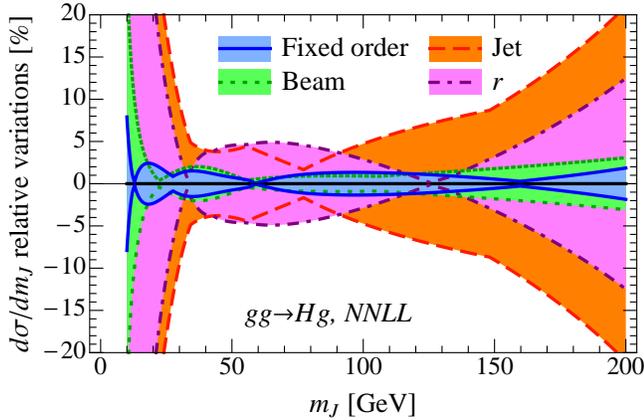
(b) Normalized jet mass spectrum for quark and gluon jets at NNLL. Compared to Fig. 4(a), the normalization significantly reduces the perturbative uncertainties.



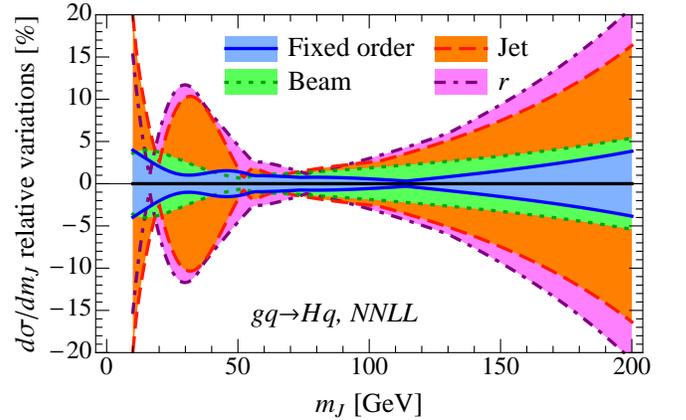
(c) Convergence of the resummed calculation for gluon jets.



(d) Convergence of the resummed calculation for quark jets.



(e) Individual scale variations that enter the uncertainty estimate for gluon jets at NNLL. Shown are the variations relative to the central NNLL curve.



(f) Individual scale variations that enter the uncertainty estimate for quark jets at NNLL. Shown are the variations relative to the central NNLL curve.

FIG. 4: Perturbative uncertainties and convergence for the jet mass spectrum in $gg \rightarrow Hg$ and $gq \rightarrow Hq$ with default inputs.

for the soft function. For the NNLL results, these individual scale variations are shown in Figs. 4(e) and 4(f) for gluon and quark jets respectively. For simplicity we combined the uncertainty from varying the jet scale μ_J

and the scale of the jet part of the soft function μ_{S_J} by taking the envelope, and similarly for the beams. It is not too surprising that the uncertainties associated with the hard and beam scale variations are smaller, since they

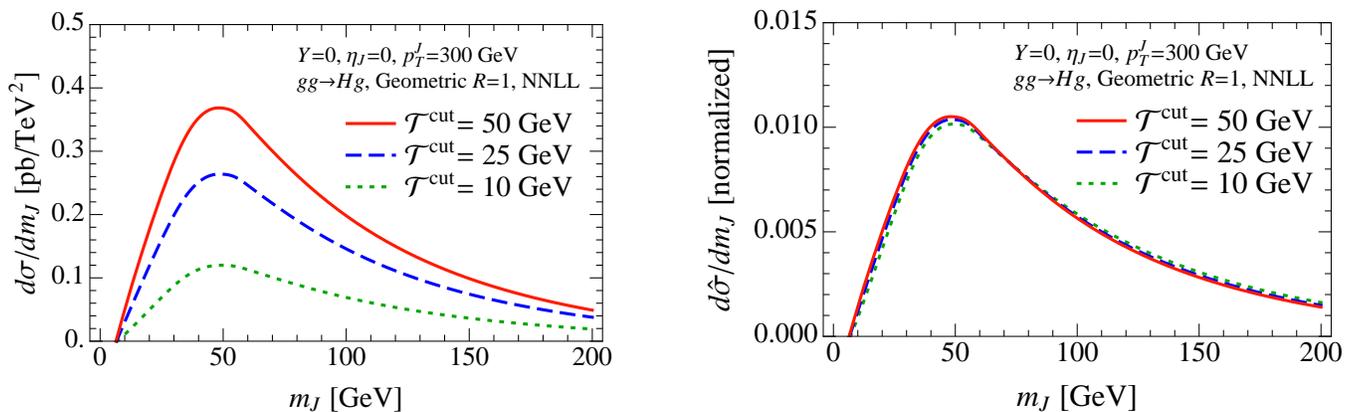


FIG. 5: Effect of using different jet veto cuts on the jet mass spectrum for $gg \rightarrow Hg$. While the unnormalized spectrum on the left is directly sensitive to the jet veto cut, this dependence almost completely cancels in the normalized spectrum on the right. The same is true for the quark channel, $qq \rightarrow Hq$, and the sum over all partonic channels.

are mostly common to the numerator and denominator of the normalized spectrum in Eq. (19). To obtain the total perturbative uncertainty we take the envelope of “Jet”, “Beam” and “r” uncertainties and combine it in quadrature with the “Fixed Order” uncertainty. The total uncertainty in the jet mass spectrum is dominated by that of the jet and by the soft function refactorization.

C. Jet Veto and Non-Global Logarithms

Next we discuss the effect of the jet veto on the jet mass spectrum. Our veto is imposed through the variable \mathcal{T}^{cut} , rather than a more traditional $p_{T,J}^{\text{cut}}$, since this simplifies the treatment of scales in the problem, and allows us to make use of a known factorization theorem. This jet veto restricts the initial and final-state collinear radiation as well as soft radiation. It turns out that the *normalized* jet mass spectrum is fairly insensitive to the value of the jet veto cut.

We start by considering the effect of the jet veto on the *unnormalized* jet mass spectrum, as shown for $gg \rightarrow Hg$ in the left panel of Fig. 5. Decreasing \mathcal{T}^{cut} imposes a stronger restriction on the initial-state radiation and reduces the unnormalized cross section. (This reduction is less strong for $qq \rightarrow Hq$, because quarks radiate less than gluons.) As the right panel of Fig. 5 shows, the normalization removes the majority of the \mathcal{T}^{cut} dependence. Note that without the refactorization of the soft function (see Sec. III B) this cancellation would be spoiled by unphysical logarithms. This strong cancellation is also the case for the other partonic channels, as well as for their sum in $pp \rightarrow H + 1$ jet. This insensitivity to \mathcal{T}^{cut} also remains valid after integrating over the jet phase space, as we show below in Fig. 8. We have also studied the dependence on \mathcal{T}^{cut} as well as a standard $p_{T,J}^{\text{cut}}$ jet veto with PYTHIA, where we also find a similar insensitivity of the normalized jet mass spectrum to the details of the

used jet-veto variable and cut values.

Next we turn to our analysis of NGLs, both in the unnormalized and normalized jet mass spectra. As explained in Sec. III D, we test for the size of the NGLs by comparing the cross section with and without these terms. The leading NGL is included in fixed-order perturbation theory, on top of which we sum an infinite series of global logarithms through the factorization formula.

In the left panel of Fig. 6 we show the unnormalized spectrum for various \mathcal{T}^{cut} values at NNLL (solid lines) and the same spectra including the NGL terms (dotted lines). As mentioned earlier, there is a point on the spectrum where the NGLs exactly cancel. This point is at $m_J \simeq 50, 110, 165, 300$ for $\mathcal{T}^{\text{cut}} = 10, 25, 50, 150$ GeV respectively. For all values of m_J shown in this figure the effect of the NGL terms is well within the perturbative uncertainty [cf. the uncertainty bands shown in Fig. 4(a)].

When we normalize the spectrum we are dividing by the cumulant with m_J^{cut} , and the jet-veto dependence does not cancel out in the presence of the non-global logarithms. There are two types of NGLs in the normalized result, terms involving $\ln[m_J^2/(p_T^J \mathcal{T}^{\text{cut}})]$ from the numerator and terms involving $\ln[m_J^{\text{cut}2}/(p_T^J \mathcal{T}^{\text{cut}})]$ from the denominator. Therefore for a fixed \mathcal{T}^{cut} there is no longer a value of m_J where all the NGLs will vanish. Results for the normalized spectrum with and without NGLs are shown in the right panel of Fig. 6. The orange band shows the NNLL result without NGLs along with its perturbative uncertainty, while the various black lines show the central values for NNLL results that have the NGLs included. For the wide range of values $25 \text{ GeV} \leq \mathcal{T}^{\text{cut}} \leq 150 \text{ GeV}$ the effect of the NGLs is of the same size as the reduced perturbative uncertainty in the normalized spectrum. This justifies our assertion that the NGLs do not have to be considered as large logarithms for a significant range of cut values, so that our NNLL result is complete at this order. In the small m_J region of the spectrum the resummation of global logarithms on top of the NGL term provides an appropriate

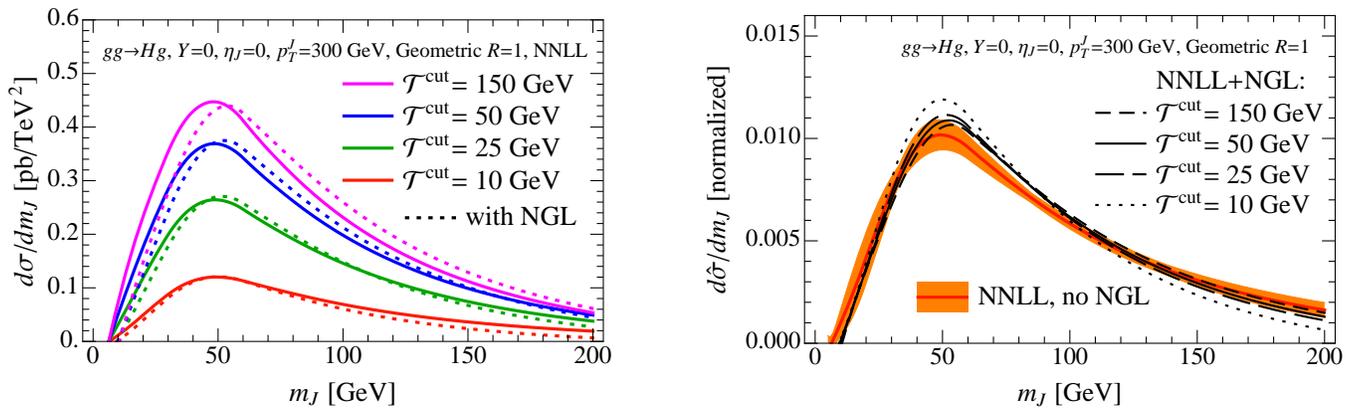


FIG. 6: Effect of non-global logarithms on the NNLL jet mass spectrum for $gg \rightarrow Hg$ for different jet veto cuts. Left panel: Including the leading NGLs (dashed lines) has a small effect on the unnormalized spectrum, and is well within the perturbative uncertainty for a wide range of jet veto cuts. Right panel: The effect of including the leading NGLs (black solid, dashed, and dotted curves) on the normalized NNLL spectrum (orange band) is still within the reduced perturbative uncertainty for a wide range of jet veto cuts, but start to become important for $\mathcal{T}^{\text{cut}} = 10$ GeV.

Sudakov suppression in the the cross section. For other m_J values, and $25 \text{ GeV} \leq \mathcal{T}^{\text{cut}} \leq 150 \text{ GeV}$, the argument of the NGL remains between $1/8$ and 8 , which is the range over which we expect that the NGLs do not dominate over nonlogarithmic corrections, as mentioned in the introduction. On the other hand, for $\mathcal{T}^{\text{cut}} = 10 \text{ GeV}$ one observes that the NGLs become large enough that they are no longer contained within the perturbative uncertainty, so this value is outside the range of validity of our normalized NNLL results (though for this value the unnormalized results in the peak region are still valid). For this value the argument of the NGL involving m_J^{cut} becomes $\simeq 13$, which is outside of the range mentioned above.

Although we have only explored the $gg \rightarrow Hg$ channel at a fixed kinematic point in this section, we have also checked explicitly that the same conclusions about NGLs hold when integrating over a kinematic range, and when considering quark jets from $gq \rightarrow Hq$.

V. RESULTS FOR $pp \rightarrow H + 1$ JET

In this section we show results for the $pp \rightarrow H + 1$ jet cross section at NNLL, summing the contributions from the various partonic channels: $gg \rightarrow Hg$, $gq \rightarrow Hq$, and the (small) $q\bar{q} \rightarrow Hg$. We present results for the dependence of the jet mass spectrum on the jet kinematics, on the choice of jet definition which affects the shape of the jets, and on the jet size R . We also compare the m_J spectrum obtained for a fixed point in the jet kinematics to that obtained from integrating over a range of jet momenta.

A. Dependence on Kinematics

For $pp \rightarrow H + 1$ jet there are three nontrivial kinematic variables: the transverse momentum of the jet p_T^J , rapidity of the jet η_J , and the total rapidity Y of the combined Higgs+jet system. We show how each of these variables affect both the unnormalized and normalized jet mass spectrum, which allows us to separate the impact of kinematics on the normalization and the shape.

The falloff of the PDFs at larger x values causes the cross section to strongly decrease for increasing p_T^J and for increasing $|\eta_J|$ (for $Y = 0$). This is shown in Figs. 7(a) and 7(c). The dependence on p_T^J and η_J in the corresponding normalized spectra are shown in Figs. 7(b) and 7(d). Here we see that there is a decrease in the height of the peak and a compensating increase in the tail height as p_T^J or $|\eta_J|$ are increased. Note that for these variables there is a marked difference between the total $pp \rightarrow H + 1$ jet process compared to the individual partonic channels (which are not shown). For each partonic channel the peak position of the jet mass spectrum increases as $m_J^{\text{peak}} \propto \sqrt{p_T^J}$ and also increases with increasing $|\eta_J|$. However, at the same time the contribution of $gq \rightarrow Hq$ relative to $gg \rightarrow Hg$ is enhanced, and the peak of the jet mass spectrum is at lower values for quark jets than for gluon jets [see Fig. 4(b)]. These two effects largely cancel for $pp \rightarrow H + 1$ jet, such that the peak position is practically unchanged with increasing p_T^J , whereas for increasing η_J a small net increase in the peak position remains.

Note that our ability to calculate the η_J dependence implies that it is trivial to impose rapidity cuts in our framework.

The main dependence on the total system rapidity Y enters through the shape of the PDFs, causing the cross section to strongly decrease with increasing $|Y|$, as Fig. 7(e) shows. (This is also the reason for taking cen-

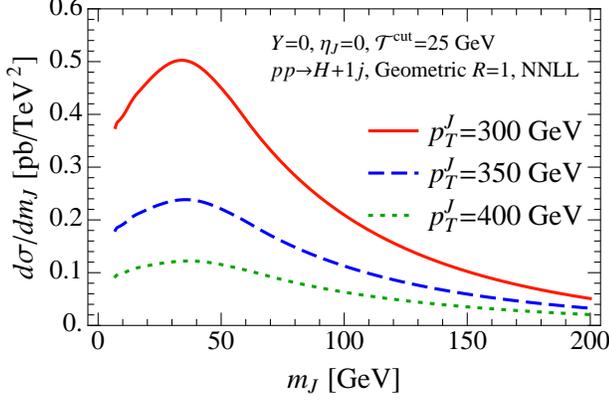
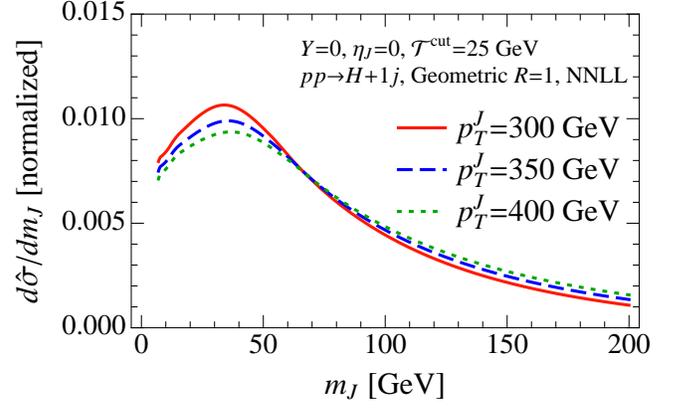
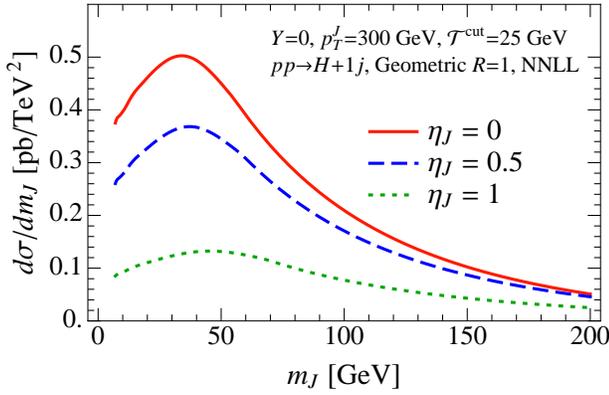
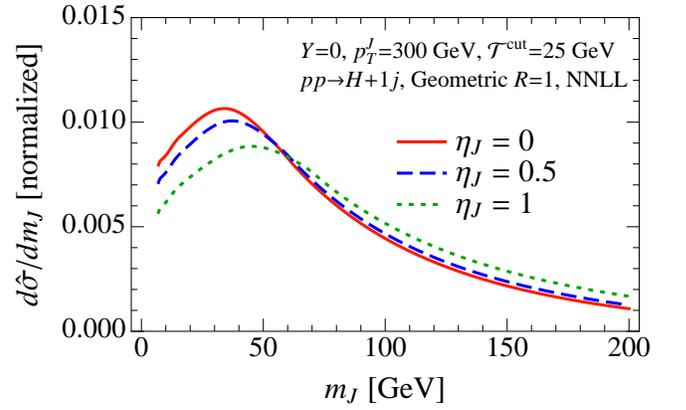
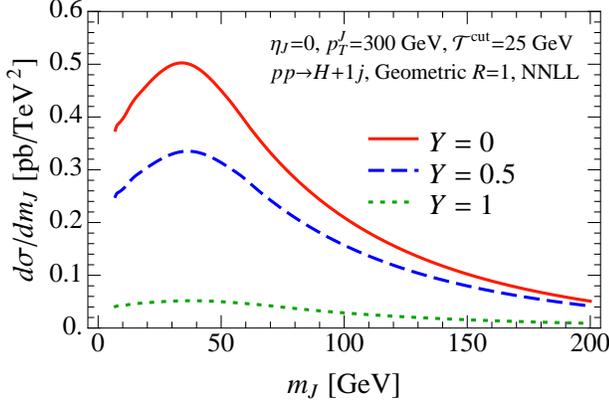
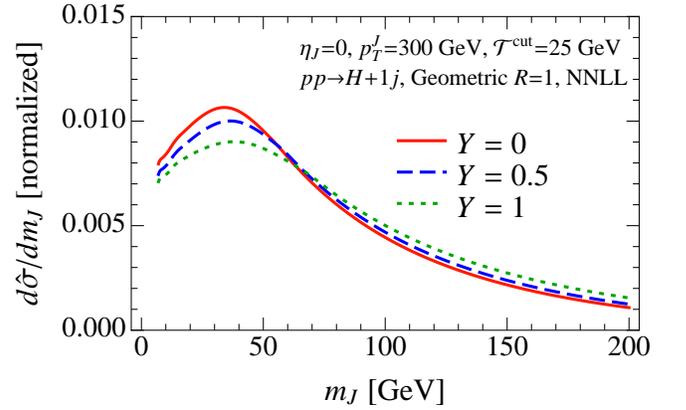
(a) The cross section decreases with increasing p_T^J .(b) For $pp \rightarrow H + 1$ jet the peak position remains stable and the spectrum slightly broadens with increasing p_T^J .(c) The cross section decreases with increasing η_J .(d) For $pp \rightarrow H + 1$ jet the peak position shifts slightly and the spectrum slightly broadens with increasing η_J .(e) The cross section quickly decreases for larger Y .(f) For $pp \rightarrow H + 1$ jet the peak position remains stable and the spectrum slightly broadens with increasing Y .

FIG. 7: Dependence on the kinematic variables p_T^J , η_J , and Y for the unnormalized and normalized NNLL jet mass spectra for $pp \rightarrow H + 1$ jet.

tral jets with $Y = 0$ for our default value when using a single phase space point.) The value of Y also affects the shape of the jet mass spectrum, as can be seen in Fig. 7(f). The jet rapidity relative to the partonic center

of mass is $\eta - Y$, so one would expect the shape change as function of Y to be similar to that as function of η_J , shown in Fig. 7(d). The agreement is close but not exact because the Y dependence induced by the shape of the

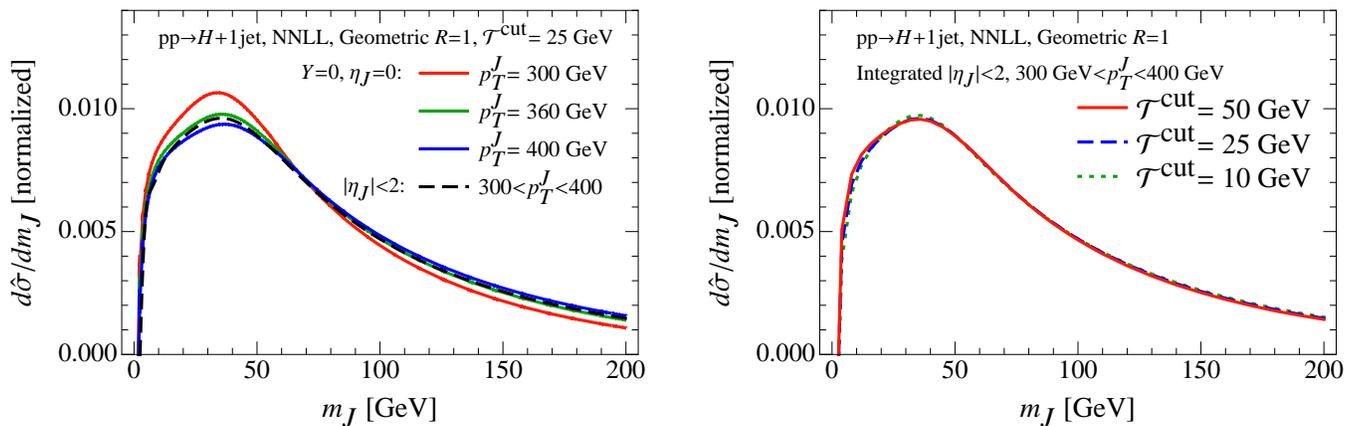


FIG. 8: Results for the normalized jet mass spectrum at NNLL for $pp \rightarrow H + 1 \text{ jet}$ after integrating over $300 \text{ GeV} < p_T^J < 400 \text{ GeV}$, $|\eta_J| < 2$, and all Y . The left panel compares the spectrum for integrated kinematics (dashed line) to those for fixed kinematics with $Y = \eta_J = 0$ and $p_T^J = 300, 360, 400 \text{ GeV}$ (solid lines from top to bottom at the peak). The right panel shows the impact of \mathcal{T}^{cut} on the normalized spectrum for integrated kinematics, which is the analog of the comparison in the right panel of Fig. 5 for fixed kinematics.

PDFs differs channel by channel, and thus affects their relative weight in the sum over channels.

B. Integrated Kinematics

So far we have shown the m_J spectra for a fixed point in p_T^J , η_J , and Y . We now consider the impact of integrating the kinematic variables over a bin with $|\eta_J| < 2$, $300 \text{ GeV} < p_T^J < 400 \text{ GeV}$, and any Y . These kinematic ranges are realistic experimentally for jets at the LHC.

In the left panel of Fig. 8 the jet mass spectrum for integrated kinematics is shown by a black dashed line, and is compared to three spectra with fixed kinematics shown by solid lines (with $Y = \eta_J = 0$ and $p_T^J = 300, 360, 400 \text{ GeV}$ from top to bottom at the peak of the spectrum). One observes that the m_J spectrum in the integrated bin is very close to the m_J spectrum with $Y = \eta_J = 0$ and near the center of the p_T^J bin. Thus our conclusions made from studies of a single kinematic point directly carry over to the results obtained by integrating over a phase space bin.

The one situation where this is not immediately obvious is the dependence of the normalized cross section on the jet-veto cut, \mathcal{T}^{cut} , shown for fixed kinematics in Fig. 5. When we integrate over the kinematic bin the hard function, including its Sudakov form factor depending on \mathcal{T}^{cut} , no longer exactly cancels between the numerator and denominator. Nevertheless, comparing the spectra for integrated kinematics and different values of \mathcal{T}^{cut} , shown in Fig. 8, we see that the normalized spectrum is still very insensitive to the details of the jet veto also after summing over partonic channels and integrating over a range of kinematics. (We have also confirmed that upon phase space integration the size of the NGL effect remains the same as shown in Fig. 6.)

C. Jet Definitions and Jet Area

In Sec. II we discussed the various N -jettiness measures (defined by the Q_i) and illustrated the corresponding size and shape of the jet regions for the geometric cases. An illustration of the more irregular regions that appear for the invariant mass measure can be found in Ref. [47]. We now study how the jet mass spectrum is affected by these different jet definitions as well as by their jet area (R dependence). We start by noting that in the N -jettiness factorization only the soft function is sensitive to the boundaries of the jet regions. Since the soft function only probes the boundaries starting at NLO, the first dependence on the jet definition appears at NNLL. The nontrivial Q_i dependence at this order in the singular terms is formally enhanced over the dependence on the jet algorithm and jet area in the power-suppressed nonsingular terms that are not part of Eq. (28).

In Fig. 9 we compare the invariant mass, geometric p_T , and geometric E measures (using $\rho = 0.5$ for the latter two) for three different kinematic configurations with $\eta_J = \{0, 1, 2\}$, $Y = 0$, and $p_T^J = 300 \text{ GeV}$. For $\eta_J = 0$ and $\eta_J = 1$ the dependence of the jet mass on the jet definition is fairly mild (for jets of similar area), with the largest visible effect at small m_J . As we saw in Fig. 1(a) the jet regions of these two geometric measures agree exactly for $\eta_J = 0$ and differ progressively for more forward jet rapidities. This is reflected in the jet mass spectrum where the largest differences are observed for $\eta_J = 2$. The two bumps in the jet mass spectrum for geometric p_T at $\eta_J = 2$ are caused by an increased separation between the peaks of the individual quark and gluon channels [see Fig. 4(b)].

In Fig. 10 we show the jet mass spectrum for the geometric R measure for various values of the jet radius R . A larger jet radius translates into a peak at slightly

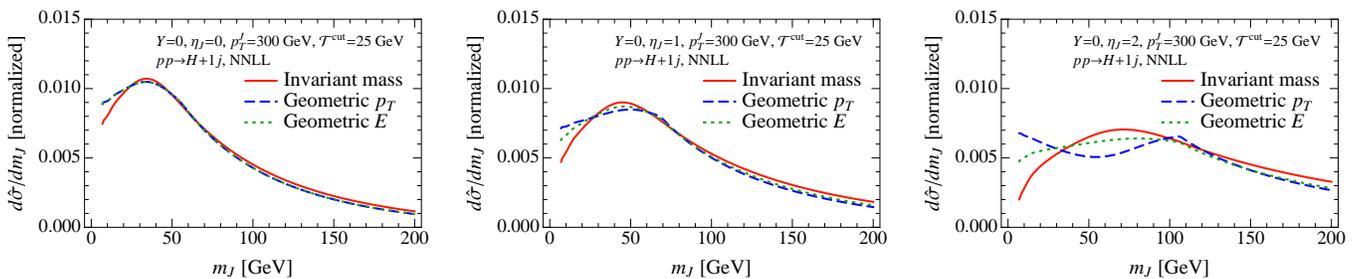


FIG. 9: Dependence of the NNLL jet mass spectrum for $pp \rightarrow H + 1$ jet on the N -jettiness measure used to define the jets.

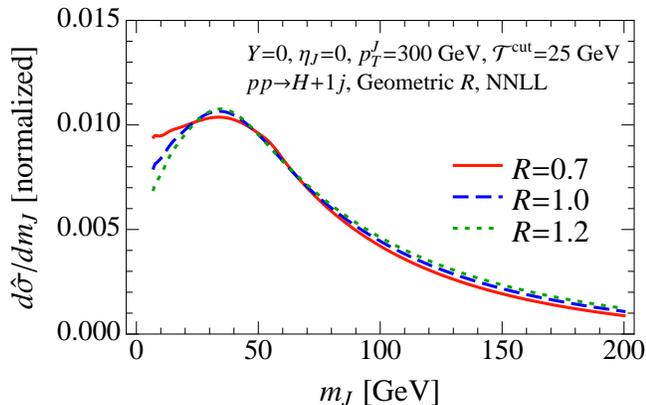


FIG. 10: Dependence of the NNLL jet mass spectrum for the geometric R measure on the jet radius R . Only the R dependence from singular terms in the factorization formula is shown here.

higher mass and slightly larger tail. Note that the biggest difference between the jet mass spectra for different values of R will be at large m_J in the tail of the spectrum, since the size of the jet puts an effective upper boundary on its mass $m_J \lesssim p_T^J R / \sqrt{2}$.⁵ At this boundary the jet mass spectrum has to fall off rapidly. This boundary is seen in PYTHIA and LHC data but is not included in our calculation, because we have not yet incorporated the nonsingular contributions to the cross section which are important for accounting for this effect and making realistic predictions in this part of the tail of the distribution.

⁵ For a jet of fixed three momentum centered at $\eta = 0$, the absolute bound on the jet mass is $m_J^2 < p_T^J{}^2(1/\cos^2 R - 1)$, which is reached when the jet consists of two energetic narrow subjets located at $(\eta, \phi) = (0, \pm R)$. Assuming a uniform energy distribution in (η, ϕ) -space, leads to a more practical bound $m_J^2 \lesssim p_T^J{}^2 R^2 / 2$.

VI. MONTE CARLO COMPARISONS

In this section we study various aspects of the jet mass spectrum in PYTHIA. Although formally the perturbative accuracy of PYTHIA is significantly lower than that of our NNLL calculation, it is also well known that after sufficient tuning PYTHIA is able to reproduce the shape of many jet observables. Here we are particularly interested in testing the impact on the jet mass spectrum from using different hard processes, using different jet algorithms, and from adding hadronization and underlying event (the latter being described by PYTHIA’s multi-parton interaction model). We also perform a comparison between our calculation and PYTHIA for the same geometric $R = 1$ N -jettiness jets used in our analysis. Finally we compare our exclusive 1-jet m_J calculation with the inclusive jet mass spectrum measured in $pp \rightarrow$ jets by ATLAS. We always use PYTHIA8 with its default tune 5 (“Tune 4C”), which as we will see provides a good description of the ATLAS jet mass data.

A. Hard Process and Jet Algorithm Dependence in PYTHIA

We start by investigating to what extent the jet mass spectrum depends on the underlying hard process in PYTHIA. In Fig. 11 we show the spectrum for a gluon jet from $gg \rightarrow gg$ (solid) and from $gg \rightarrow Hg$ (dotted), demonstrating that in PYTHIA there is essentially negligible process dependence for individual partonic channels. This is true both at the partonic level (blue curves with peak on the left) and after including hadronization and multiple interactions (red curves with peak on the right). In reality one expects some differences from the hard process due to the additional soft radiation produced with more available colored particles, and from the different color flow, where in particular $gg \rightarrow gg$ involves a matrix of color channels with nontrivial interference. These effects may not be sufficiently described by PYTHIA so one should not conclude that the hard process dependence on the jet mass spectrum is as small as is shown.

Next, we look at the difference in PYTHIA between the jet mass for exclusive and inclusive jet production. We use the process $gg \rightarrow Hg$, imposing the jet veto

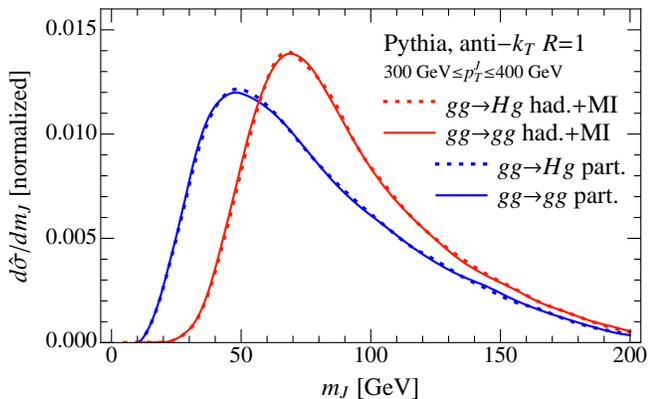


FIG. 11: The gluon jet mass spectrum in PYTHIA does not depend on the underlying hard process producing the jets. This is true both for partons (left peaks) and with hadronization and underlying event (right peaks).

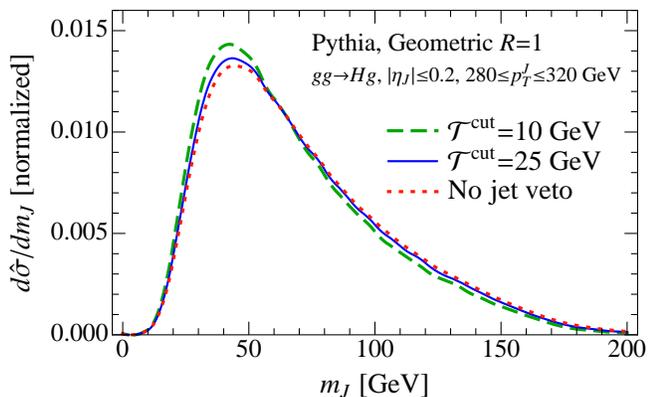


FIG. 12: Comparison of the normalized jet mass spectra for exclusive and inclusive jet samples in PYTHIA.

$\mathcal{T}^{\text{cut}} = 10, 25$ GeV to obtain two exclusive samples, and using no jet veto for our inclusive sample. The resulting normalized jet mass spectra are shown in Fig. 12. The difference between $\mathcal{T}^{\text{cut}} = 25$ GeV (our default value) and the inclusive case is small, allowing our calculation to be compared to inclusive spectra. The difference is slightly larger for $\mathcal{T}^{\text{cut}} = 10$ GeV and increases significantly for smaller values of \mathcal{T}^{cut} . However, we will not consider such strong jet vetos, as they lead to large NGLs (see Sec. IV C).

In Fig. 13 we compare the jet mass spectrum from PYTHIA for different jet algorithms, specifically our 1-jettiness $R = 1$ -algorithm, Cambridge-Aachen with $R = 1$, and anti- k_T with $R = 1$ and $R = 1.2$ [77]. To stay close to a calculation for a single phase space point, we restrict the jet to a narrow p_T and rapidity bin, and impose a veto using $\mathcal{T}^{\text{cut}} = 25$ GeV. The differences between the $R = 1$ curves are within the size of the uncertainty band from our NNLL calculation in the same phase space bin. This result agrees with the small differences observed in each of the panels of Fig. 9 from comparing

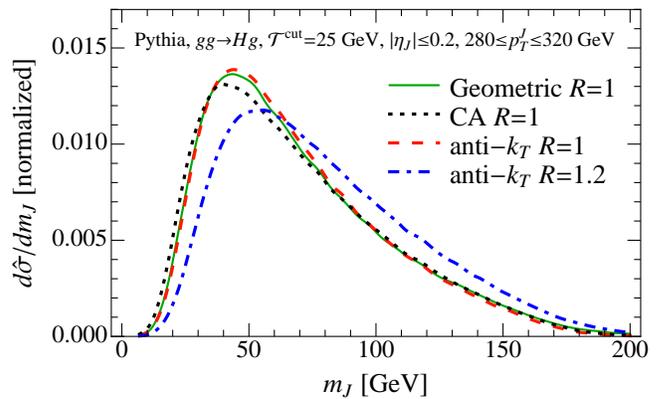


FIG. 13: Comparison of the anti- k_T , CA, and geometric R jet algorithms in PYTHIA.

different jet measures for 1-jettiness jets. The difference between $R = 1$ and $R = 1.2$ for anti- k_T is larger than that observed in our calculation using geometric R jets in Fig. 10. In PYTHIA the difference between $R = 1$ and $R = 1.2$ becomes smaller when \mathcal{T}^{cut} is decreased, since with a stronger jet veto less additional radiation is present that would be absorbed by larger jets. To be specific, the 15% difference in the peak heights for anti- k_T with $R = 1$ and $R = 1.2$ for $\mathcal{T}^{\text{cut}} = 25$ GeV reduces to 7% for $\mathcal{T}^{\text{cut}} = 5$ GeV. From these results we conclude that in our NNLL calculation, the R dependence of the non-singular terms that are not included in our analysis may well be numerically comparable or larger in size than the formally leading R dependence that we have computed.

B. Comparison of NNLL with PYTHIA

A comparison between our NNLL calculation and partonic PYTHIA results for $gg \rightarrow Hg$ are shown in Fig. 14, both using the geometric $R = 1$ jet definition. The peak positions in both cases agree very well. To ensure that this is not an accident and that the peak position in PYTHIA does not depend on the PDF set used by our default tune, we checked that an alternative tune (number 10, which is based on our default PYTHIA tune but uses MSTW2008 LO PDFs) only shifts the peak by a small amount, similar to the small difference in peak positions between PYTHIA and our NNLL calculation. However, as seen in Fig. 14, the NNLL calculation has a lower peak and a correspondingly higher tail. Since the spectrum is normalized these two effects are related, namely higher values in the tail must be compensated by a lower peak. There are several possibilities that may account for this difference. Due to the stability of our order-by-order results in Fig. 4(c) it is unlikely to be related to the lower order accuracy of PYTHIA's LL parton shower resummation. Most likely the differences are related to the fact that we have not yet included nonsingular contributions to the spectrum which are important in the tail region,

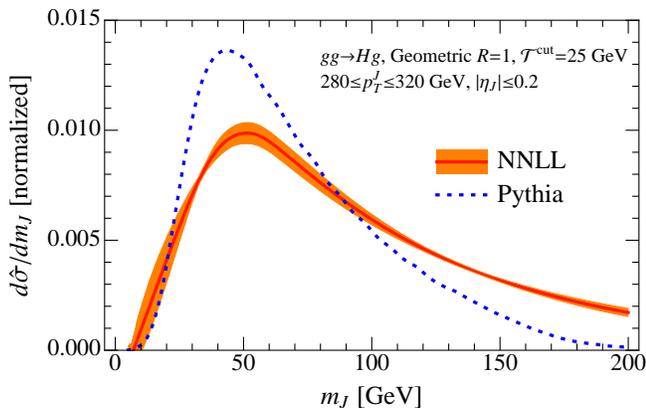


FIG. 14: Comparison between our NNLL calculation and partonic PYTHIA for the $gg \rightarrow Hg$ channel. Both results use geometric $R = 1$ jets and the same kinematic cuts.

in particular for the spectrum to fall off rapidly enough. Due to the fact that the results are normalized, this mismatch in the tail then also leads to a disagreement of the peak heights. Thus we expect that the inclusion of the nonsingular contributions will reduce this difference.

C. Hadronization in PYTHIA

We now explore the effect of hadronization on the jet mass spectrum using PYTHIA. In the factorization formula the hadronization is encoded through nonperturbative corrections in the soft function S at a scale $\sim \Lambda_{\text{QCD}}$, which must be separated from perturbative corrections at the soft scale $\mu_S \sim m_J^2/p_T^J$. For $e^+e^- \rightarrow 2$ jets there is an analytic understanding of the analogous nonperturbative corrections originating in Refs. [78–81] as well as a modern understanding in terms of field theory operators [19, 82–84]. For these processes, as soon as the relevant soft scale μ_S is perturbative, the nonperturbative corrections can be power expanded in $\Lambda_{\text{QCD}}/\mu_S$, and the dominant power correction simply shifts the event shape distribution, $e \rightarrow e - \Omega_e/Q$. In the case at hand, the nonperturbative soft function is built from more than two Wilson lines, so the description of the power corrections becomes more complicated. Nevertheless, for a given kinematic configuration we still expect that the dominant effect will be described by a shift involving a parameter $\Omega \sim \Lambda_{\text{QCD}}$. For a jet mass $m_J^2 \simeq p_T^+ p_T^-$ this shift occurs due to nonperturbative soft radiation causing a shift in the small momentum p_T^+ , so it takes the form

$$m_J^2 \rightarrow m_J^2 - 2\Omega p_T^J R. \quad (35)$$

The factor of R accounts for the fact that there is a decreased amount of soft momentum contamination in the jet for decreasing R [85]. It is straightforward to test whether this shift agrees with the hadronization model

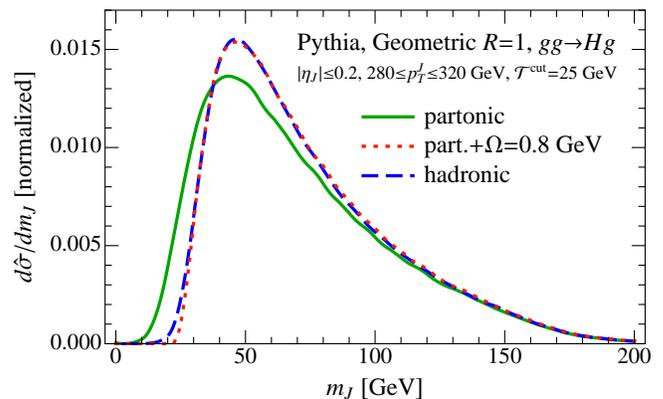


FIG. 15: The nonperturbative hadronization correction in PYTHIA is well described by a shift in m_J^2 .

in PYTHIA, by comparing the results with and without hadronization. As demonstrated in Fig. 15, a shift with the choice $\Omega = 0.8$ GeV works very well, in reasonable agreement with the $\Omega = 1.0$ GeV found earlier in Ref. [14] for the inclusive ≥ 1 jet cross section.

D. Underlying Event and ATLAS Data

In PYTHIA the effect of the underlying event is modeled by multiple partonic interactions, and its effect on the jet mass spectrum is more pronounced than that of hadronization. This is shown in Fig. 16 where we plot the jet mass spectrum for inclusive $pp \rightarrow$ jets from PYTHIA at parton level, including hadronization, and including hadronization and multiple interactions. Also shown are the corresponding ATLAS data from Ref. [26], where the uncertainty bars are from linearly combining the statistical and systematic uncertainties. This channel is dominated by the copious $pp \rightarrow$ dijet production at the LHC. We use the same inputs and cuts as ATLAS, namely $E_{\text{cm}} = 7$ TeV, anti- k_T jets with $R = 1$, $|\eta_J| \leq 2$, and consider both $300 \text{ GeV} \leq p_T^J \leq 400 \text{ GeV}$ and $500 \text{ GeV} \leq p_T^J \leq 600 \text{ GeV}$. The shift to the peak location from hadronization is of similar magnitude as that for $gg \rightarrow Hg$ in Fig. 15, namely $\simeq 3.0$ GeV for $gg \rightarrow Hg$ compared to $\simeq 8.0$ GeV for the $300 \text{ GeV} \leq p_T^J \leq 400 \text{ GeV}$ inclusive jets which have a slightly larger average p_T^J . For the inclusive $pp \rightarrow$ jets in PYTHIA the additional shift to the peak location from the underlying event is $\simeq 17.4$ GeV. The final PYTHIA results agree well with the ATLAS data for both p_T^J bins. In a NNLL calculation the effect of hadronization and part of the effect of the underlying event will be captured by corrections to the soft function, but it is not clear if hadronic corrections in the multi-jet soft function will fully capture the effect of the underlying event.

Given that PYTHIA agrees well with the ATLAS inclusive dijet spectrum, one might wonder what the purpose of a higher-order NNLL dijet calculation would be. An

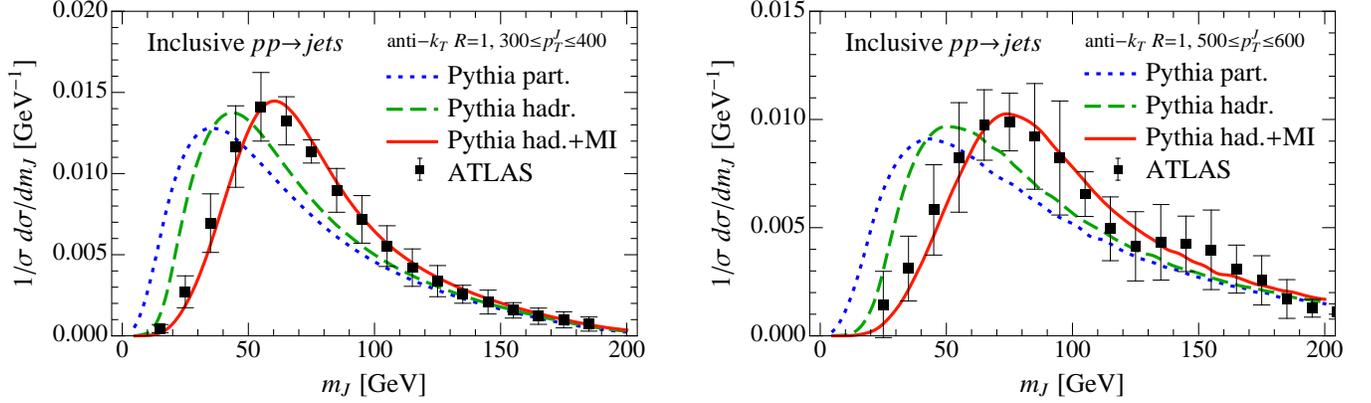


FIG. 16: Comparison of the PYTHIA jet mass spectrum for inclusive $pp \rightarrow$ jets to the corresponding ATLAS data [26]. PYTHIA results are shown at parton level (dotted), including hadronization (dashed), and including hadronization and multiple interactions (solid). The final PYTHIA results reproduce the data well.

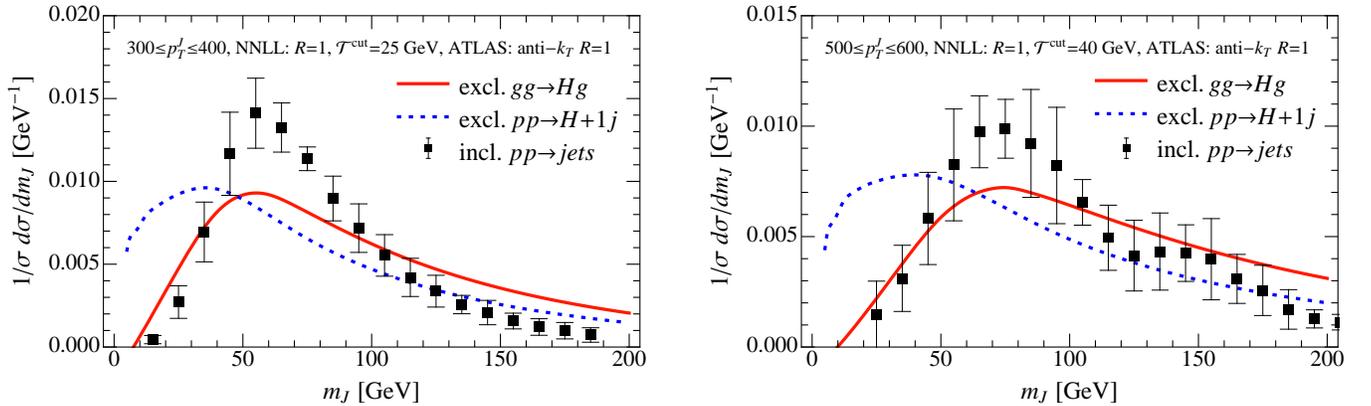


FIG. 17: Comparison of our exclusive NNLL calculation with ATLAS inclusive jet mass data [26]. The peak position of our gluon jets from $gg \rightarrow Hg$ agrees remarkably well with the inclusive dijet data. For the ATLAS data there is presumably a shift to lower values due to quark jets which is compensated by a shift to higher values due to hadronization and multiple interactions.

advantage of our calculational framework over PYTHIA is that it follows from first principles and does not involve the modeling and tuning present in PYTHIA. Specifically, the input to our calculation is limited to $\alpha_s(m_Z)$, the parton distributions functions, and simple soft function parameters like Ω for the hadronic effects. Furthermore, we have a rigorous estimate of the higher-order perturbative uncertainty from scale variation, as well as from order-by-order convergence, which enable us to fully assess the reliability of the result. Finally, it should be emphasized that our calculation is fully analytic (up to the numerical convolution with the PDFs) and hence provides an analytic QCD calculation of an LHC spectrum for jets.

To the extent that the normalized jet mass spectrum is independent of the hard process and independent of using an inclusive or exclusive jet sample, which PYTHIA seems to suggest in Figs. 11 and 12, a comparison between jet mass spectra involving different hard processes and with and without jet veto cuts is appropriate. The approx-

imate hard process independence only holds separately for gluon or quark jets, which themselves have fairly different jet mass spectra, see Fig. 4(b). Therefore when varying the hard process we expect the dominant change in the jet mass spectrum to be related to the process dependent fraction of quark and gluon jets produced.

In Fig. 17 we compare our NNLL result for $pp \rightarrow H+1$ jet and for $gg \rightarrow Hg$ to the ATLAS data for $pp \rightarrow$ jets. Recall that the peak location of the NNLL $H+1$ jet calculation matches well with that from PYTHIA, see Fig. 14. Because of the significant contribution from quark jets the $H+1$ jet spectrum peaks to the left of the spectrum from dijets. On the other hand, the peak location with pure gluon jets ($gg \rightarrow Hg$) agrees remarkably well with the data on dijets. From the results already obtained above, we expect only small differences (comparable to the ATLAS error bars) for effects related to the choice of the jet algorithm, the choice of inclusive versus exclusive jets, or the choice of looking at gluon jets in dijets

or in Higgs production. On the other hand there will be a more significant shift of the spectrum to the left from quark channels in the dijet production, and a shift to the right from adding hadronization and underlying event, neither of which is included in the solid red curve. The agreement between peak locations seems to indicate that these two effects largely compensate for one another. Finally, there will be an effect related to the fact that there are nontrivial color correlations in $gg \rightarrow gg$ which are not present in $gg \rightarrow Hg$ (these effects are not apparent in PYTHIA, see Fig. 11).

One may also look at the peak heights in Fig. 17, for which the agreement is not as good. As described earlier, this effect is related to the fact that we have not yet included nonsingular corrections. These corrections are known to decrease the tail to enable it to rapidly fall off by $m_J^2 \sim p_T^J R^2/2$, and they also affect the peak directly through the normalization. Since with additional work these can be included in future results the difference in peak heights is not of too much concern.

Finally one may also compare the results in Figs. 16 and 17 for the $300 \text{ GeV} \leq p_T^J \leq 400 \text{ GeV}$ and $500 \text{ GeV} \leq p_T^J \leq 600 \text{ GeV}$ bins. For dijets the peak location moves to higher m_J with increased p_T^J , unlike for $pp \rightarrow H + 1$ jet, again indicating that gluon jets likely dominate. The conclusions from the comparison with PYTHIA and the contrast to our NNLL calculation remains the same for these two ranges of p_T^J .

VII. CONCLUSIONS

In this paper we calculated the jet mass spectrum for $pp \rightarrow H + 1$ jet to NNLL order. For this exclusive 1-jet cross section we veto additional jets with the 1-jettiness event shape, and used the 1-jettiness factorization formula in terms of hard, beam, jet, and soft functions to obtain our results. For the *normalized* jet mass spectrum the remaining higher-order perturbative uncertainties from scale variation are at the $\simeq 5 - 10\%$ level at NNLL order, and in addition our results exhibit excellent order-by-order convergence.

The normalized NNLL spectrum is quite insensitive to the jet veto over a wide range of values, even when accounting for non-global logarithms. Thus in our framework non-global logarithms can be accurately treated as fixed-order contributions to the soft function, upon which additional global logarithms are automatically resummed in the factorization framework. An essential ingredient in the resummation of the global logarithms was the refactorization of the soft function, which we demonstrate is required to avoid introducing spurious leading logarithms in certain regions of phase space. Our treatment of the NNLL exclusive cross section with a jet veto has significantly smaller non-global logarithmic terms when compared to the size of these terms observed in the earlier inclusive NLL analysis in Ref. [10], and the earlier inclusive partial NNLL analysis in Ref. [31]. Finally, we note

that in PYTHIA the inclusive jet mass spectrum and the exclusive jet mass spectrum with our default jet veto are essentially identical.

Utilizing our calculation we investigated the dependence of the jet mass spectrum on various parameters of the exclusive jet cross section. Part of the power of our framework is that the factorization formula is fully differential in the jet kinematics (p_T^J , η_J , and Y), allowing us to vary the definition of the jets and the jet area, and can be easily separated into quark jet and gluon jet channels. As expected we find that the spectrum peaks at larger m_J values for gluon jets than for quark jets. For a given partonic channel the factorization framework predicts little sensitivity to the underlying hard process, and this result is also found to be the case in PYTHIA. The main process dependence is therefore the relative mix of quark and gluon jets. The peak of our NNLL m_J spectrum moves to the right for larger p_T^J and for larger $|\eta_J|$, but more so for the individual partonic channels than for $pp \rightarrow H + 1$ jet, where the change to the mix of quarks and gluons provides a compensating effect. The complete description of the various kinematic variables also makes it trivial to implement rapidity cuts. For a bin $|\eta_J| < 2$ and a not too large bin in p_T^J , we find that the integrated NNLL result is very consistent with the NNLL result for fixed kinematic variables taken at the center of the bin.

Varying the jet definition with fixed jet area leads to very little change in the jet mass spectrum, both for various jet definitions in our NNLL result and for anti- k_T , CA, and geometric- R jets in PYTHIA. This suggests that there are only small differences between the spectrum for 1-jettiness jets and traditional jet algorithms. On the other hand, PYTHIA exhibits a larger dependence on the jet radius R than our NNLL results. This is presumably due to the fact that we have not included nonsingular terms in our analysis. These terms become important in the tail region and are also needed to ensure that the jet mass spectrum dies off at $m_J^{\text{max}} \sim p_T^J R/\sqrt{2}$. This leads to a larger tail in our NNLL spectrum than in PYTHIA, and correspondingly a smaller peak height in the normalized NNLL result. On the other hand, the peak location agrees very well between our NNLL calculation and PYTHIA. An analysis of these additional nonsingular terms will be carried out in the future.

We investigated the dependence of the jet mass spectrum on hadronization and underlying event using PYTHIA. Hadronization is very well described by a shift to the mass spectrum, $m_J^2 \rightarrow m_J^2 - (2Rp_T^J)\Omega$ with $\Omega \sim \Lambda_{\text{QCD}}$, which is the anticipated result from non-perturbative soft gluon contributions in our factorization formula's soft function. In PYTHIA the underlying event is modeled by multiple partonic interactions and leads to a somewhat larger shift to the spectrum than for hadronization. It plays an important role in obtaining agreement with the ATLAS jet mass results for inclusive dijets. Comparing our results to ATLAS we find that the NNLL $pp \rightarrow H + 1$ jet spectrum peaks to the left of the dijet data, whereas the NNLL $gg \rightarrow Hg$ spectrum

peaks in the same location. The comparison made so far with the ATLAS data is promising. The extension of our NNLL calculation to $pp \rightarrow$ dijets is completely feasible using 2-jettiness, and it will be interesting to see to what extent the contributions from quark channels, color mixing, and hadronization and underlying event will affect this comparison with the data. Theoretically, the only remaining challenge to a complete comparison appears to be incorporating the effect of the underlying event from first principles rather than relying on its modeling via Monte Carlo.

Acknowledgments

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Appendix A: Perturbative Inputs

In this section we collect the fixed-order ingredients and evolution kernels for evaluating the jet mass cross section for $pp \rightarrow H + 1j$ in Eqs. (17) and (28) at NNLL order. We first give expressions for the hard, jet, beam and soft functions at next-to-leading order. This is followed by the evolution kernels and the coefficients that they depend on.

1. Hard Function

The hard functions H_κ for the various partonic channels κ that contribute to $pp \rightarrow H + 1$ jet can be obtained from the finite part of the helicity amplitudes A determined in Ref. [56],

$$\begin{aligned}
H_{ggg}(\{q_i^\mu\}, \mu_H) &= \frac{16\alpha_s(\mu_H)^3 C_A^2 C_F}{9\pi v^2} \frac{1}{[2(N_c^2 - 1)]^2} \left[|A(1_g^+, 2_g^+, 3_g^+; 4_H)|^2 + |A(1_g^+, 2_g^+, 3_g^-; 4_H)|^2 \right. \\
&\quad \left. + |A(1_g^+, 3_g^+, 2_g^-; 4_H)|^2 + |A(3_g^+, 2_g^+, 1_g^-; 4_H)|^2 \right], \\
H_{g\bar{q}\bar{q}}(\{q_i^\mu\}, \mu_H) &= \frac{8\alpha_s(\mu_H)^3 C_A C_F}{9\pi v^2} \frac{1}{2N_c} \frac{1}{2(N_c^2 - 1)} \left[|A(1_g^+, 2_q^+, 3_{\bar{q}}^-; 4_H)|^2 + |A(1_g^-, 2_q^+, 3_{\bar{q}}^-; 4_H)|^2 \right], \\
H_{\bar{q}g\bar{q}}(\{q_i^\mu\}, \mu_H) &= \frac{8\alpha_s(\mu_H)^3 C_A C_F}{9\pi v^2} \frac{1}{2N_c} \frac{1}{2(N_c^2 - 1)} \left[|A(2_g^+, 1_q^+, 3_{\bar{q}}^-; 4_H)|^2 + |A(2_g^-, 1_q^+, 3_{\bar{q}}^-; 4_H)|^2 \right], \\
H_{q\bar{q}g}(\{q_i^\mu\}, \mu_H) &= \frac{8\alpha_s(\mu_H)^3 C_A C_F}{9\pi v^2} \frac{1}{(2N_c)^2} \left[|A(3_g^+, 2_q^+, 1_{\bar{q}}^-; 4_H)|^2 + |A(3_g^-, 2_q^+, 1_{\bar{q}}^-; 4_H)|^2 \right], \\
H_{gqq}(\{q_i^\mu\}, \mu_H) &= H_{g\bar{q}\bar{q}}(\{q_i^\mu\}, \mu_H), \quad H_{qqq}(\{q_i^\mu\}, \mu_H) = H_{\bar{q}g\bar{q}}(\{q_i^\mu\}, \mu_H), \\
H_{\bar{q}qg}(\{q_i^\mu\}, \mu_H) &= H_{q\bar{q}g}(\{q_i^\mu\}, \mu_H).
\end{aligned} \tag{A1}$$

The factors of $1/(2N_c)$ and $1/[2(N_c^2 - 1)]$ arise from averaging over the spins and colors of the colliding quarks and gluons. The arguments of a helicity amplitude A have the form i_t^h , where i denotes the momentum q_i^μ , t denotes the parton type, and h denotes the helicity of this particle. *Only in the helicity amplitudes* will we use an outgoing convention for all these quantities, to make crossing symmetry direct. This implies that if we want to convert to the convention used in the main text, then the s_{ij} 's in the helicity amplitudes below will pick up additional minus signs if one of the particles i and j is in and the other is out. The amplitudes that enter in Eq. (A1) are given by

$$\begin{aligned}
A(1_g^+, 2_g^+, 3_g^+; 4_H) &= \frac{m_H^4}{\sqrt{2|s_{12}s_{13}s_{23}}|} \left\{ 1 + \frac{\alpha_s(\mu_H)}{4\pi} \left[f(s_{12}, s_{13}, s_{23}, m_H^2, \mu_H) + \frac{1}{3}(C_A - 2T_F n_f) \frac{s_{12}s_{13} + s_{12}s_{23} + s_{13}s_{23}}{m_H^4} \right] \right\}, \\
A(1_g^+, 2_g^+, 3_g^-; 4_H) &= \frac{s_{12}^2}{\sqrt{2|s_{12}s_{13}s_{23}}|} \left\{ 1 + \frac{\alpha_s(\mu_H)}{4\pi} \left[f(s_{12}, s_{13}, s_{23}, m_H^2, \mu_H) + \frac{1}{3}(C_A - 2T_F n_f) \frac{s_{13}s_{23}}{s_{12}^2} \right] \right\}, \\
A(1_g^+, 2_q^+, 3_{\bar{q}}^-; 4_H) &= \frac{s_{12}}{\sqrt{2|s_{23}}|} \left\{ 1 + \frac{\alpha_s(\mu_H)}{4\pi} \left[g(s_{12}, s_{13}, s_{23}, m_H^2, \mu_H) + (C_F - C_A) \frac{s_{23}}{s_{12}} \right] \right\}, \\
A(1_g^-, 2_q^+, 3_{\bar{q}}^-; 4_H) &= \frac{s_{13}}{\sqrt{2|s_{23}}|} \left\{ 1 + \frac{\alpha_s(\mu_H)}{4\pi} \left[g(s_{12}, s_{13}, s_{23}, m_H^2, \mu_H) + (C_F - C_A) \frac{s_{23}}{s_{13}} \right] \right\}, \\
f(s_{12}, s_{13}, s_{23}, m_H^2, \mu_H) &= -C_A \left[\frac{1}{2}(L_{12}^2 + L_{13}^2 + L_{23}^2) + L_{12/H}L_{13/H} + L_{12/H}L_{23/H} + L_{13/H}L_{23/H} \right]
\end{aligned}$$

$$\begin{aligned}
& + 2\text{Li}_2\left(1 - \frac{s_{12}}{m_H^2}\right) + 2\text{Li}_2\left(1 - \frac{s_{13}}{m_H^2}\right) + 2\text{Li}_2\left(1 - \frac{s_{23}}{m_H^2}\right) - 5 - \frac{3\pi^2}{4} \Big] - 3C_F, \\
g(s_{12}, s_{13}, s_{23}, m_H^2, \mu_H) = & C_A \left[-\frac{1}{2}(L_{12}^2 + L_{13}^2 - L_{23}^2) + L_{12/H}L_{13/H} - (L_{12/H} + L_{13/H})L_{23/H} - 2\text{Li}_2\left(1 - \frac{s_{23}}{m_H^2}\right) \right. \\
& + \frac{22}{3} + \frac{\pi^2}{4} \Big] + C_F \left[-L_{23}^2 + 3L_{23} - 2L_{12/H}L_{13/H} - 2\text{Li}_2\left(1 - \frac{s_{12}}{m_H^2}\right) - 2\text{Li}_2\left(1 - \frac{s_{13}}{m_H^2}\right) \right. \\
& \left. - 11 + \frac{\pi^2}{2} \right] + \beta_0 \left(-L_{23} + \frac{5}{3} \right). \tag{A2}
\end{aligned}$$

Here we use the shorthand notation

$$\begin{aligned}
L_{ij} &= \ln\left(-\frac{s_{ij}}{\mu_H^2} - i0\right), \\
L_{ij/H} &= \ln\left(-\frac{s_{ij}}{\mu_H^2} - i0\right) - \ln\left(-\frac{m_H^2}{\mu_H^2} - i0\right). \tag{A3}
\end{aligned}$$

The $\mathcal{L}_n(x)$ integrate to zero if the range in x is $[0, 1]$.

2. Jet Functions

The one-loop jet functions are given by [58–60]

$$\begin{aligned}
J_q(s, \mu_J) &= \delta(s) + \frac{\alpha_s(\mu_J)C_F}{2\pi} \left[\frac{2}{\mu_J^2} \mathcal{L}_1\left(\frac{s}{\mu_J^2}\right) - \frac{3}{2\mu_J^2} \mathcal{L}_0\left(\frac{s}{\mu_J^2}\right) - \left(\frac{\pi^2}{2} - \frac{7}{2}\right) \delta(s) \right], \\
J_g(s, \mu_J) &= \delta(s) + \frac{\alpha_s(\mu_J)}{2\pi} \left\{ \frac{2C_A}{\mu_J^2} \mathcal{L}_1\left(\frac{s}{\mu_J^2}\right) - \frac{\beta_0}{2\mu_J^2} \mathcal{L}_0\left(\frac{s}{\mu_J^2}\right) + \left[\left(\frac{2}{3} - \frac{\pi^2}{2}\right) C_A + \frac{5}{6} \beta_0 \right] \delta(s) \right\}, \tag{A4}
\end{aligned}$$

where the plus distributions \mathcal{L}_n are defined as

$$\mathcal{L}_n(x) \equiv \left[\frac{\theta(x) \ln^n x}{x} \right]_+$$

3. Beam Functions

The beam functions can be expressed in terms of standard gluon and quark PDFs using an operator product expansion [61, 86],

$$\begin{aligned}
B_i(t, x, \mu_B) &= \sum_{j=\{g, q, \bar{q}\}} \int_x^1 \frac{d\xi}{\xi} \mathcal{I}_{ij}\left(t, \frac{x}{\xi}, \mu_B\right) f_j(\xi, \mu_B) \\
&\times \left[1 + \mathcal{O}\left(\frac{\Lambda_{\text{QCD}}^2}{t}\right) \right]. \tag{A6}
\end{aligned}$$

The one-loop matching coefficients are [62, 64] are

$$\begin{aligned}
\mathcal{I}_{qq}(t, z, \mu_B) &= \delta(t) \delta(1-z) + \frac{\alpha_s(\mu_B)C_F}{2\pi} \theta(z) \left\{ \frac{2}{\mu_B^2} \mathcal{L}_1\left(\frac{t}{\mu_B^2}\right) \delta(1-z) + \frac{1}{\mu_B^2} \mathcal{L}_0\left(\frac{t}{\mu_B^2}\right) P_{qq}(z) \right. \\
&\quad \left. + \delta(t) \left[\mathcal{L}_1(1-z)(1+z^2) - P_{qq}(z) \ln z - \frac{\pi^2}{6} \delta(1-z) + \theta(1-z)(1-z) \right] \right\}, \\
\mathcal{I}_{qg}(t, z, \mu_B) &= \frac{\alpha_s(\mu_B)T_F}{2\pi} \theta(z) \left\{ \frac{1}{\mu_B^2} \mathcal{L}_0\left(\frac{t}{\mu_B^2}\right) P_{qg}(z) + \delta(t) \left[P_{qg}(z) \left(\ln \frac{1-z}{z} - 1 \right) + \theta(1-z) \right] \right\} \\
\mathcal{I}_{gg}(t, z, \mu_B) &= \delta(t) \delta(1-z) + \frac{\alpha_s(\mu_B)C_A}{2\pi} \theta(z) \left\{ \frac{2}{\mu_B^2} \mathcal{L}_1\left(\frac{t}{\mu_B^2}\right) \delta(1-z) + \frac{1}{\mu_B^2} \mathcal{L}_0\left(\frac{t}{\mu_B^2}\right) P_{gg}(z) \right. \\
&\quad \left. + \delta(t) \left[\mathcal{L}_1(1-z) \frac{2(1-z+z^2)^2}{z} - P_{gg}(z) \ln z - \frac{\pi^2}{6} \delta(1-z) \right] \right\}, \\
\mathcal{I}_{gq}(t, z, \mu_B) &= \frac{\alpha_s(\mu_B)C_F}{2\pi} \theta(z) \left\{ \frac{1}{\mu_B^2} \mathcal{L}_0\left(\frac{t}{\mu_B^2}\right) P_{gq}(z) + \delta(t) \left[P_{gq}(z) \ln \frac{1-z}{z} + \theta(1-z)z \right] \right\}. \tag{A7}
\end{aligned}$$

The splitting functions in this equation are defined as

$$P_{qq}(z) = \mathcal{L}_0(1-z)(1+z^2),$$

$$P_{qg}(z) = \theta(1-z)[(1-z)^2 + z^2],$$

$$\begin{aligned}
P_{gg}(z) &= 2\mathcal{L}_0(1-z)z + 2\theta(1-z)\left[\frac{1-z}{z} + z(1-z)\right], \\
P_{gq}(z) &= \theta(1-z)\frac{1+(1-z)^2}{z}.
\end{aligned} \tag{A8}$$

4. Factorized Soft Function

We now give expressions for the N -jettiness soft function, showing explicitly how the factorization in Eq. (24) is implemented. We remind the reader that their is some

freedom in this refactorization, and that the corresponding uncertainty is probed by varying the parameter r in Eq. (31).

Up to NLO the 1-jettiness soft function is given by

$$S_\kappa(\{k_i\}, \{\mu_{S_i}\}) = \prod_{i=a,b,J} S_i(k_i, \{\hat{q}_i^\mu\}, \mu_{S_i}) + \mathcal{O}(\alpha_s^2). \tag{A9}$$

From the NLO calculation in Ref. [47] we obtain

$$\begin{aligned}
S_i(k_i, \{\hat{q}_i^\mu\}, \mu_{S_i}) &= \mathbf{1} \delta(k_i) + \frac{\alpha_s(\mu_{S_i})}{\pi} \sum_{j \neq i} \left\{ \mathbf{T}_i \cdot \mathbf{T}_j \left[\frac{2}{\sqrt{\hat{s}_{ij}} \mu_{S_i}} \mathcal{L}_1\left(\frac{k_i}{\sqrt{\hat{s}_{ij}} \mu_{S_i}}\right) - \frac{\pi^2}{24} \delta(k_i) \right] \right. \\
&+ \sum_{m \neq i,j} \left[\left\{ \mathbf{T}_i \cdot \mathbf{T}_j I_0\left(\frac{\hat{s}_{jm}}{\hat{s}_{ij}}, \frac{\hat{s}_{im}}{\hat{s}_{ij}}\right) - \mathbf{T}_m \cdot \mathbf{T}_j I_0\left(\frac{\hat{s}_{ij}}{\hat{s}_{mj}}, \frac{\hat{s}_{im}}{\hat{s}_{mj}}\right) \right\} \frac{1}{\mu} \mathcal{L}_0\left(\frac{k_i}{\mu_{S_i}}\right) \right. \\
&\left. \left. + \frac{1}{6} \left\{ \mathbf{T}_i \cdot \mathbf{T}_j \left[I_0\left(\frac{\hat{s}_{jm}}{\hat{s}_{ij}}, \frac{\hat{s}_{im}}{\hat{s}_{ij}}\right) \ln \frac{\hat{s}_{jm}}{\hat{s}_{ij}} + I_1\left(\frac{\hat{s}_{jm}}{\hat{s}_{ij}}, \frac{\hat{s}_{im}}{\hat{s}_{ij}}\right) \right] + 5 \text{ permutations of } (i, j, k) \right\} \delta(k_i) \right] \right\}.
\end{aligned} \tag{A10}$$

Here the two integrals are

$$\begin{aligned}
I_0(\alpha, \beta) &= \frac{1}{\pi} \int_{-\pi}^{\pi} d\phi \int \frac{dy}{y} \theta(y - \sqrt{\beta/\alpha}) \theta(1/\alpha - 1 - y^2 + 2y \cos \phi), \\
I_1(\alpha, \beta) &= \frac{1}{\pi} \int_{-\pi}^{\pi} d\phi \int \frac{dy}{y} \ln(1 + y^2 - 2y \cos \phi) \theta(y - \sqrt{\beta/\alpha}) \theta(1/\alpha - 1 - y^2 + 2y \cos \phi),
\end{aligned} \tag{A11}$$

and

$$\begin{aligned}
gg \rightarrow Hg : \quad \mathbf{T}_a^2 &= \mathbf{T}_b^2 = \mathbf{T}_J^2 = C_A, \\
\mathbf{T}_a \cdot \mathbf{T}_b &= \mathbf{T}_a \cdot \mathbf{T}_J = \mathbf{T}_b \cdot \mathbf{T}_J = -\frac{C_A}{2}, \\
gq \rightarrow Hq : \quad \mathbf{T}_a^2 &= C_A, \quad \mathbf{T}_b^2 = \mathbf{T}_J^2 = C_F, \\
\mathbf{T}_a \cdot \mathbf{T}_b &= \mathbf{T}_a \cdot \mathbf{T}_J = -\frac{C_A}{2}, \\
\mathbf{T}_b \cdot \mathbf{T}_J &= \frac{C_A}{2} - C_F.
\end{aligned} \tag{A12}$$

5. Evolution Factors

Following the discussion in Sec. III B, we give expressions for the factorized evolution of the hard function,

$$\begin{aligned}
H_\kappa(\{q_j^\mu\}, \{\mu_i\}) &= H_\kappa(\{q_i^\mu\}, \mu_H) \prod_{i=a,b,J} U_{H_{\kappa_i}}(\{q_j^\mu\}, \mu_H, \mu_i), \\
U_{H_{\kappa_i}}(\{q_j^\mu\}, \mu_H, \mu_i) &= \left| e^{K_H^i} \prod_{j \neq i} \left(\frac{-s_{ij} - i0}{\mu_H^2} \right)^{\mathbf{T}_i \cdot \mathbf{T}_j \eta_H} \right|, \\
K_H^i(\mu_H, \mu_i) &= -2K_{\Gamma^{\kappa_i}}(\mu_H, \mu_i) + K_{\gamma_H^{\kappa_i}}(\mu_H, \mu_i), \\
\eta_H(\mu_H, \mu_i) &= -\frac{\eta_{\Gamma^q}(\mu_H, \mu_i)}{C_F} = -\frac{\eta_{\Gamma^g}(\mu_H, \mu_i)}{C_A}.
\end{aligned} \tag{A13}$$

Here the products over i and j run over all colored particles, with corresponding flavor κ_i and κ_j . For each channel contributing to $pp \rightarrow H + 1j$ there is only one color structure so $\mathbf{T}_i \cdot \mathbf{T}_j$ is simply a number [see Eq. (A12)]. The functions K_Γ , η_Γ and K_γ are given below in Eq. (A16).

The solution of the RG evolution of the jet function is given by [37, 73, 87, 88]

$$\begin{aligned}
J_{\kappa_i}(s, \mu) &= \int ds' J_{\kappa_i}(s - s', \mu_J) U_{J_{\kappa_i}}(s', \mu_J, \mu), \\
U_{J_{\kappa_i}}(s, \mu_J, \mu) &= \frac{e^{K_J^i - \gamma_E \eta_J^i}}{\Gamma(1 + \eta_J^i)} \left[\frac{\eta_J^i}{\mu_J^2} \mathcal{L}^{\eta_J^i}\left(\frac{s}{\mu_J^2}\right) + \delta(s) \right], \\
K_J^i(\mu_J, \mu) &= 4K_{\Gamma^{\kappa_i}}(\mu_J, \mu) + K_{\gamma_J^{\kappa_i}}(\mu_J, \mu), \\
\eta_J^i(\mu_J, \mu) &= -2\eta_{\Gamma^{\kappa_i}}(\mu_J, \mu).
\end{aligned} \tag{A14}$$

The plus distribution \mathcal{L}^η is defined as

$$\begin{aligned}
\mathcal{L}^\eta(x) &\equiv \left[\frac{\theta(x)}{x^{1-\eta}} \right]_+ \\
&= \lim_{\beta \rightarrow 0} \left[\frac{\theta(x - \beta)}{x^{1-\eta}} + \delta(x - \beta) \frac{x^\eta - 1}{\eta} \right].
\end{aligned} \tag{A15}$$

General relations for the rescaling and convolutions of $\mathcal{L}_n(x)$ in Eq. (A5) and $\mathcal{L}^\eta(x)$ can be found in App. B

of Ref. [73]. The renormalization group evolution of the beam functions is identical [62] and can be obtained from the above expressions by replacing $J_i(s, \mu) \rightarrow B_i(t, x, \mu)$. We do not give the evolution of the soft function, as it is not needed for evaluating Eq. (28). It can be obtained

from the evolution of the hard function and beam function by using the μ -independence of the cross section.

The functions $K_\Gamma(\mu_0, \mu)$, $\eta_\Gamma(\mu_0, \mu)$, $K_\gamma(\mu_0, \mu)$ in the above RGE solutions at NNLL are given by,

$$\begin{aligned}
K_\Gamma(\mu_0, \mu) &= -\frac{\Gamma_0}{4\beta_0^2} \left\{ \frac{4\pi}{\alpha_s(\mu_0)} \left(1 - \frac{1}{r} - \ln r \right) + \left(\frac{\Gamma_1}{\Gamma_0} - \frac{\beta_1}{\beta_0} \right) (1 - r + \ln r) + \frac{\beta_1}{2\beta_0} \ln^2 r \right. \\
&\quad \left. + \frac{\alpha_s(\mu_0)}{4\pi} \left[\left(\frac{\beta_1^2}{\beta_0^2} - \frac{\beta_2}{\beta_0} \right) \left(\frac{1-r^2}{2} + \ln r \right) + \left(\frac{\beta_1\Gamma_1}{\beta_0\Gamma_0} - \frac{\beta_1^2}{\beta_0^2} \right) (1 - r + r \ln r) - \left(\frac{\Gamma_2}{\Gamma_0} - \frac{\beta_1\Gamma_1}{\beta_0\Gamma_0} \right) \frac{(1-r)^2}{2} \right] \right\}, \\
\eta_\Gamma(\mu_0, \mu) &= -\frac{\Gamma_0}{2\beta_0} \left[\ln r + \frac{\alpha_s(\mu_0)}{4\pi} \left(\frac{\Gamma_1}{\Gamma_0} - \frac{\beta_1}{\beta_0} \right) (r-1) + \frac{\alpha_s^2(\mu_0)}{16\pi^2} \left(\frac{\Gamma_2}{\Gamma_0} - \frac{\beta_1\Gamma_1}{\beta_0\Gamma_0} + \frac{\beta_1^2}{\beta_0^2} - \frac{\beta_2}{\beta_0} \right) \frac{r^2-1}{2} \right], \\
K_\gamma(\mu_0, \mu) &= -\frac{\gamma_0}{2\beta_0} \left[\ln r + \frac{\alpha_s(\mu_0)}{4\pi} \left(\frac{\gamma_1}{\gamma_0} - \frac{\beta_1}{\beta_0} \right) (r-1) \right]. \tag{A16}
\end{aligned}$$

Here, $r = \alpha_s(\mu)/\alpha_s(\mu_0)$ and the running coupling at the scale μ is given in terms of that at the reference scale μ_0 by the three-loop expression

$$\begin{aligned}
\frac{1}{\alpha_s(\mu)} &= \frac{X}{\alpha_s(\mu_0)} + \frac{\beta_1}{4\pi\beta_0} \ln X + \frac{\alpha_s(\mu_0)}{16\pi^2} \left[\frac{\beta_2}{\beta_0} \left(1 - \frac{1}{X} \right) \right. \\
&\quad \left. + \frac{\beta_1^2}{\beta_0^2} \left(\frac{\ln X}{X} + \frac{1}{X} - 1 \right) \right], \tag{A17}
\end{aligned}$$

where $X \equiv 1 + \alpha_s(\mu_0)\beta_0 \ln(\mu/\mu_0)/(2\pi)$.

6. RGE Coefficients

Up to three loops, the coefficients of the beta function [89, 90] and cusp anomalous dimension [65, 66] in $\overline{\text{MS}}$ are

$$\begin{aligned}
\beta_0 &= \frac{11}{3} C_A - \frac{4}{3} T_F n_f, \tag{A18} \\
\beta_1 &= \frac{34}{3} C_A^2 - \left(\frac{20}{3} C_A + 4C_F \right) T_F n_f, \\
\beta_2 &= \frac{2857}{54} C_A^3 + \left(C_F^2 - \frac{205}{18} C_F C_A - \frac{1415}{54} C_A^2 \right) 2T_F n_f \\
&\quad + \left(\frac{11}{9} C_F + \frac{79}{54} C_A \right) 4T_F^2 n_f^2, \\
\Gamma_0^q &= 4C_F, \\
\Gamma_1^q &= 4C_F \left[\left(\frac{67}{9} - \frac{\pi^2}{3} \right) C_A - \frac{20}{9} T_F n_f \right], \\
\Gamma_2^q &= 4C_F \left[\left(\frac{245}{6} - \frac{134\pi^2}{27} + \frac{11\pi^4}{45} + \frac{22\zeta_3}{3} \right) C_A^2 \right. \\
&\quad \left. + \left(-\frac{418}{27} + \frac{40\pi^2}{27} - \frac{56\zeta_3}{3} \right) C_A T_F n_f \right. \\
&\quad \left. + \left(-\frac{55}{3} + 16\zeta_3 \right) C_F T_F n_f - \frac{16}{27} T_F^2 n_f^2 \right], \\
\Gamma_n^q &= \frac{C_A}{C_F} \Gamma_n^q \quad \text{for } n \leq 2.
\end{aligned}$$

Up to two loops, the $\overline{\text{MS}}$ non-cusp anomalous dimension for the hard function [91, 92] and jet and beam functions [60, 62, 64, 70] are

$$\begin{aligned}
\gamma_{H0}^q &= -6C_F, \tag{A19} \\
\gamma_{H1}^q &= -C_F \left[\left(\frac{82}{9} - 52\zeta_3 \right) C_A + (3 - 4\pi^2 + 48\zeta_3) C_F \right. \\
&\quad \left. + \left(\frac{65}{9} + \pi^2 \right) \beta_0 \right], \\
\gamma_{H0}^g &= -2\beta_0, \\
\gamma_{H1}^g &= \left(-\frac{118}{9} + 4\zeta_3 \right) C_A^2 + \left(-\frac{38}{9} + \frac{\pi^2}{3} \right) C_A \beta_0 - 2\beta_1, \\
\gamma_{J0}^q &= 6C_F, \\
\gamma_{J1}^q &= C_F \left[\left(\frac{146}{9} - 80\zeta_3 \right) C_A + (3 - 4\pi^2 + 48\zeta_3) C_F \right. \\
&\quad \left. + \left(\frac{121}{9} + \frac{2\pi^2}{3} \right) \beta_0 \right], \\
\gamma_{J0}^g &= 2\beta_0, \\
\gamma_{J1}^g &= \left(\frac{182}{9} - 32\zeta_3 \right) C_A^2 + \left(\frac{94}{9} - \frac{2\pi^2}{3} \right) C_A \beta_0 + 2\beta_1.
\end{aligned}$$

Appendix B: Running Scales

We now present the remaining ingredients that enter in the running scales in Sec. III C. First of all, μ_{run} is defined as

$$\begin{aligned}
\mu_{\text{run}}(\tau, \mu) &= \begin{cases} \mu_0 + a\tau^2/\tau_1 & \tau \leq \tau_1, \\ 2a\tau + b & \tau_1 \leq \tau \leq \tau_2, \\ \mu - a(\tau - \tau_3)^2/(\tau_3 - \tau_2) & \tau_2 \leq \tau \leq \tau_3, \\ \mu & \tau > \tau_3, \end{cases} \\
a &= \frac{\mu_0 - \mu}{\tau_1 - \tau_2 - \tau_3}, \quad b = \frac{\mu\tau_1 - \mu_0(\tau_2 + \tau_3)}{\tau_1 - \tau_2 - \tau_3}. \tag{B1}
\end{aligned}$$

The expressions for a and b follow from demanding that $\mu_{\text{run}}(\tau)$ is continuous and has a continuous derivative. For our central scale choice we use

$$\begin{aligned} \mu &= Q, \quad e_i = e_{S_i} = 0, \quad \mu_0 = 2\text{GeV}, \\ \tau_1 &= \frac{5\text{GeV}}{Q}, \quad \tau_2 = 0.4, \quad \tau_3 = 0.6, \quad r = 0.2. \end{aligned} \quad (\text{B2})$$

To estimate the perturbative uncertainty we vary the above parameters within reasonable ranges. Since the cross section is most sensitive to μ , e_i , e_{S_i} and r , we restrict ourselves to the following separate variations,

$$\text{a) } \mu = 2^{\pm 1}Q, \quad e_J = e_B = e_{S_J} = e_{S_B} = 0, \quad r = 0.2,$$

$$\begin{aligned} \text{b) } &\mu = Q, \quad e_J = \pm 0.5, \quad e_B = e_{S_J} = e_{S_B} = 0, \quad r = 0.2, \\ \text{c) } &\mu = Q, \quad e_B = \pm 0.5, \quad e_J = e_{S_J} = e_{S_B} = 0, \quad r = 0.2, \\ \text{d) } &\mu = Q, \quad e_{S_J} = \pm 0.5, \quad e_J = e_B = e_{S_B} = 0, \quad r = 0.2, \\ \text{e) } &\mu = Q, \quad e_{S_B} = \pm 0.5, \quad e_J = e_B = e_{S_J} = 0, \quad r = 0.2, \\ \text{f) } &\mu = Q, \quad e_J = e_B = e_{S_J} = e_{S_B} = 0, \quad r = 0.2 \pm 0.2. \end{aligned} \quad (\text{B3})$$

Following our discussion in Refs. [93, 94], we take the envelope of variations b) through f) and add this in quadrature with variation a).

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- [1] A. Altheimer *et al.*, J. Phys. G **G39**, 063001 (2012), [arXiv:1201.0008].
- [2] S. D. Ellis, A. Hornig, C. Lee, C. K. Vermilion, and J. R. Walsh, Phys. Lett. B **689**, 82 (2010), [arXiv:0912.0262].
- [3] C. W. Bauer, F. J. Tackmann, J. R. Walsh, and S. Zuberi, Phys. Rev. D **85**, 074006 (2012), [arXiv:1106.6047].
- [4] I. Feige, M. D. Schwartz, I. W. Stewart, and J. Thaler, Phys. Rev. Lett. **109**, 092001 (2012), [arXiv:1204.3898].
- [5] A. J. Larkoski, Phys. Rev. D **86**, 054004 (2012), [arXiv:1207.1437].
- [6] D. Krohn, T. Lin, M. D. Schwartz, and W. J. Waalewijn, arXiv:1209.2421.
- [7] W. J. Waalewijn, Phys. Rev. D **86**, 094030 (2012), [arXiv:1209.3019].
- [8] S. D. Ellis, C. K. Vermilion, J. R. Walsh, A. Hornig, and C. Lee, JHEP **11**, 101 (2010), [arXiv:1001.0014].
- [9] T. T. Jouttenus, Phys. Rev. D **81**, 094017 (2010), [arXiv:0912.5509].
- [10] A. Banfi, M. Dasgupta, K. Khelifa-Kerfa, and S. Marzani, JHEP **08**, 064 (2010), [arXiv:1004.3483].
- [11] R. Kelley, M. D. Schwartz, and H. X. Zhu, arXiv:1102.0561.
- [12] R. Kelley, M. D. Schwartz, R. M. Schabinger, and H. X. Zhu, Phys. Rev. D **86**, 054017 (2012), [arXiv:1112.3343].
- [13] H.-n. Li, Z. Li, and C.-P. Yuan, arXiv:1206.1344.
- [14] M. Dasgupta, K. Khelifa-Kerfa, S. Marzani, and M. Spannowsky, JHEP **1210**, 126 (2012), [arXiv:1207.1640].
- [15] Y.-T. Chien, R. Kelley, M. D. Schwartz, and H. X. Zhu, Phys. Rev. D **87**, 014010 (2013), [arXiv:1208.0010].
- [16] D. Krohn, J. Thaler, and L.-T. Wang, JHEP **1002**, 084 (2010), [arXiv:0912.1342].
- [17] J. M. Butterworth, A. R. Davison, M. Rubin, and G. P. Salam, Phys. Rev. Lett. **100**, 242001 (2008), [arXiv:0802.2470].
- [18] S. D. Ellis, C. K. Vermilion, and J. R. Walsh, Phys. Rev. D **80**, 051501 (2009), [arXiv:0903.5081].
- [19] G. P. Korchemsky and G. Sterman, Nucl. Phys. B **555**, 335 (1999), [hep-ph/9902341].
- [20] C. W. Bauer, D. Pirjol, and I. W. Stewart, Phys. Rev. D **65**, 054022 (2002), [hep-ph/0109045].
- [21] S. Fleming, A. H. Hoang, S. Mantry, and I. W. Stewart, Phys. Rev. D **77**, 074010 (2008), [hep-ph/0703207].
- [22] T. Becher and M. D. Schwartz, JHEP **07**, 034 (2008), [arXiv:0803.0342].
- [23] R. Abbate, M. Fickinger, A. H. Hoang, V. Mateu, and I. W. Stewart, Phys. Rev. D **83**, 074021 (2011), [arXiv:1006.3080].
- [24] T. Gehrmann, G. Luisoni, and P. F. Monni, arXiv:1210.6945.
- [25] L. J. Dixon, L. Magnea, and G. Sterman, JHEP **08**, 022 (2008), [arXiv:0805.3515].
- [26] ATLAS Collaboration Collaboration, G. Aad *et al.*, JHEP **1205**, 128 (2012), [arXiv:1203.4606].
- [27] M. Cacciari, G. P. Salam, and G. Soyez, JHEP **04**, 063 (2008), [arXiv:0802.1189].
- [28] A. Banfi and M. Dasgupta, Phys. Lett. B **628**, 49 (2005), [hep-ph/0508159].
- [29] Y. Delenda, R. Appleby, M. Dasgupta, and A. Banfi, JHEP **0612**, 044 (2006), [hep-ph/0610242].
- [30] K. Khelifa-Kerfa, JHEP **1202**, 072 (2012), [arXiv:1111.2016].
- [31] R. Kelley, J. R. Walsh, and S. Zuberi, JHEP **1209**, 117 (2012), [arXiv:1202.2361].
- [32] M. Dasgupta and G. P. Salam, Phys. Lett. B **512**, 323 (2001), [hep-ph/0104277].
- [33] M. Dasgupta and G. P. Salam, JHEP **0208**, 032 (2002), [hep-ph/0208073].
- [34] L. Clavelli, Phys. Lett. B **85**, 111 (1979).
- [35] T. Chandramohan and L. Clavelli, Nucl. Phys. B **184**, 365 (1981).
- [36] L. Clavelli and D. Wyler, Phys. Lett. B **103**, 383 (1981).
- [37] S. Fleming, A. H. Hoang, S. Mantry, and I. W. Stewart, Phys. Rev. D **77**, 114003 (2008), [arXiv:0711.2079].
- [38] D. Appell, G. Sterman, and P. B. Mackenzie, Nucl. Phys. B **309**, 259 (1988).
- [39] S. Catani, M. L. Mangano, and P. Nason, JHEP **07**, 024 (1998), [hep-ph/9806484].
- [40] T. Becher, M. Neubert, and G. Xu, JHEP **07**, 030 (2008), [arXiv:0710.0680].
- [41] R. Kelley, M. D. Schwartz, R. M. Schabinger, and H. X. Zhu, Phys. Rev. D **84**, 045022 (2011), [arXiv:1105.3676].
- [42] A. Hornig, C. Lee, I. W. Stewart, J. R. Walsh, and S. Zuberi, JHEP **1108**, 054 (2011), [arXiv:1105.4628].
- [43] B. Mellado, W. Quayle, and S. L. Wu, Phys. Lett. B **611**, 60 (2005), [hep-ph/0406095].
- [44] B. Mellado, W. Quayle, and S. L. Wu, Phys. Rev. D **76**, 093007 (2007), [arXiv:0708.2507].

- [45] X. Liu and F. Petriello, *Phys. Rev. D* **87**, 014018 (2013), [arXiv:1210.1906].
- [46] I. W. Stewart, F. J. Tackmann, and W. J. Waalewijn, *Phys. Rev. Lett.* **105**, 092002 (2010), [arXiv:1004.2489].
- [47] T. T. Jouttenus, I. W. Stewart, F. J. Tackmann, and W. J. Waalewijn, *Phys. Rev. D* **83**, 114030 (2011), [arXiv:1102.4344].
- [48] C. W. Bauer, S. Fleming, and M. E. Luke, *Phys. Rev. D* **63**, 014006 (2000), [hep-ph/0005275].
- [49] C. W. Bauer, S. Fleming, D. Pirjol, and I. W. Stewart, *Phys. Rev. D* **63**, 114020 (2001), [hep-ph/0011336].
- [50] C. W. Bauer and I. W. Stewart, *Phys. Lett. B* **516**, 134 (2001), [hep-ph/0107001].
- [51] T. Sjöstrand, S. Mrenna, and P. Skands, *JHEP* **05**, 026 (2006), [hep-ph/0603175].
- [52] T. Sjöstrand, S. Mrenna, and P. Skands, *Comput. Phys. Commun.* **178**, 852 (2008), [arXiv:0710.3820].
- [53] M. Cacciari and G. P. Salam, *Phys. Lett. B* **659**, 119 (2008), [arXiv:0707.1378].
- [54] G. Soyez, G. P. Salam, J. Kim, S. Dutta, and M. Cacciari, arXiv:1211.2811.
- [55] J. Thaler and K. Van Tilburg, *JHEP* **1202**, 093 (2012), [arXiv:1108.2701].
- [56] C. R. Schmidt, *Phys. Lett. B* **413**, 391 (1997), [hep-ph/9707448].
- [57] I. W. Stewart, F. J. Tackmann, and W. J. Waalewijn, arXiv:1211.2305.
- [58] C. W. Bauer and A. V. Manohar, *Phys. Rev. D* **70**, 034024 (2004), [hep-ph/0312109].
- [59] S. Fleming, A. K. Leibovich, and T. Mehen, *Phys. Rev. D* **68**, 094011 (2003), [hep-ph/0306139].
- [60] T. Becher and M. D. Schwartz, *JHEP* **02**, 040 (2010), [arXiv:0911.0681].
- [61] I. W. Stewart, F. J. Tackmann, and W. J. Waalewijn, *Phys. Rev. D* **81**, 094035 (2010), [arXiv:0910.0467].
- [62] I. W. Stewart, F. J. Tackmann, and W. J. Waalewijn, *JHEP* **09**, 005 (2010), [arXiv:1002.2213].
- [63] S. Mantry and F. Petriello, *Phys. Rev. D* **81**, 093007 (2010), [arXiv:0911.4135].
- [64] C. F. Berger, C. Marcantonini, I. W. Stewart, F. J. Tackmann, and W. J. Waalewijn, *JHEP* **04**, 092 (2011), [arXiv:1012.4480].
- [65] G. P. Korchemsky and A. V. Radyushkin, *Nucl. Phys. B* **283**, 342 (1987).
- [66] S. Moch, J. A. M. Vermaseren, and A. Vogt, *Nucl. Phys. B* **688**, 101 (2004), [hep-ph/0403192].
- [67] G. Kramer and B. Lampe, *Z. Phys. C* **34**, 497 (1987), [Erratum-ibid. C **42**, 504 (1989)].
- [68] R. V. Harlander, *Phys. Lett. B* **492**, 74 (2000), [hep-ph/0007289].
- [69] S. Mert Aybat, L. J. Dixon, and G. Sterman, *Phys. Rev. D* **74**, 074004 (2006), [hep-ph/0607309].
- [70] T. Becher and M. Neubert, *Phys. Lett. B* **637**, 251 (2006), [hep-ph/0603140].
- [71] J.-y. Chiu, R. Kelley, and A. V. Manohar, *Phys. Rev. D* **78**, 073006 (2008), [arXiv:0806.1240].
- [72] T. Becher and M. Neubert, *JHEP* **06**, 081 (2009), [arXiv:0903.1126].
- [73] Z. Ligeti, I. W. Stewart, and F. J. Tackmann, *Phys. Rev. D* **78**, 114014 (2008), [arXiv:0807.1926].
- [74] ATLAS Collaboration Collaboration, G. Aad *et al.*, *Phys. Lett. B* **716**, 1 (2012), [arXiv:1207.7214].
- [75] CMS Collaboration Collaboration, S. Chatrchyan *et al.*, *Phys. Lett. B* **716**, 30 (2012), [arXiv:1207.7235].
- [76] A. D. Martin, W. J. Stirling, R. S. Thorne, and G. Watt, *Eur. Phys. J. C* **64**, 653 (2009), [arXiv:0905.3531].
- [77] M. Cacciari, G. P. Salam, and G. Soyez, *Eur. Phys. J. C* **72**, 1896 (2012), [arXiv:1111.6097].
- [78] Y. L. Dokshitzer, G. Marchesini, and B. R. Webber, *Nucl. Phys. B* **469**, 93 (1996), [hep-ph/9512336].
- [79] Y. L. Dokshitzer and B. R. Webber, *Phys. Lett. B* **352**, 451 (1995), [hep-ph/9504219].
- [80] Y. L. Dokshitzer, A. Lucenti, G. Marchesini, and G. Salam, *Nucl. Phys. B* **511**, 396 (1998), [hep-ph/9707532].
- [81] G. P. Salam and D. Wicke, *JHEP* **05**, 061 (2001), [hep-ph/0102343].
- [82] C. Lee and G. F. Sterman, *eConf C0601121*, A001 (2006), [hep-ph/0603066].
- [83] A. H. Hoang and I. W. Stewart, *Phys. Lett. B* **660**, 483 (2008), [arXiv:0709.3519].
- [84] V. Mateu, I. W. Stewart, and J. Thaler, *Phys. Rev. D* **87**, 014025 (2013), [arXiv:1209.3781].
- [85] M. Dasgupta, L. Magnea, and G. P. Salam, *JHEP* **02**, 055 (2008), [arXiv:0712.3014].
- [86] S. Fleming, A. K. Leibovich, and T. Mehen, *Phys. Rev. D* **74**, 114004 (2006), [hep-ph/0607121].
- [87] C. Balzereit, T. Mannel, and W. Kilian, *Phys. Rev. D* **58**, 114029 (1998), [hep-ph/9805297].
- [88] M. Neubert, *Eur. Phys. J. C* **40**, 165 (2005), [hep-ph/0408179].
- [89] O. V. Tarasov, A. A. Vladimirov, and A. Y. Zharkov, *Phys. Lett. B* **93**, 429 (1980).
- [90] S. A. Larin and J. A. M. Vermaseren, *Phys. Lett. B* **303**, 334 (1993), [hep-ph/9302208].
- [91] A. Idilbi, X. dong Ji, and F. Yuan, *Nucl. Phys. B* **753**, 42 (2006), [hep-ph/0605068].
- [92] T. Becher, M. Neubert, and B. D. Pecjak, *JHEP* **01**, 076 (2007), [hep-ph/0607228].
- [93] I. W. Stewart and F. J. Tackmann, *Phys. Rev. D* **85**, 034011 (2012), [arXiv:1107.2117].
- [94] S. Dittmaier *et al.*, arXiv:1201.3084.